

Array Methods: DLR 1

BeBeC 2016

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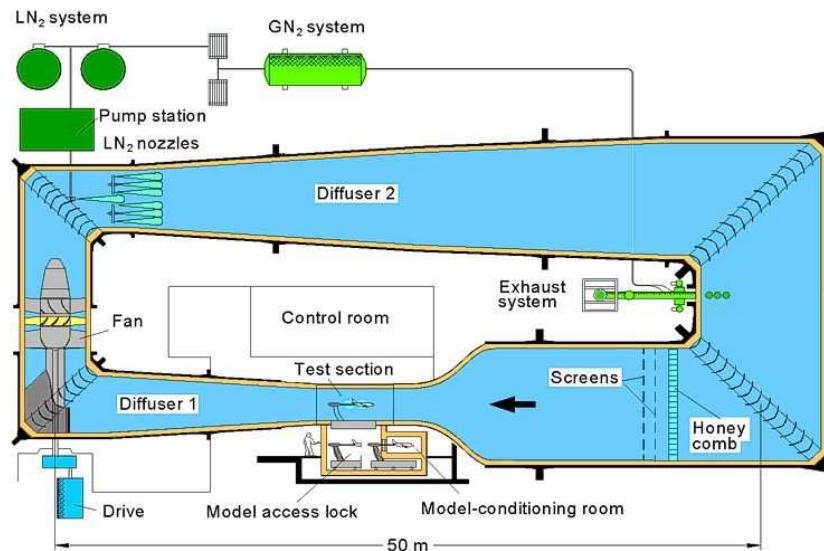
Overview

- Measurement Setup
- Model Details and Potential Sources
- Comparison of Contributions
- Conclusion



Measurement Setup: Wind Tunnel

- Cryogenic wind tunnel located at the DLR Cologne Site (DNW-KKK)



- Göttingen type wind tunnel
- 2.4 m x 2.4 m closed test section
- 300 K > T > 100 K ; 0.1 < Ma < 0.38



Measurement Setup: Array and Model

Microphone array

- 144 microphones (135 used)
- arranged in spiral arms

DO-728 half model

- scale: 1 : 9.24
- $\frac{1}{2}$ - spanwidth: 1.44 m
- chord length: 0.338 m
- no transition fixation on slat

Measurement parameter

- angle of attack: 3° , 5° , 9°
- Mach number: 0.125, 0.2, 0.25
- temperature: ~ 290 K



DO-728 half-model at landing configuration



Model Details and Potential Noise Sources

flap

flap track

flap side edge/flap tip

leading edge slat

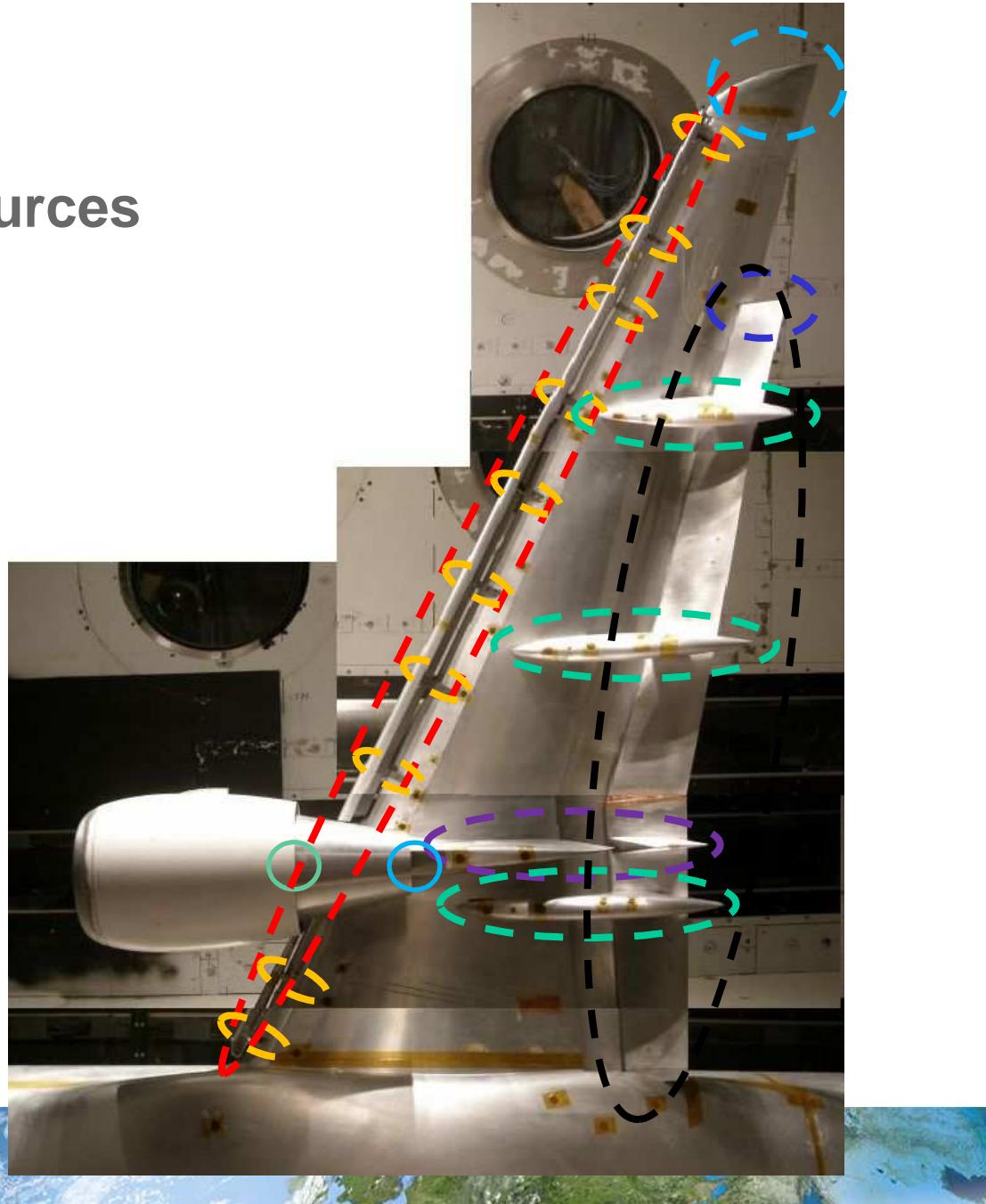
slat track

wing tip

core nozzle

fan nozzle

engine mount



Available Data at TU-Cottbus Server

1) Data

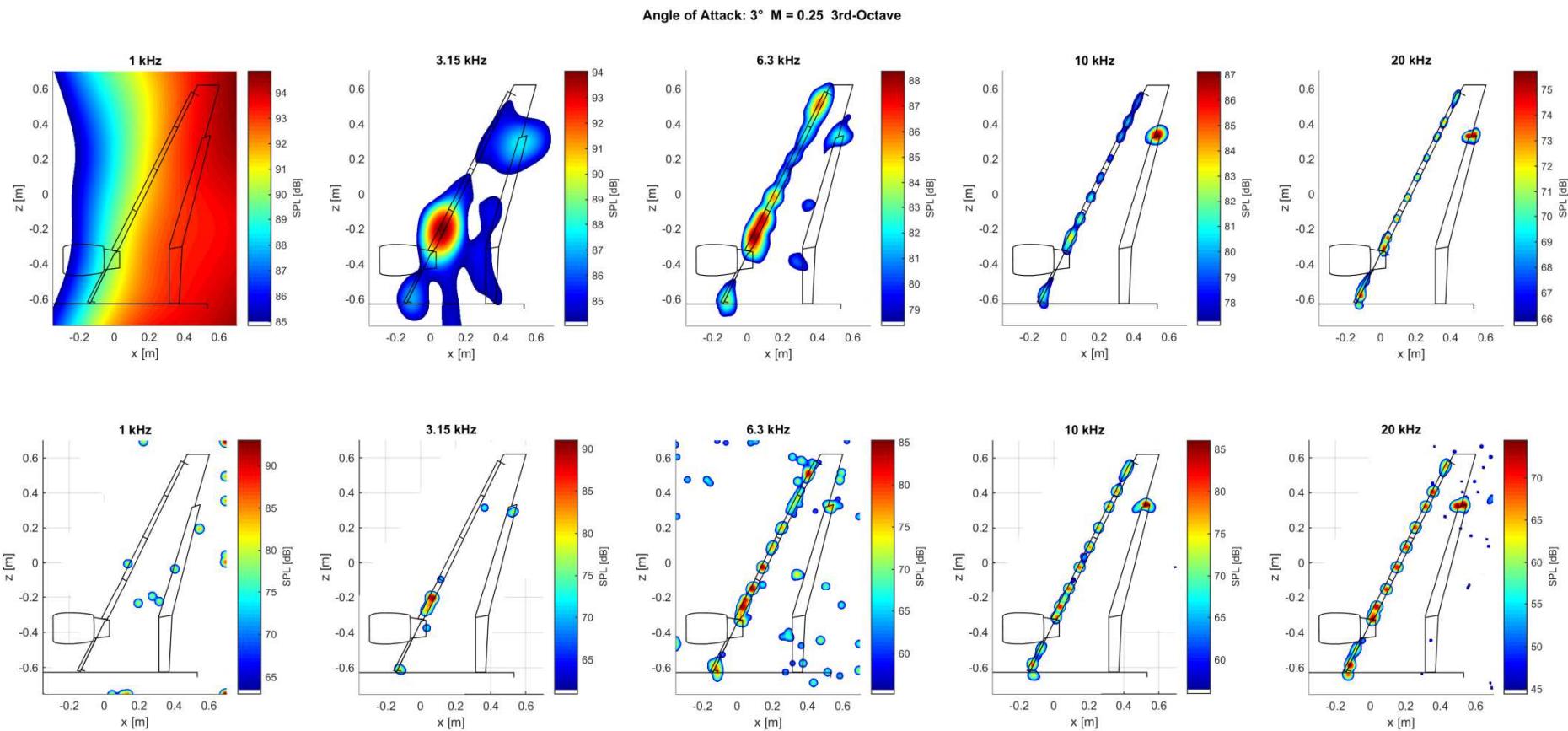
- Time Data (*_TimeSeriesEssential.h5)
- CSM Data (*._CsmEssential.h5)

2) Results

- Standard Dirty Maps (*_CsmOptional_dx_0.01_DirtyMap.h5)
- CleanSC Maps (*_CsmOptional_dx_0.01_CleanSC.h5)
- Resolution
 $\Delta x = 0.01m (\sim 15k Pts.), 0.02m (\sim 4k Pts.), 0.05m (\sim 700 Pts)$



Overview of Sources



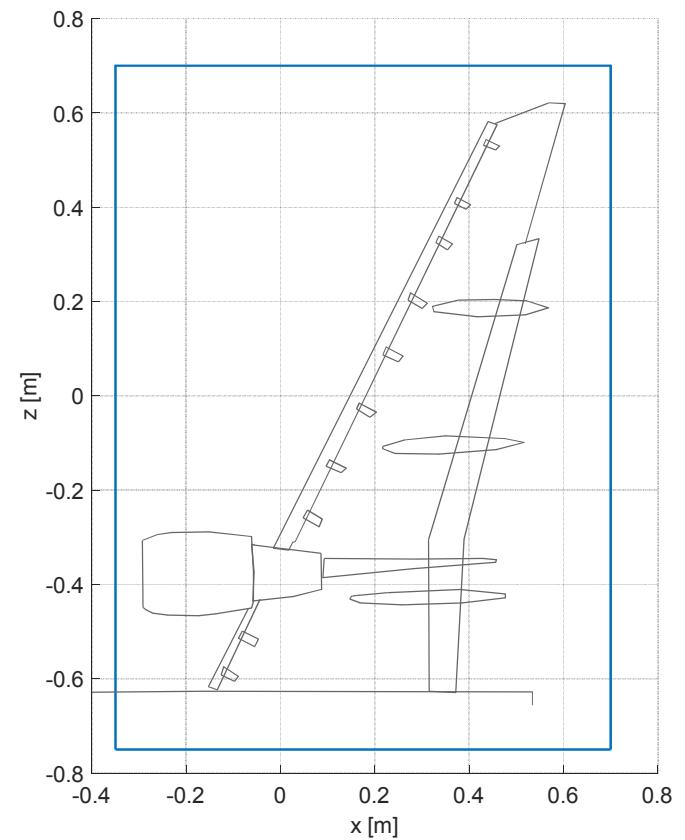
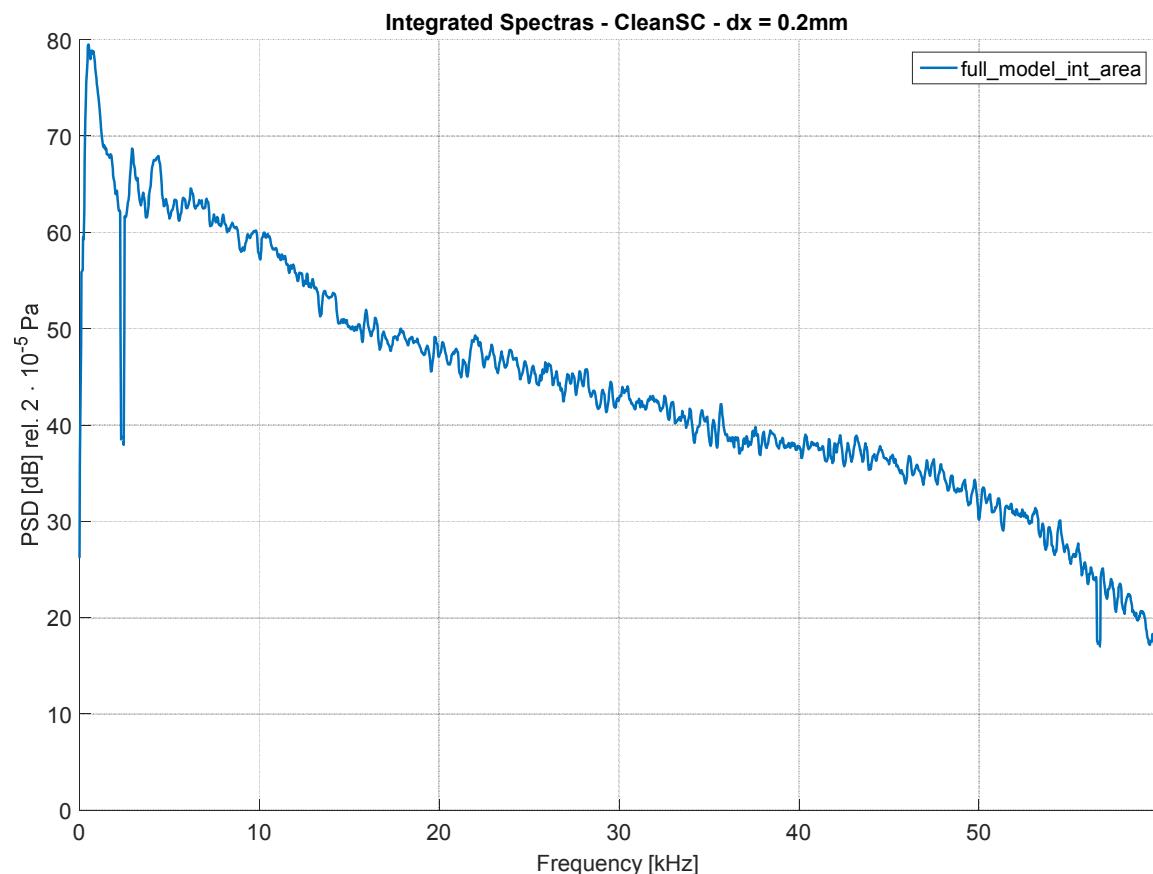
Upper row: Frequency-domain beamforming (Dynamic: 10 dB)

Lower row: CleanSC (Dynamic: 30 dB)

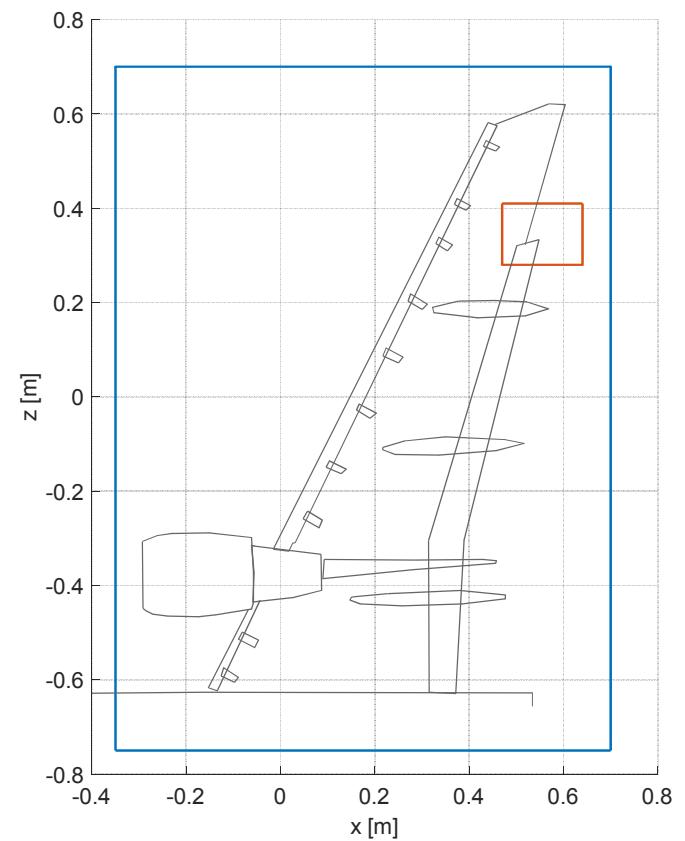
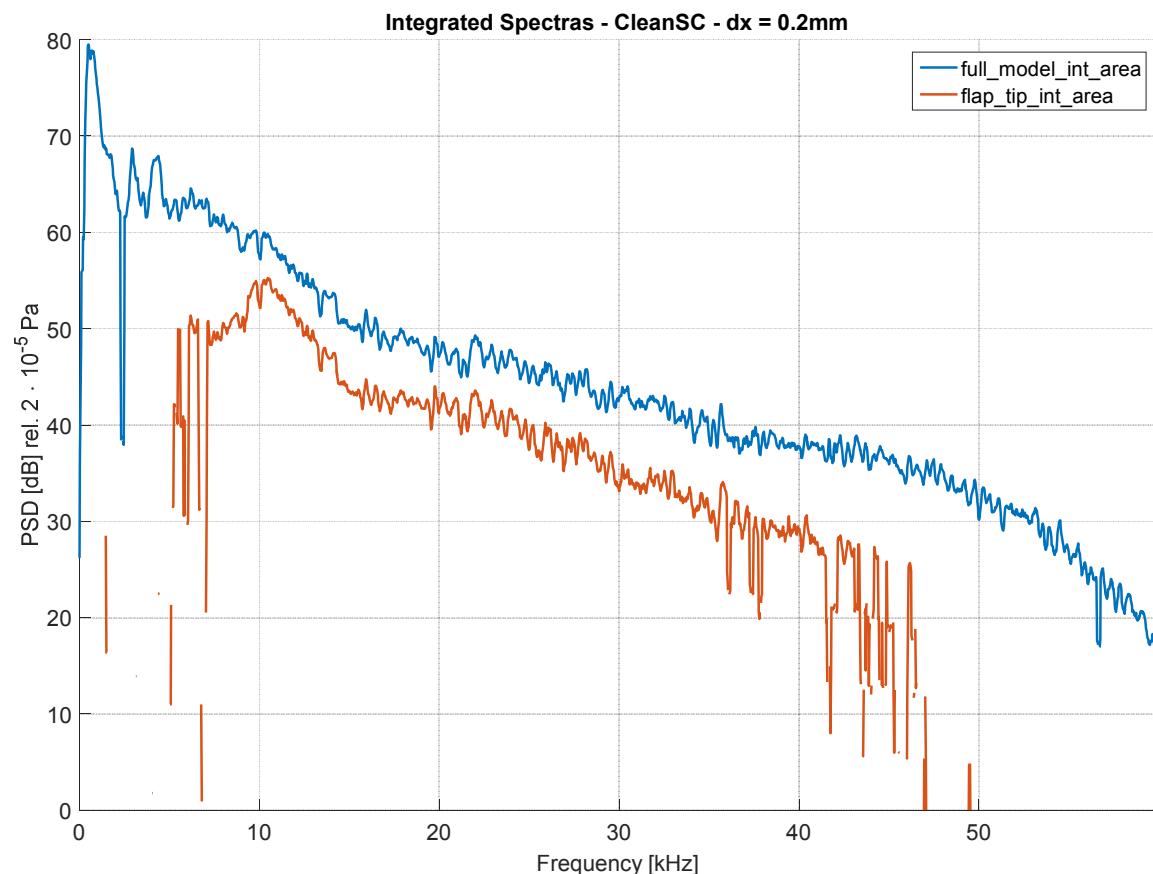
Resolution $\Delta x = 0.01m$



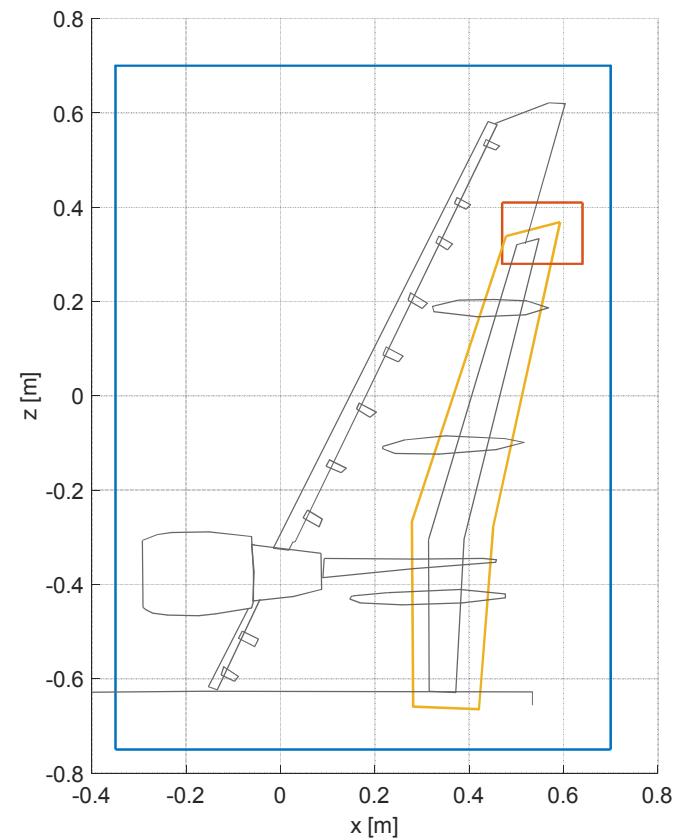
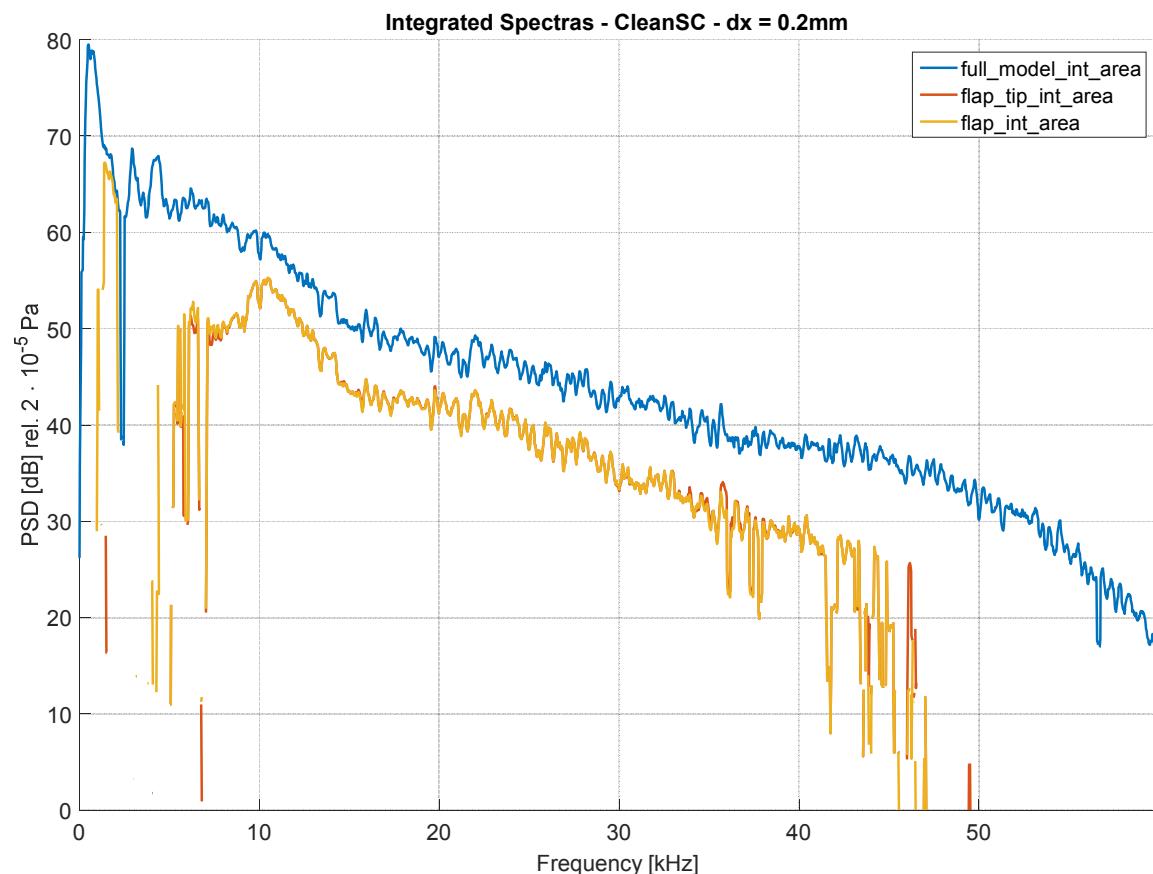
Integrated Spectra of Deconvoluted Source Maps



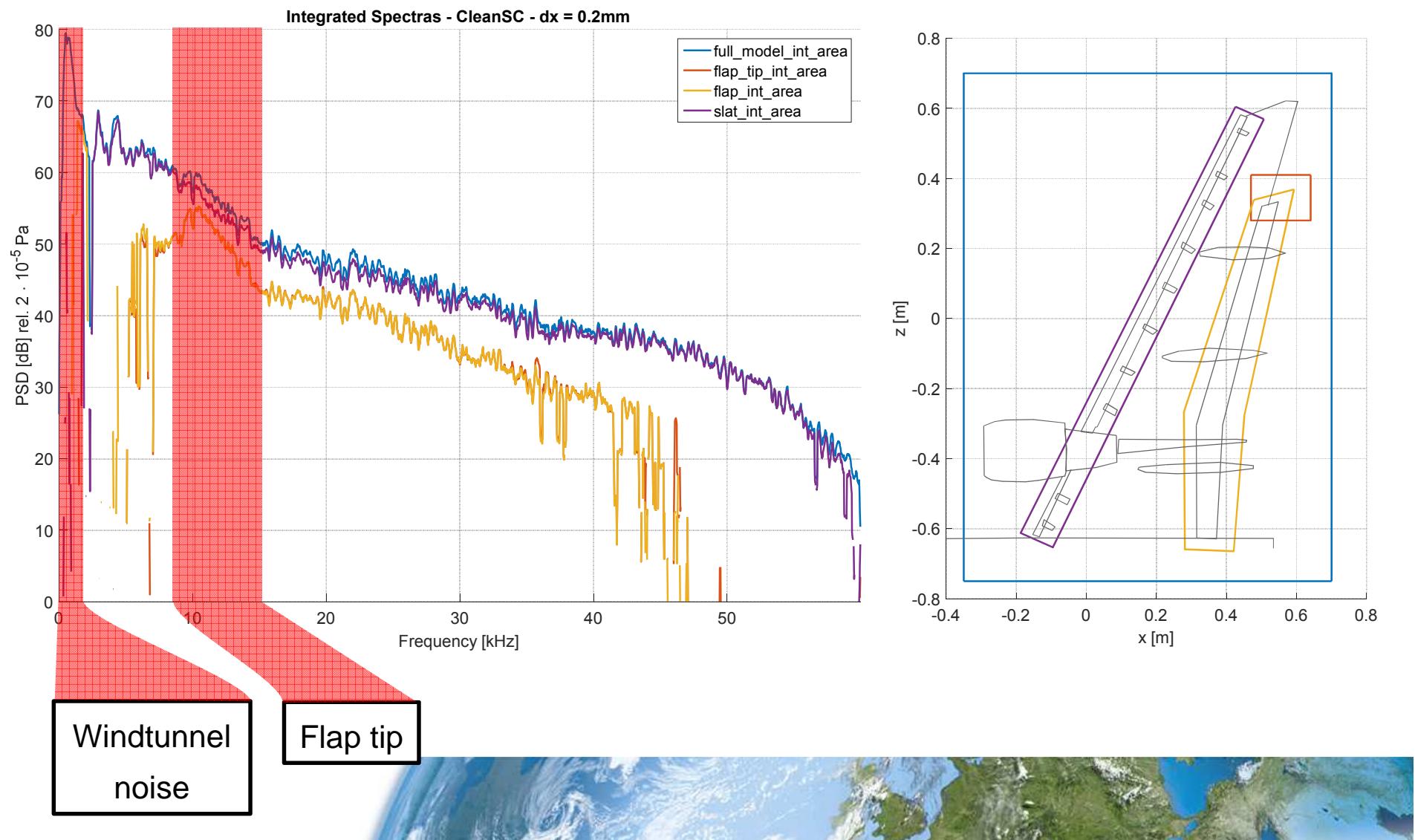
Integrated Spectra of Deconvoluted Source Maps



Integrated Spectra of Deconvoluted Source Maps



Integrated Spectra of Deconvoluted Source Maps



Beamforming Maps

low resolution (high computational effort due to deconvolution)

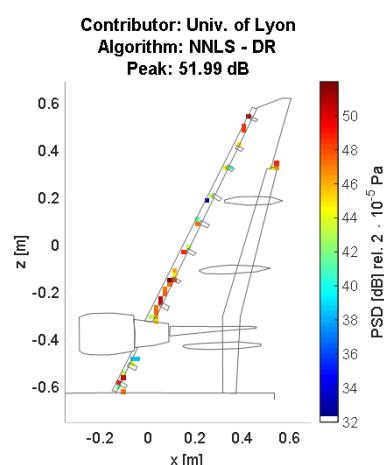
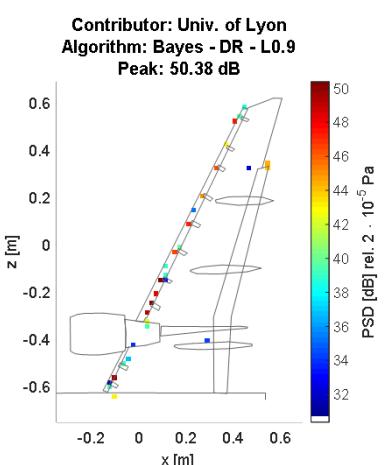
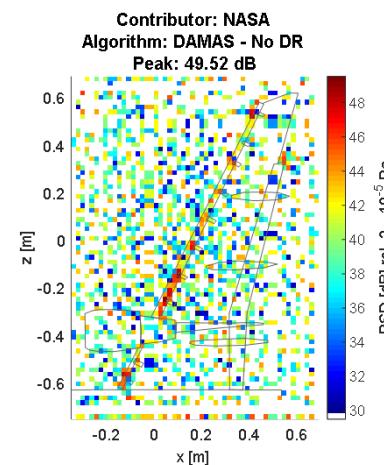
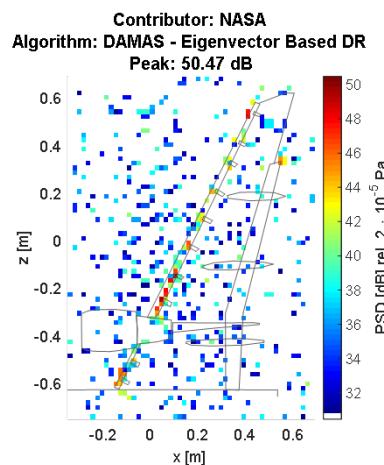
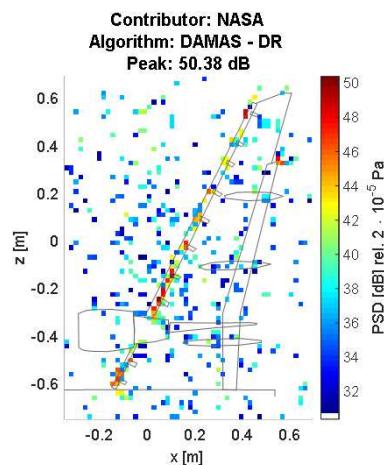
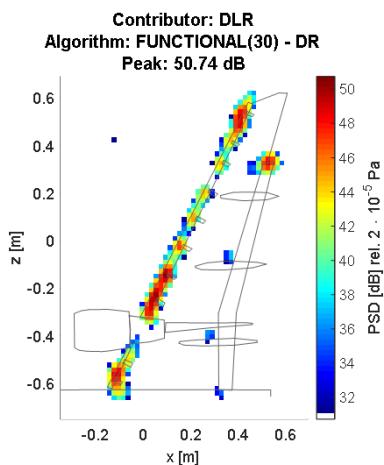
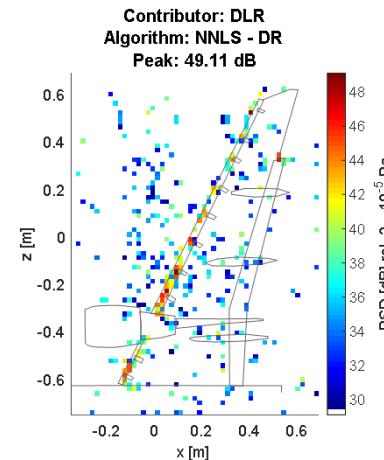
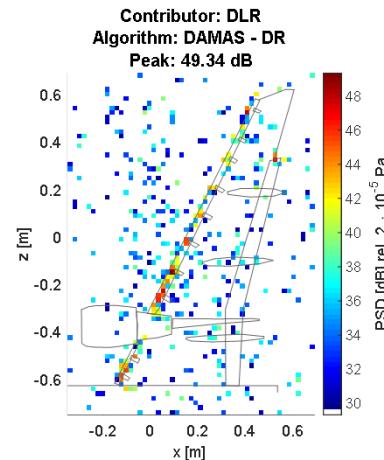
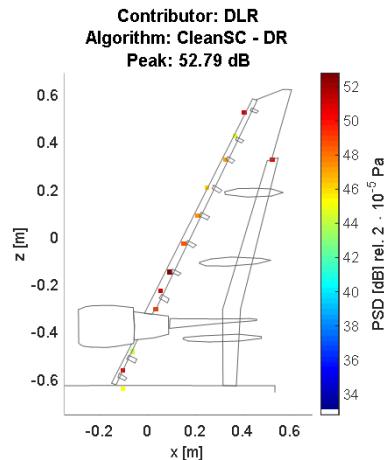
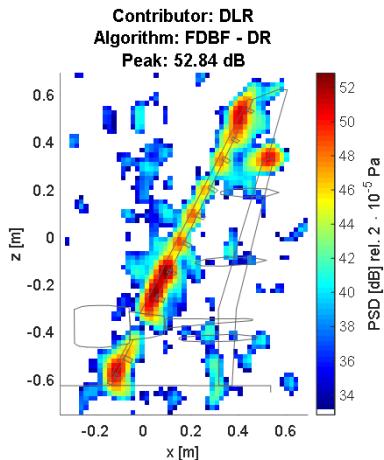
Contributors:

Chris Bahr, NASA
DAMAS and derivatives

Antonio Pereira, University of Lyon
NNLS, Bayes

Daniel Ernst, DLR
DAMAS, FDBF, Functional, CleanSC, NNLS

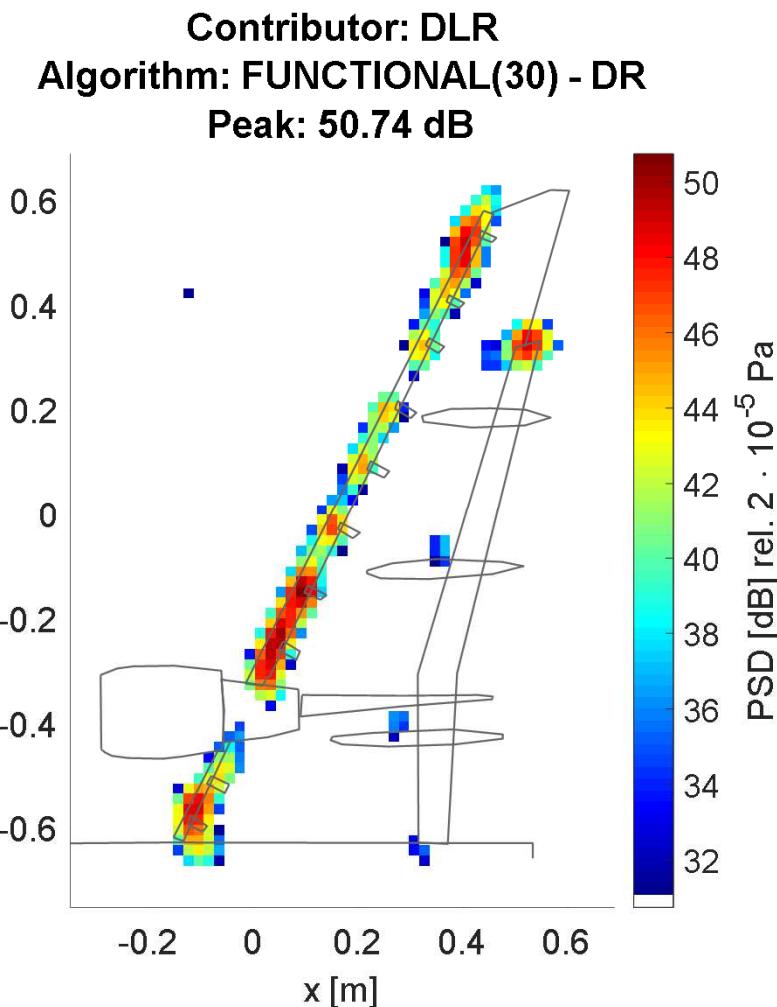
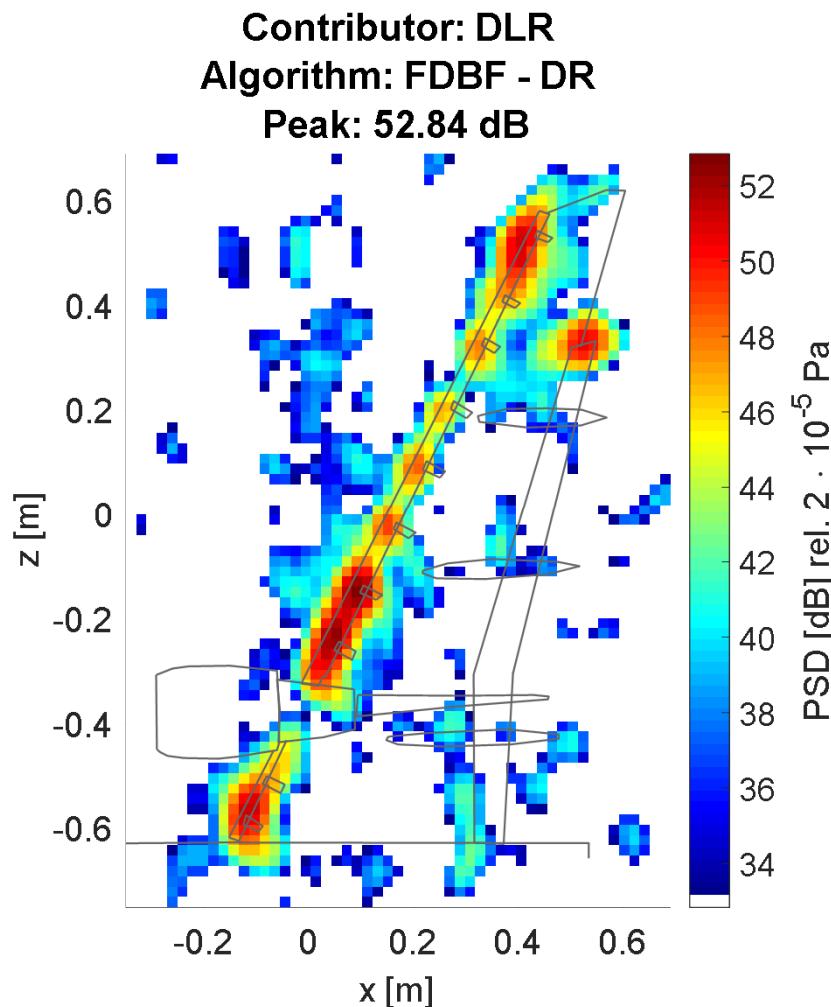


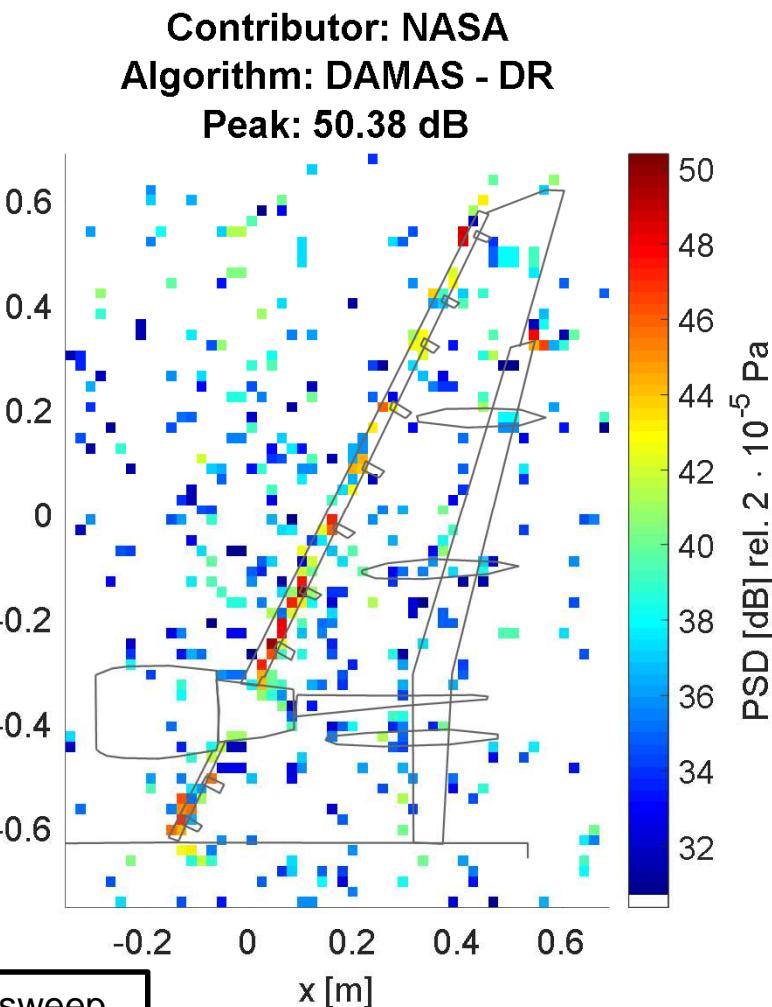
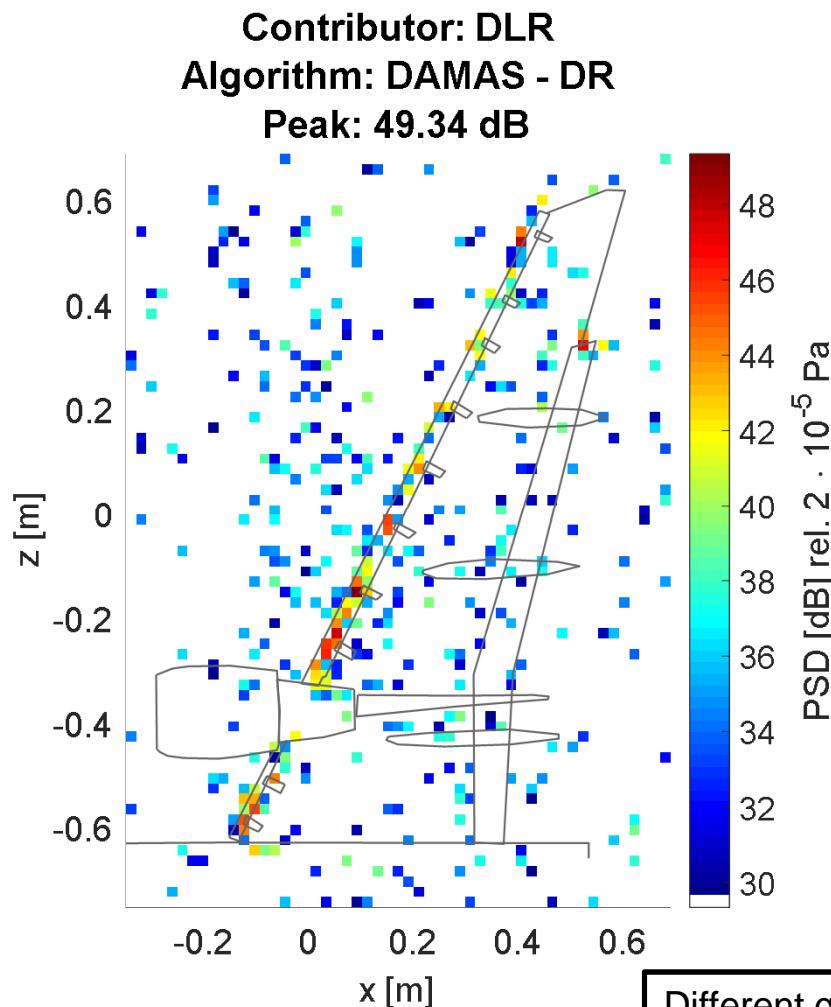


Narrow band source maps @ 8496 Hz
Dynamic range = 20 dB
Grid resolution = 20 mm



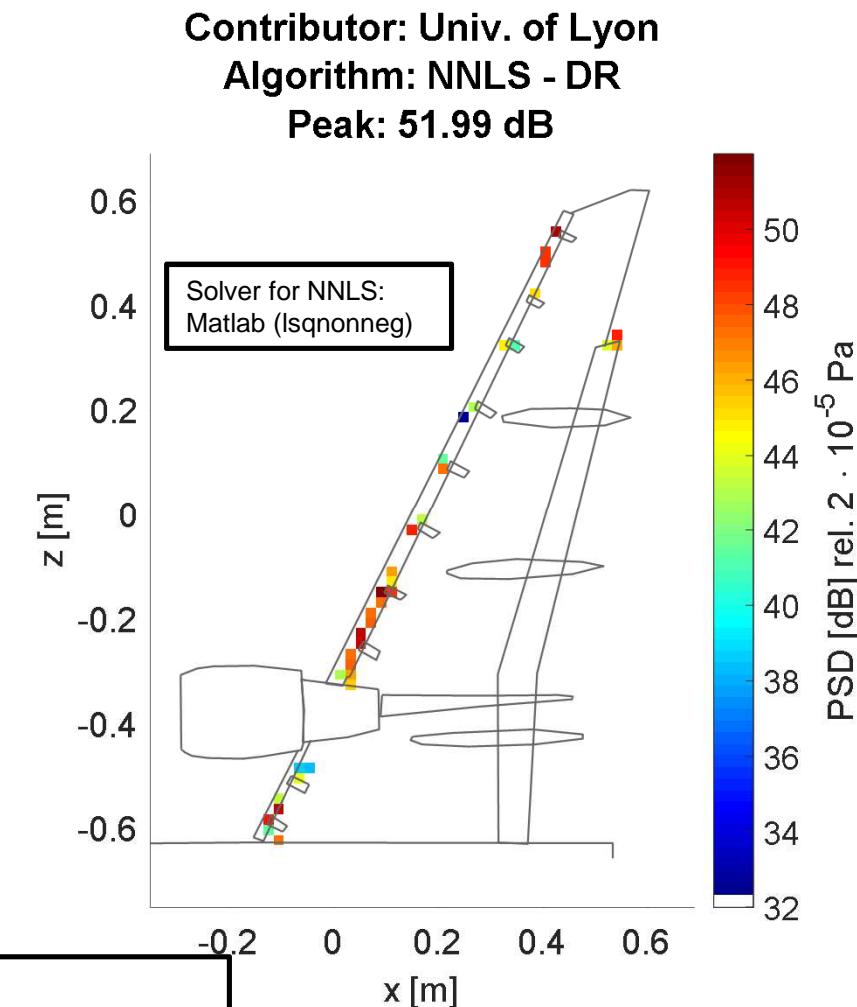
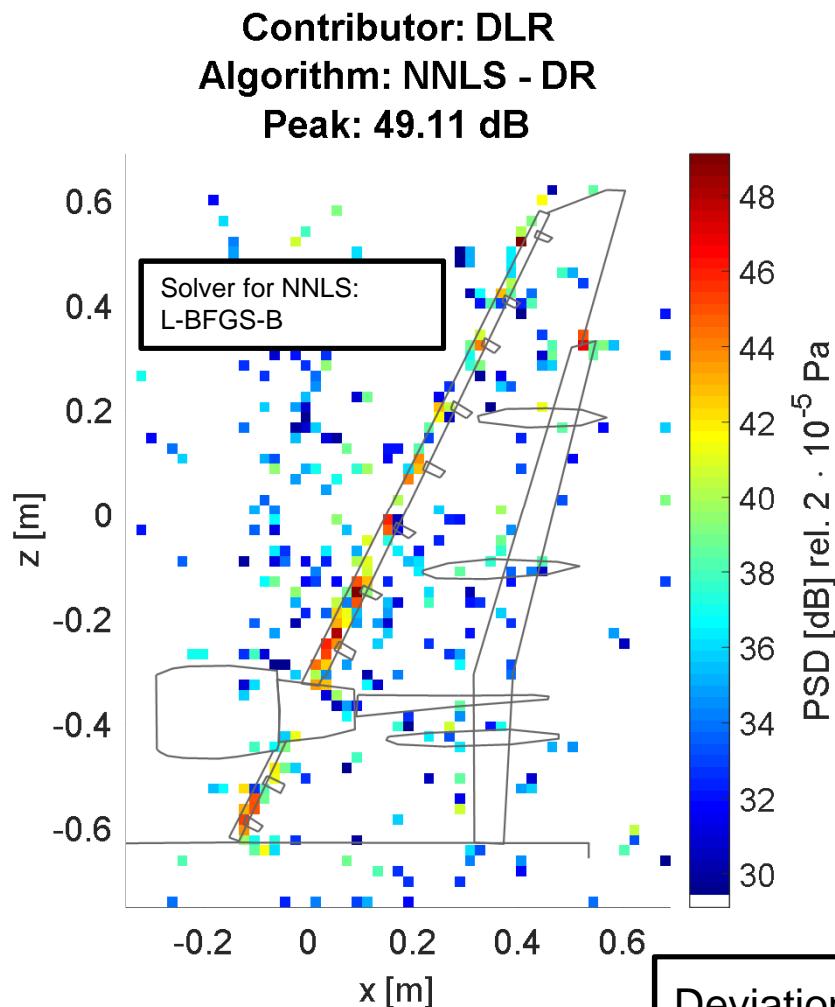
Narrow Band Source Maps @ 8496 Hz
Dynamic Range = 20 dB
Grid Resolution = 20 mm





Different grids and sweep
strategies where used.
Both 200 Iterations.

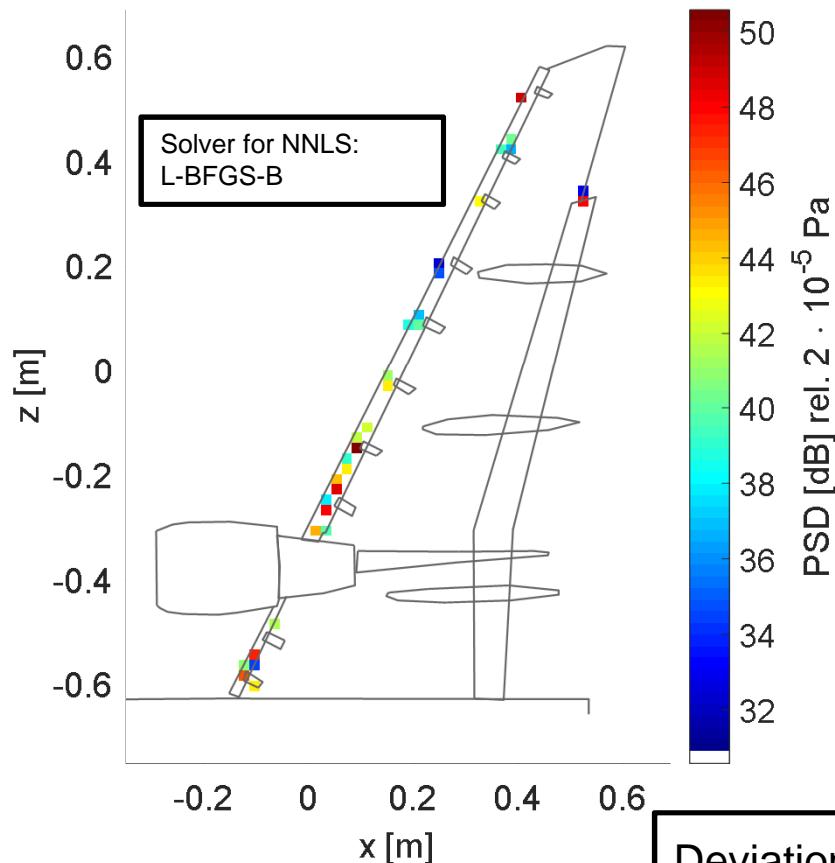




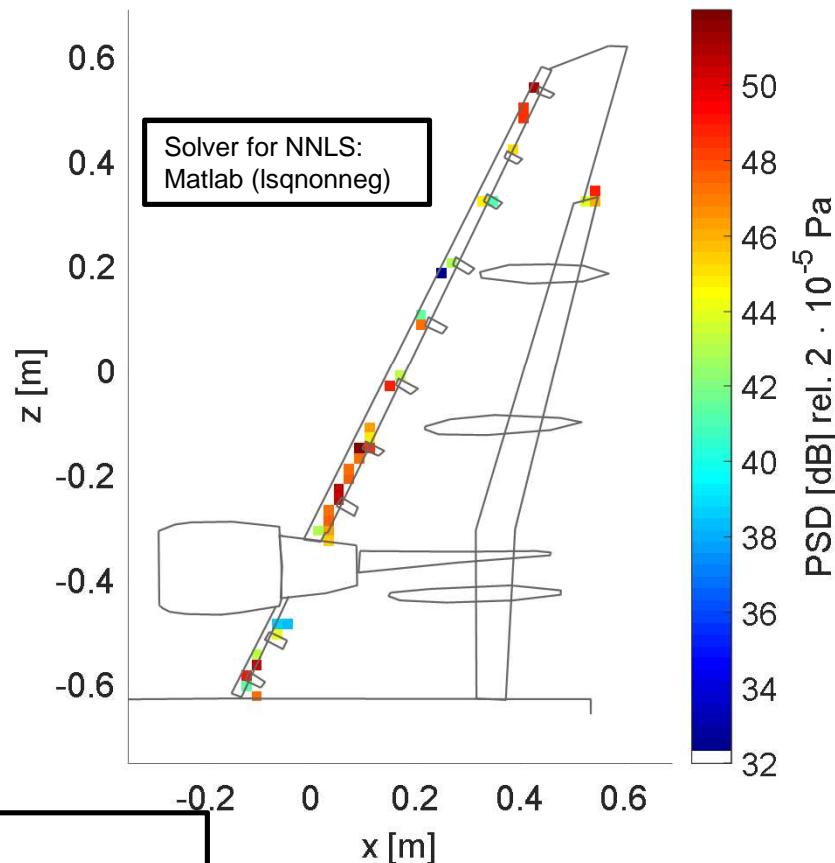
Deviations:
Maybe slightly
different approach to
compute the psf



Contributor: DLR
Algorithm: NNLS - DR PSF>0 NoPsfDR
Peak: 50.58 dB



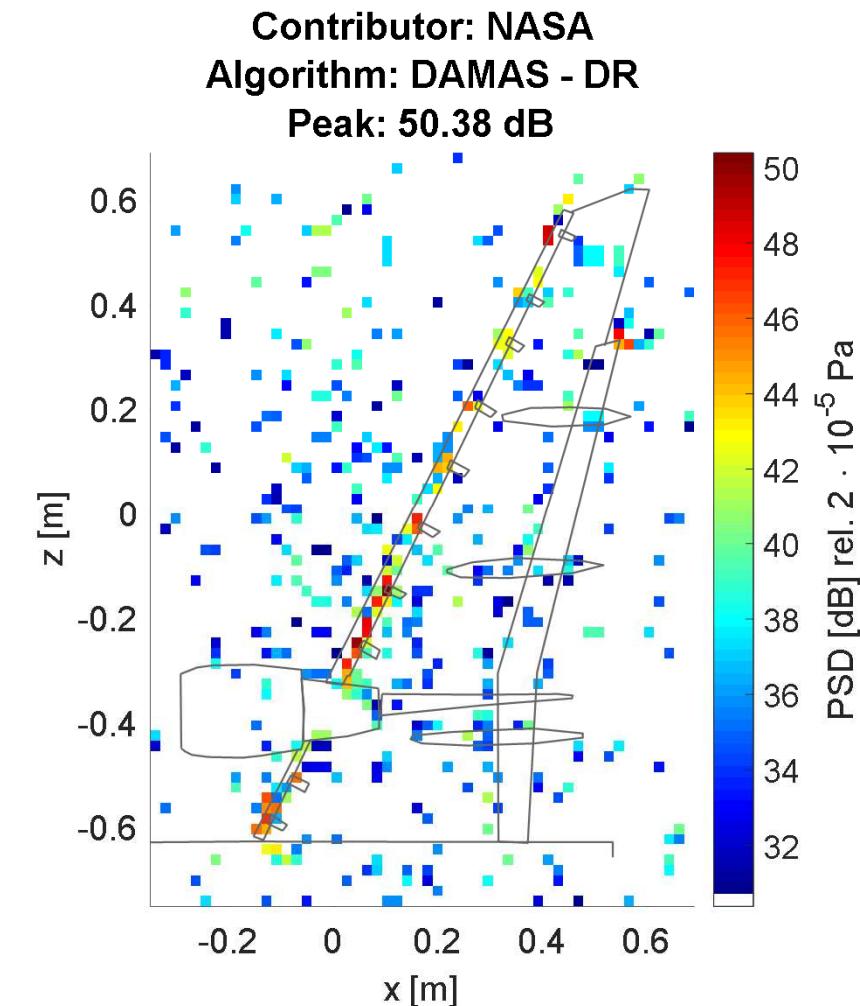
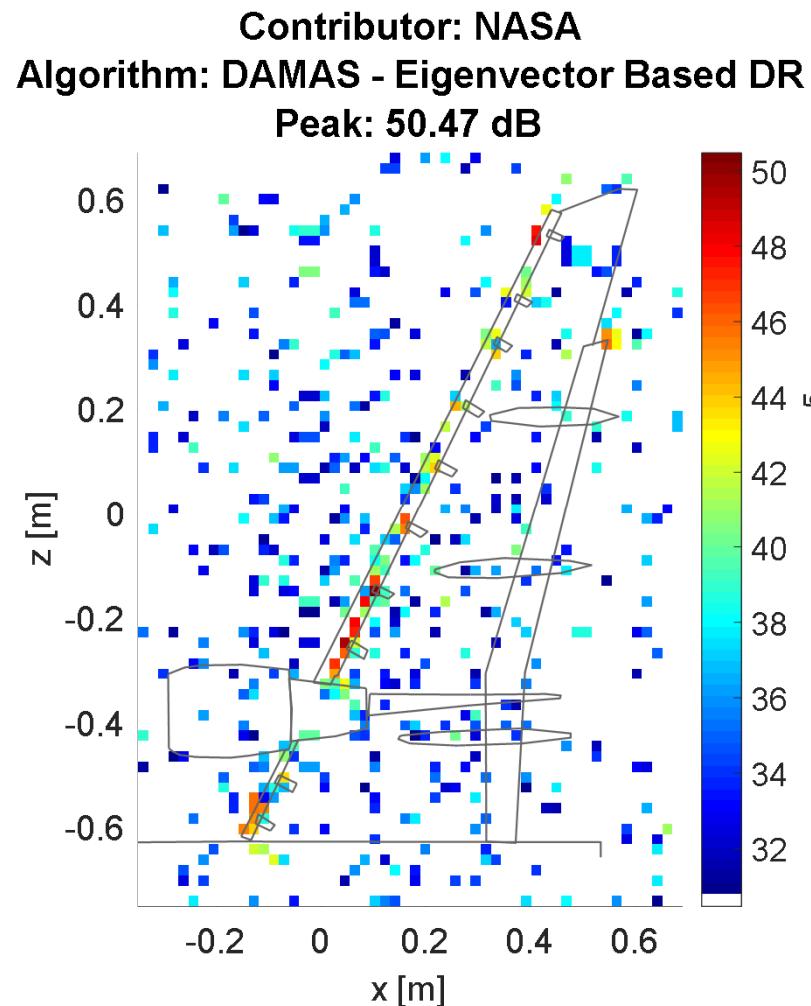
Contributor: Univ. of Lyon
Algorithm: NNLS - DR
Peak: 51.99 dB



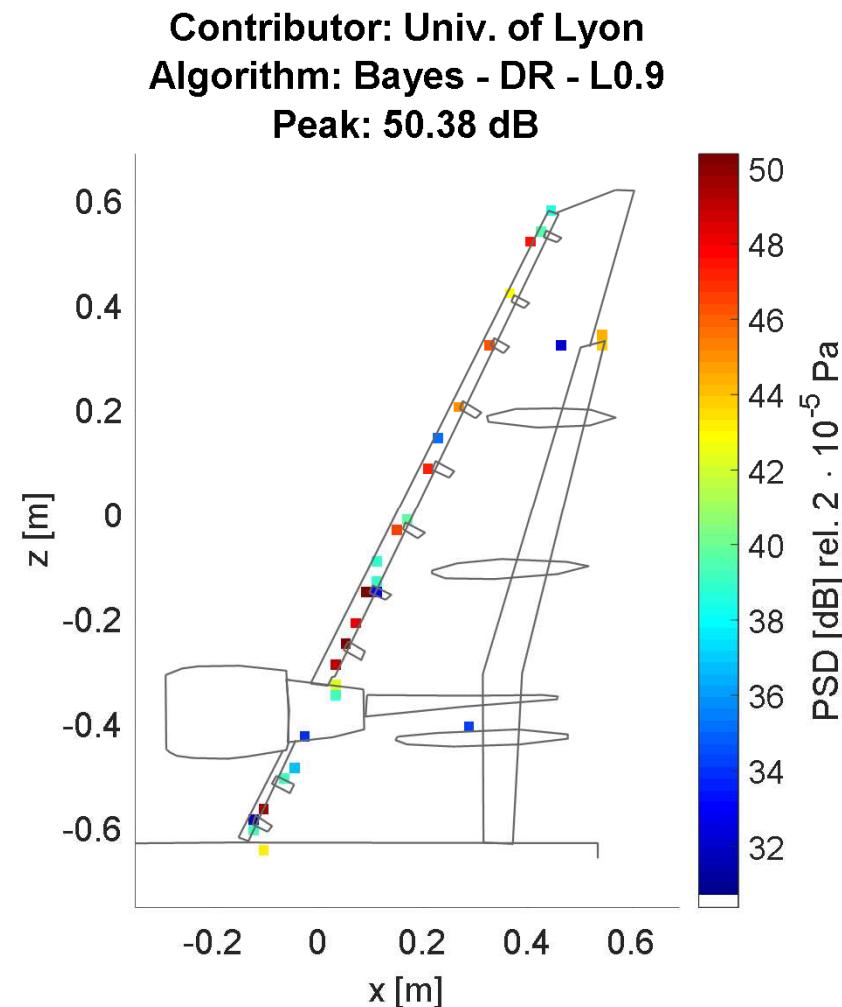
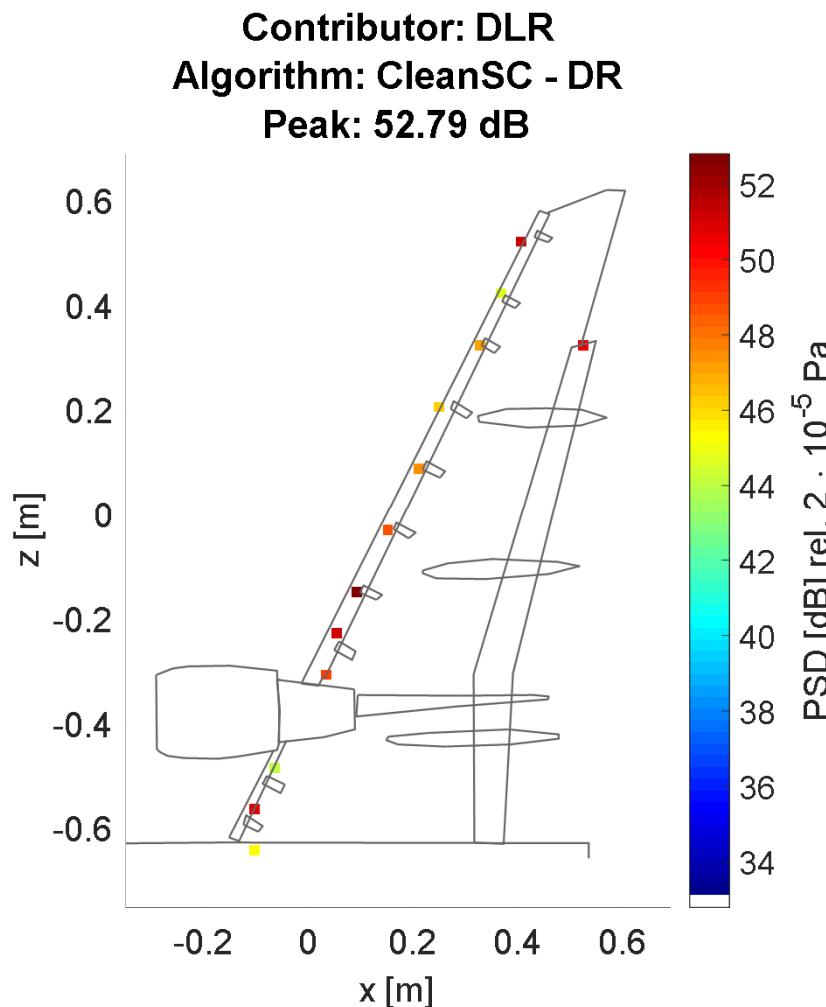
Deviations:
Maybe slightly
different approach to
compute the psf



Narrow Band Source Maps @ 8496 Hz
Dynamic Range = 20 dB
Grid Resolution = 20 mm

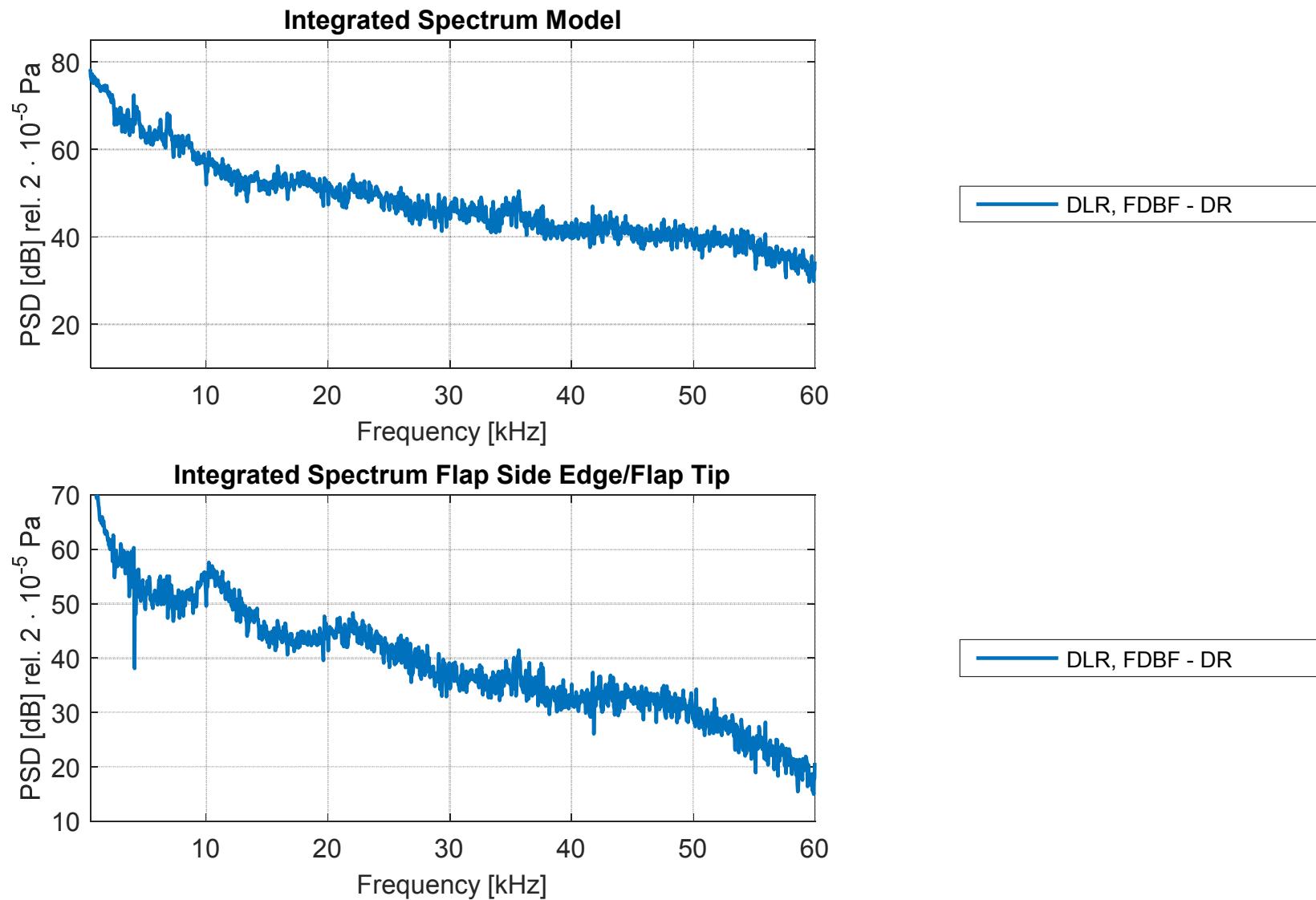


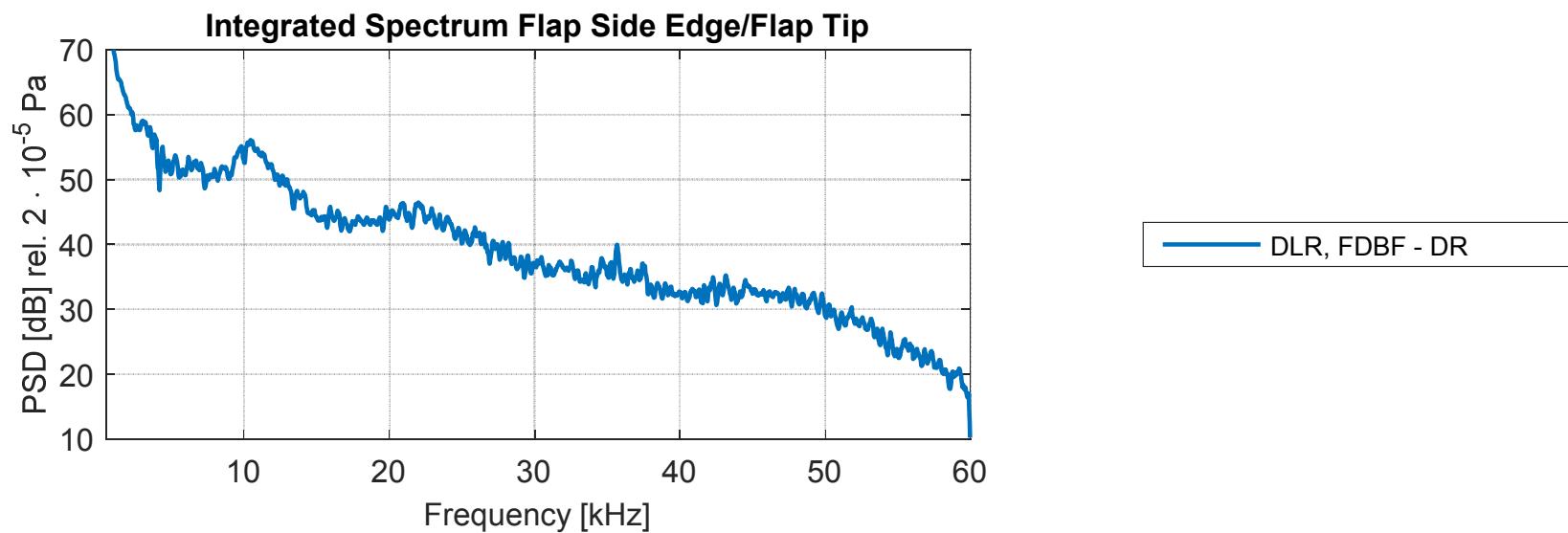
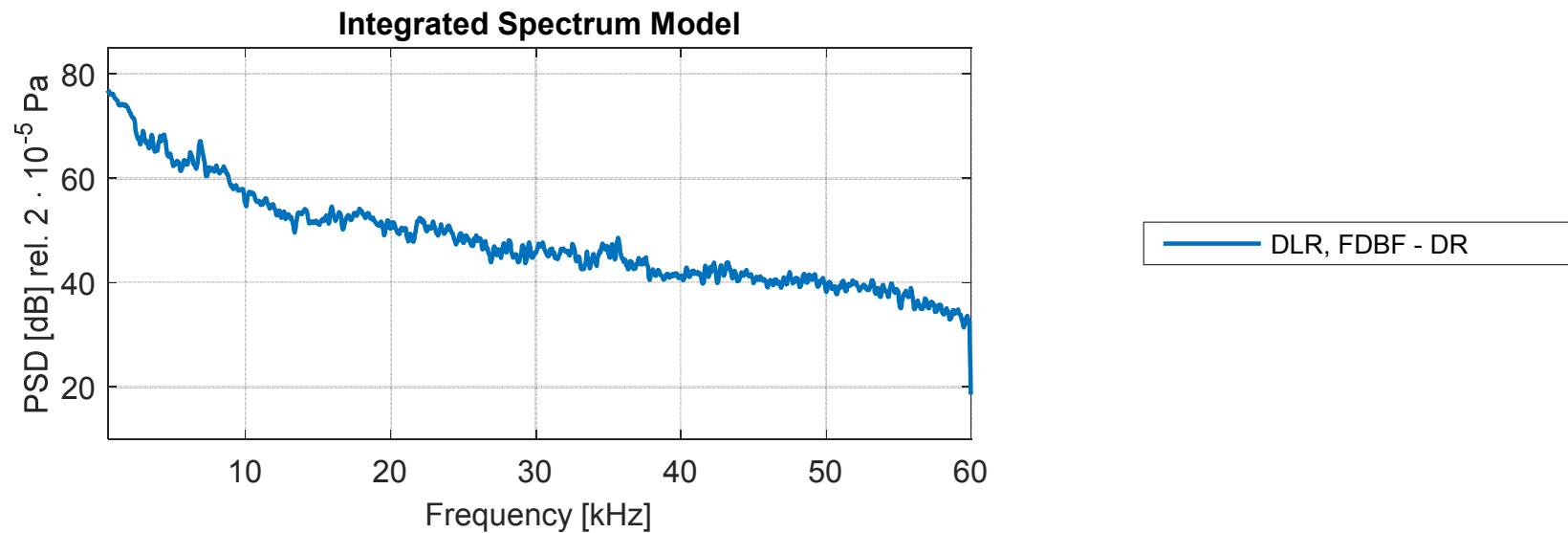
Narrow Band Source Maps @ 8496 Hz
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Grid Resolution = 20 mm



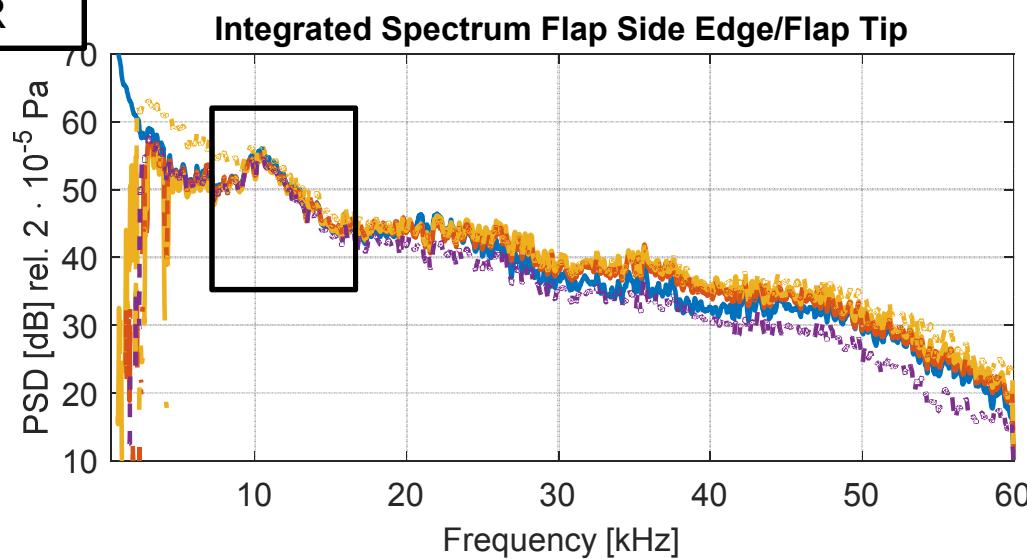
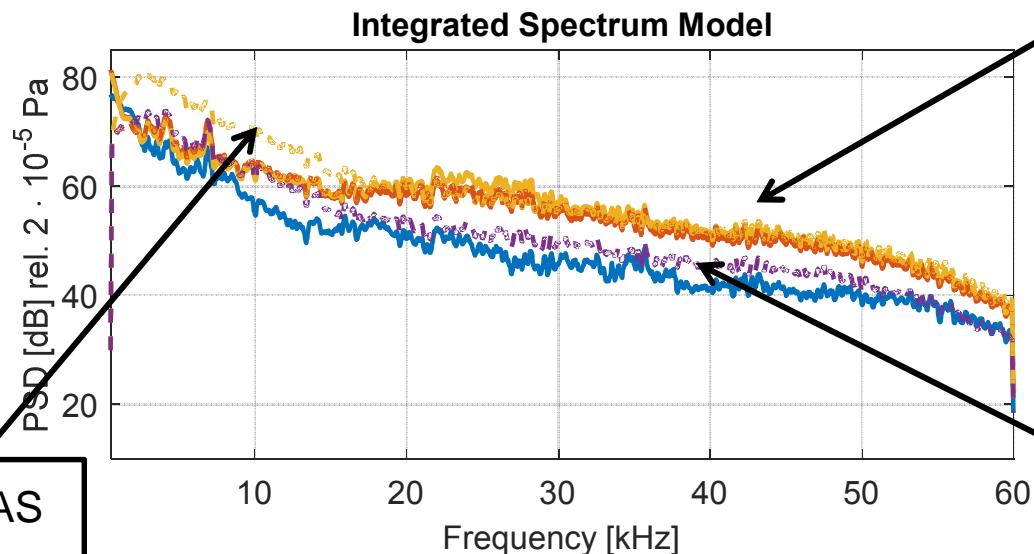
Integrated Spectra

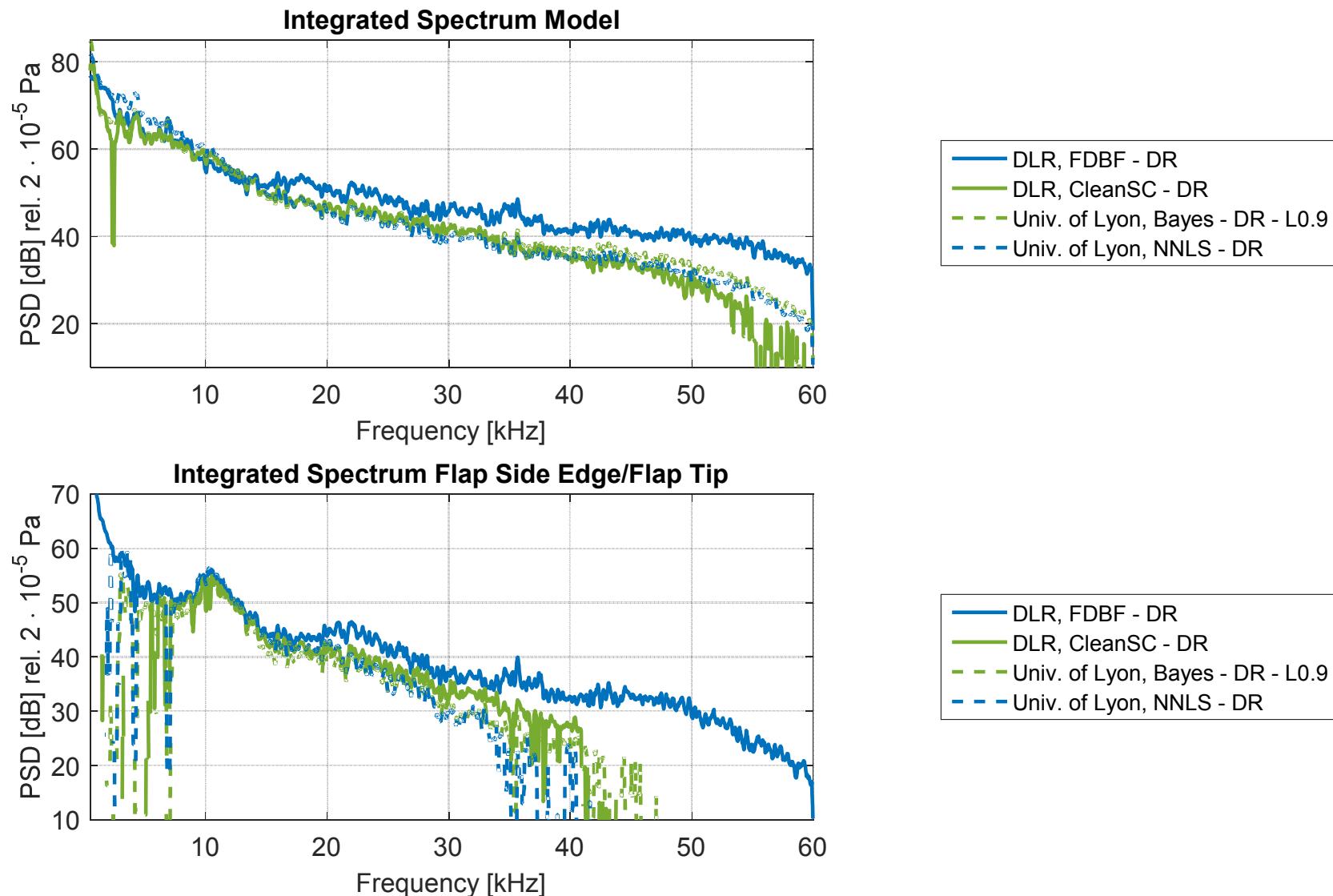


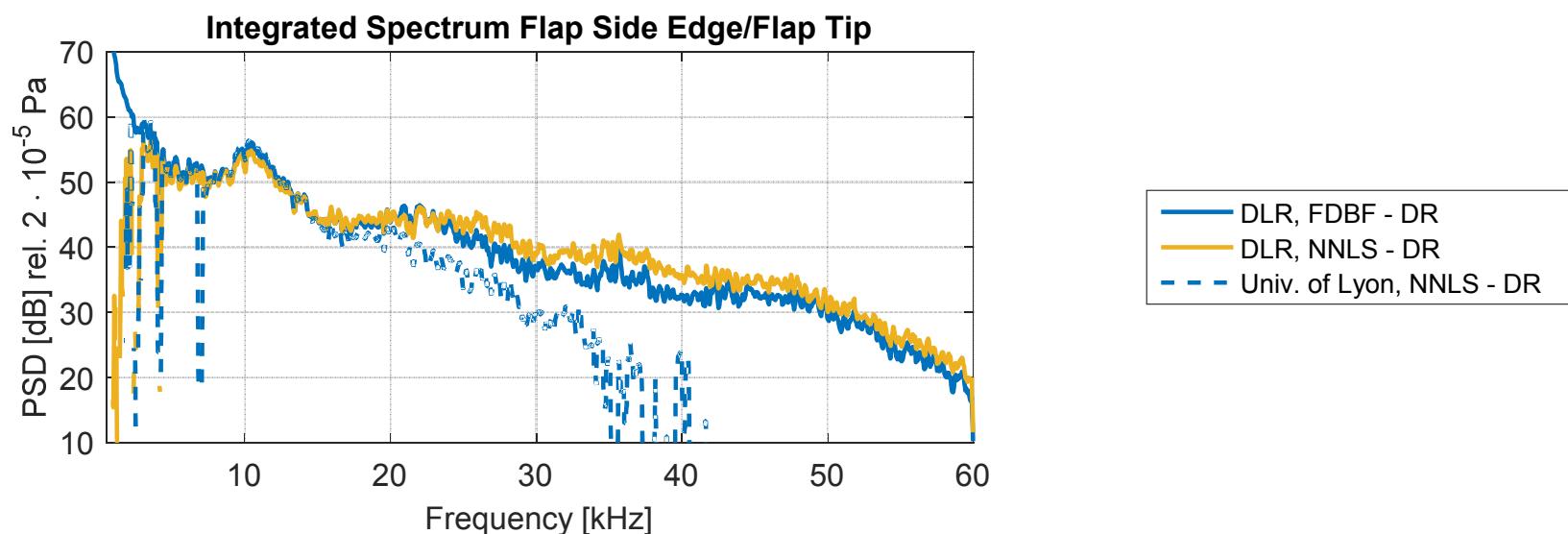
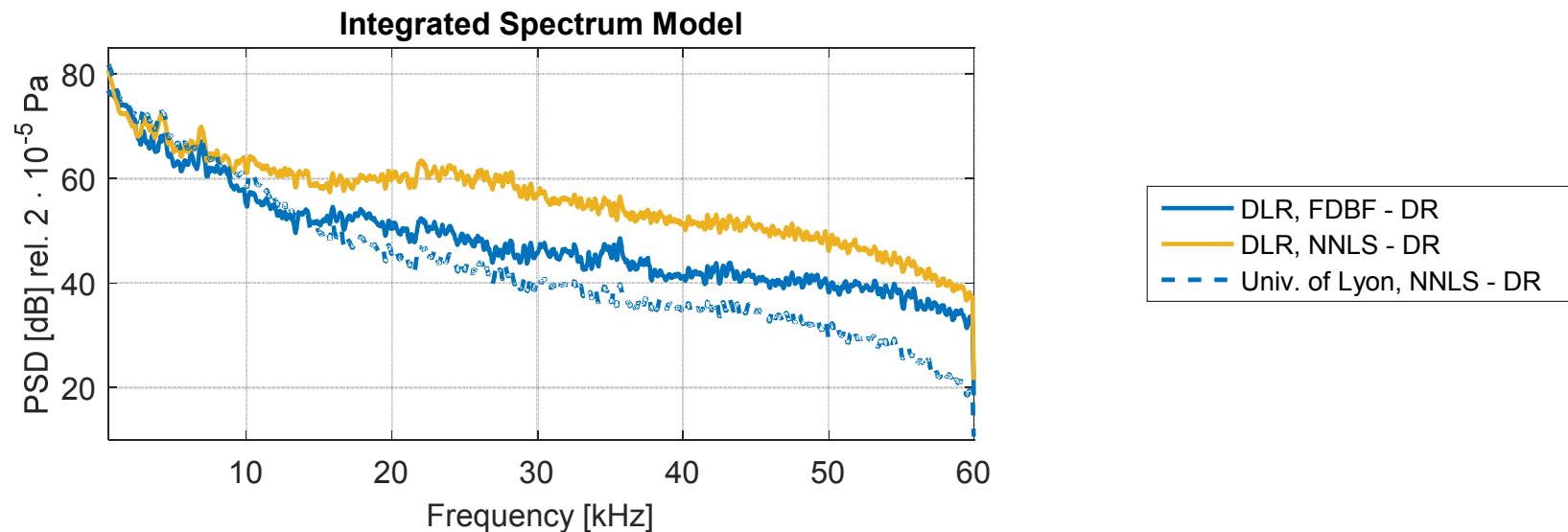


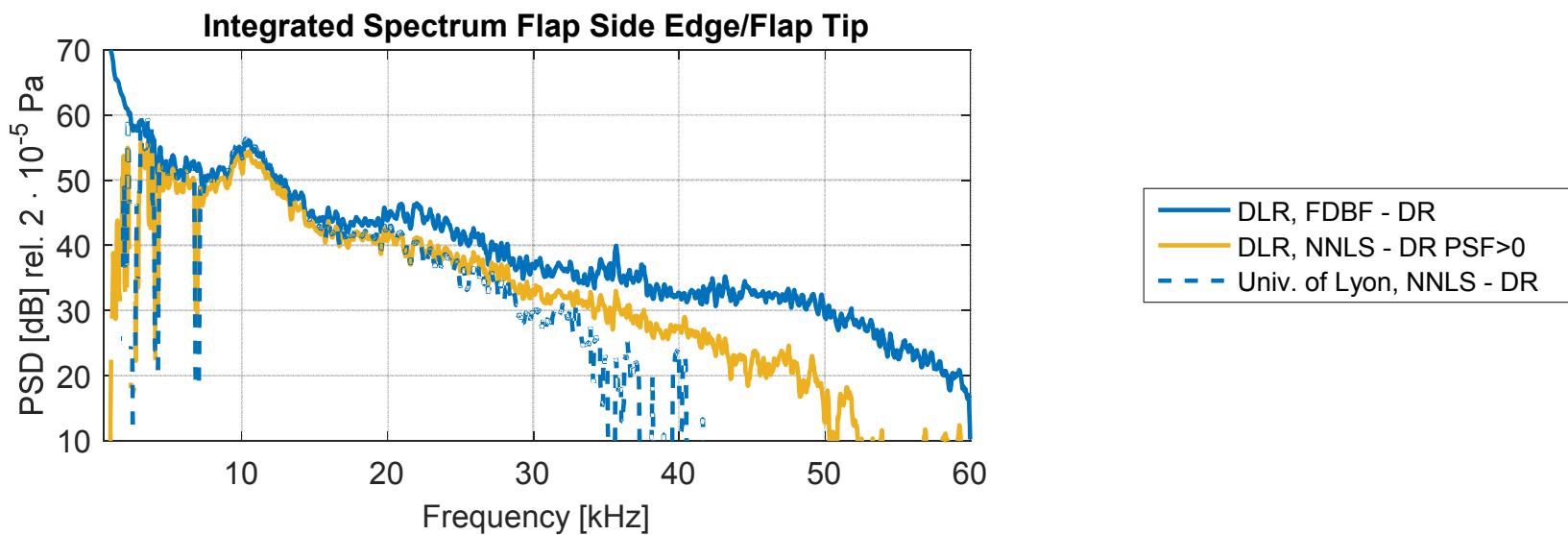
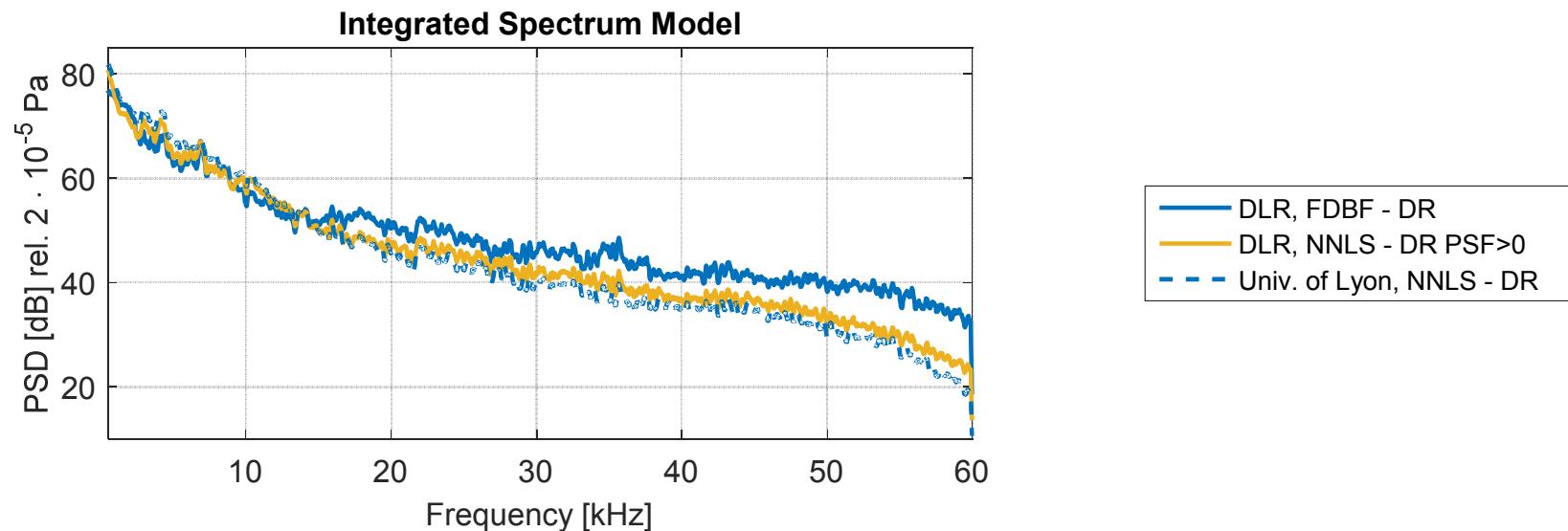


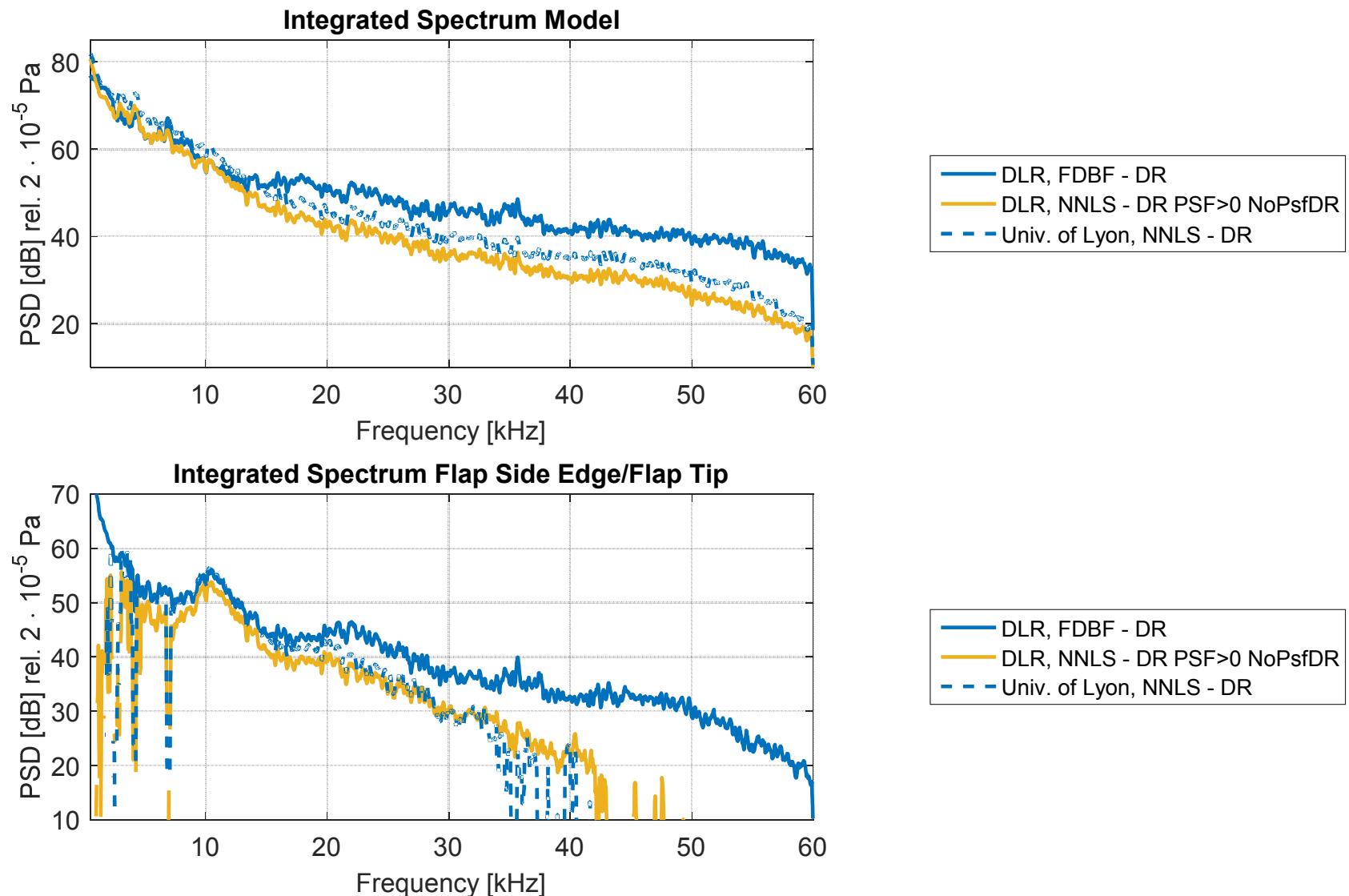
DAMAS NASA/DLR
psf and dirty map allowed
to be negative

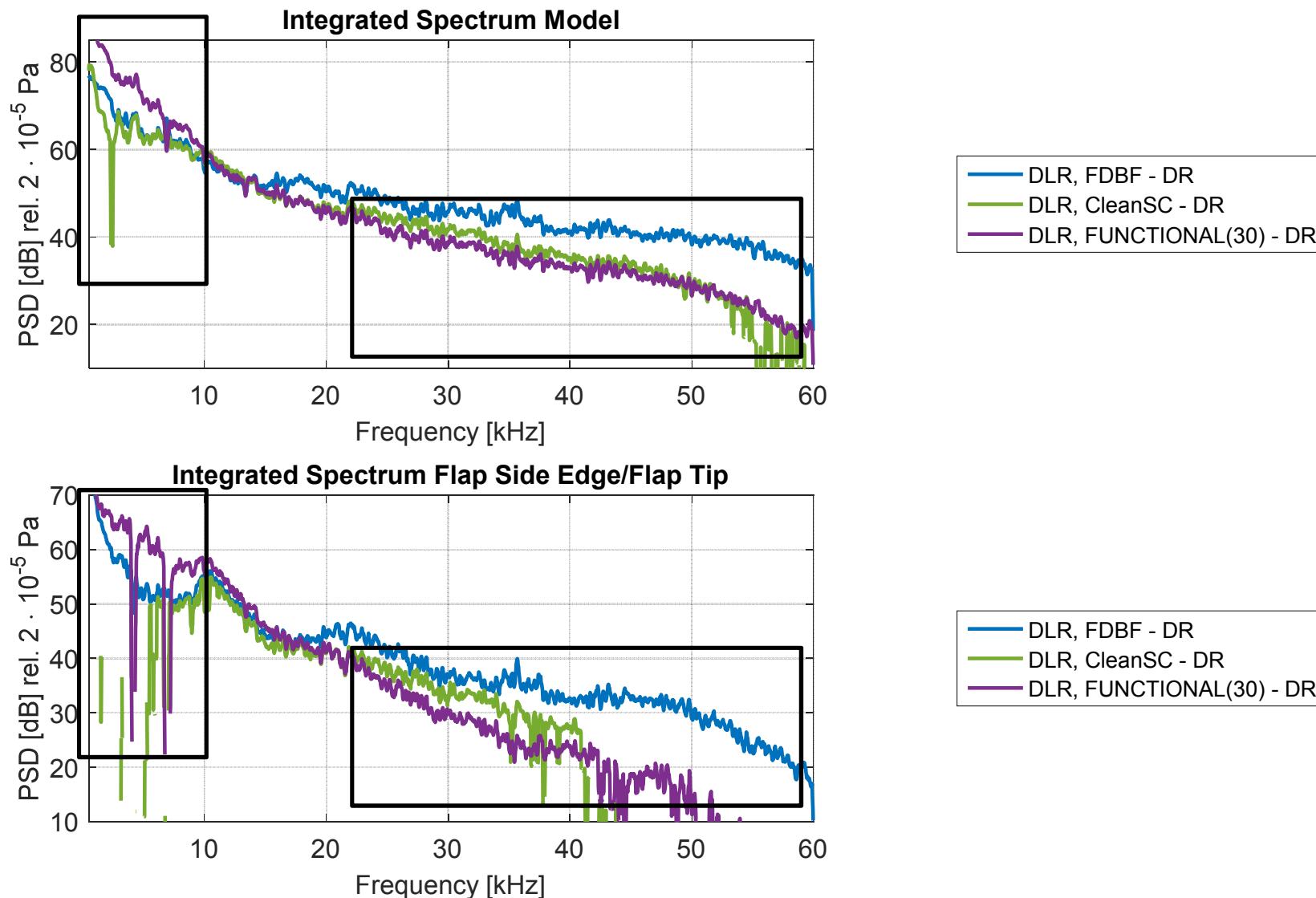






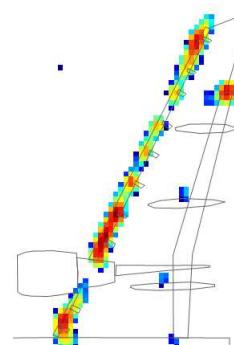
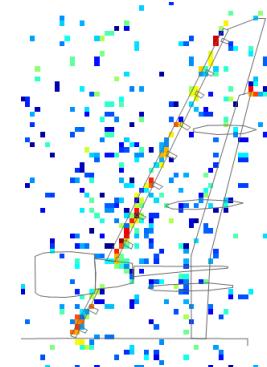
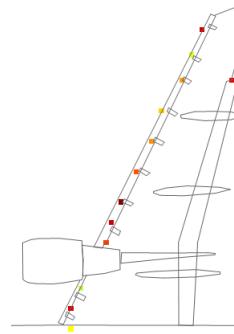






Conclusion

- CleanSC/Bayes/NNLS show similar results source position and amplitude as expected (based on experience)
- DAMAS/NNLS: difficulties at high frequencies (without non-negativity constraint on the psf)
- Definition of the psf and dirty-map is more important than the algorithm
- Functional Beamforming: fastest algorithm, all sources can be identified psf artifacts highly reduced





Submission to Testcase DLR 1

Create a HDF5 – CsmOptional File with your beamforming result

Testcase: 2009_DO728_dp59

M = 0.25

Angle of Attack: 3°

and **provide this File to** one of the following

Daniel Ernst Daniel.Ernst@dlr.de

Thomas Ahlefeldt Thomas.Ahlefeldt@dlr.de

Carsten Spehr Carsten.Spehr@dlr.de

We can provide transfer for large files upon request!

We will

1. Check your submission
2. Calculate the integrated spectra
3. Upload your solution



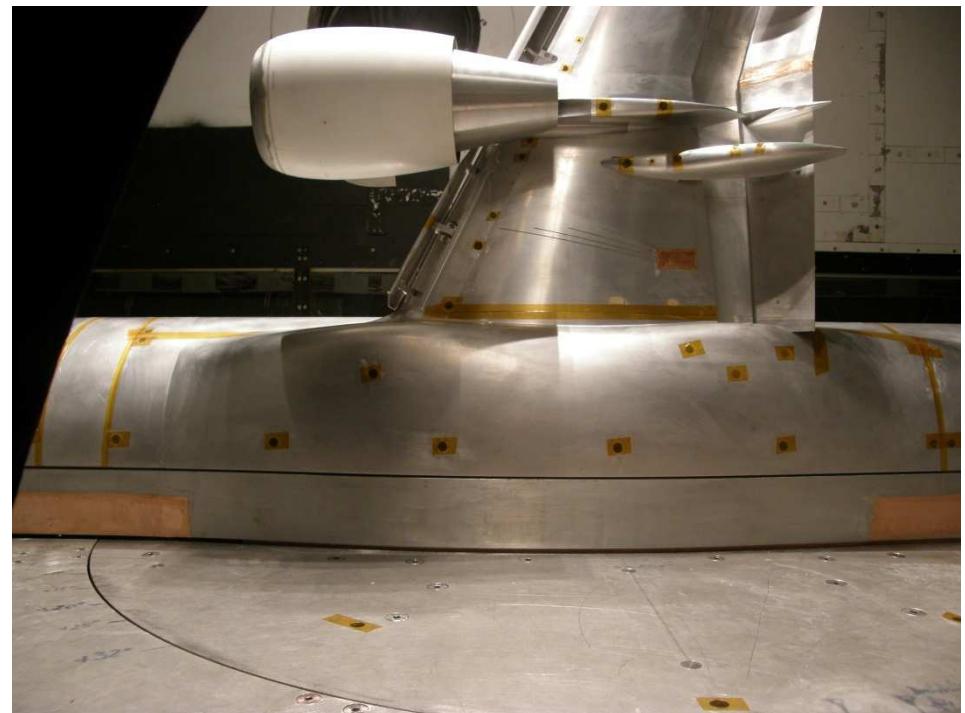
Appendix

- 1) Integrated Spectra „Full Model“ and „Flap Tip“
- 2) 3rd Octave - Source Maps



Detail Pictures of the DO-728 Model

view from microphone array



© Thomas Ahlefeldt



Detail Pictures of the DO-728 Model

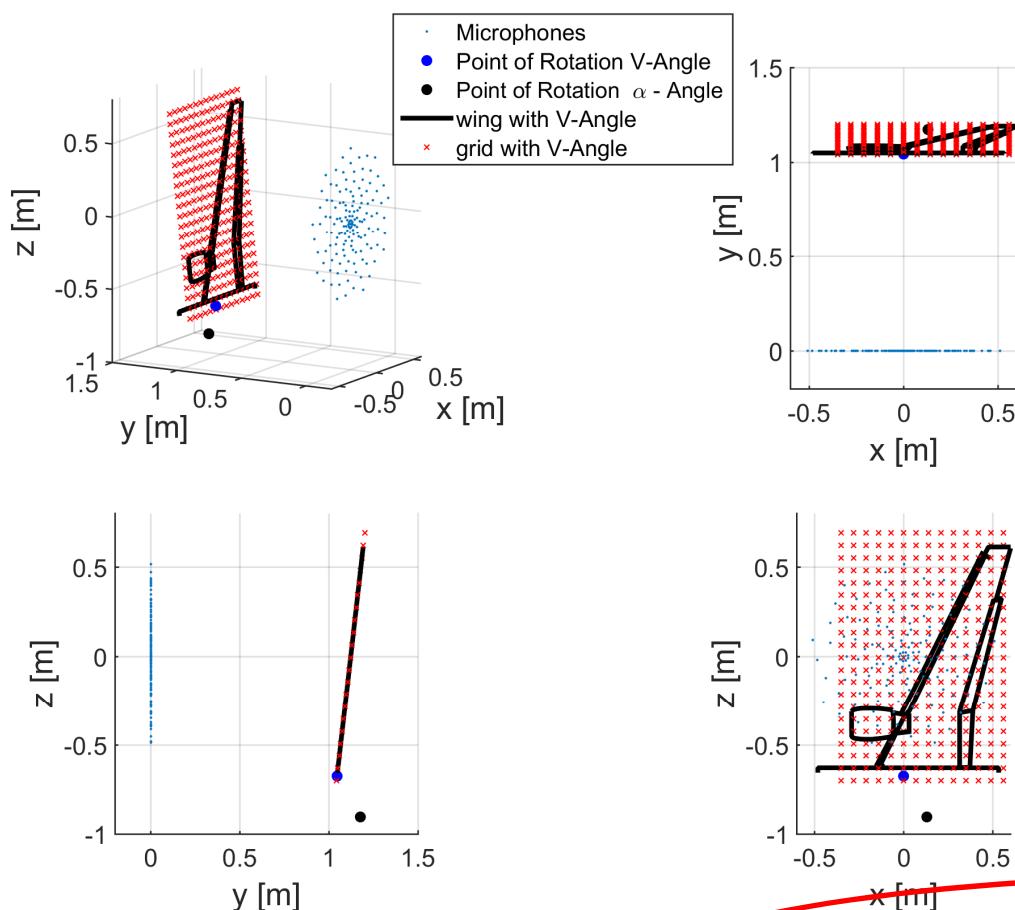
view from microphone array



© Thomas Ahlefeldt



Focus-Grid Rotation



Focus grid

$$x = -0.35, \dots, 0.7 \text{ m}$$

$$y = 1.045 \text{ m}$$

$$z = -0.75, \dots, 0.7 \text{ m}$$

V-angle: -6.5°

Rotation axis: x (right hand rule)

Point of rotation:

$$x = 0\text{m}, y = 1.045\text{m}, z = -0.675\text{m}$$

Angle of attack (negative): -3°

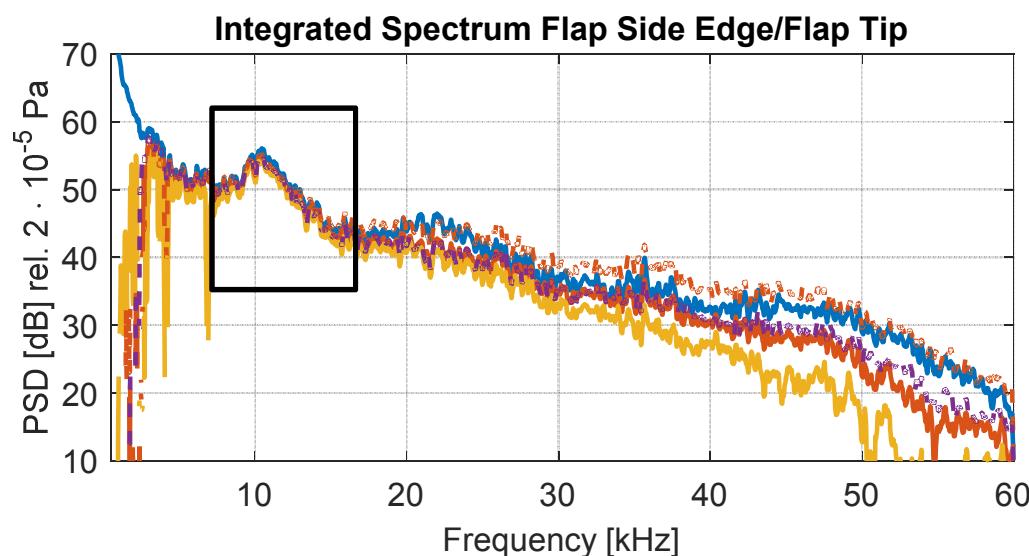
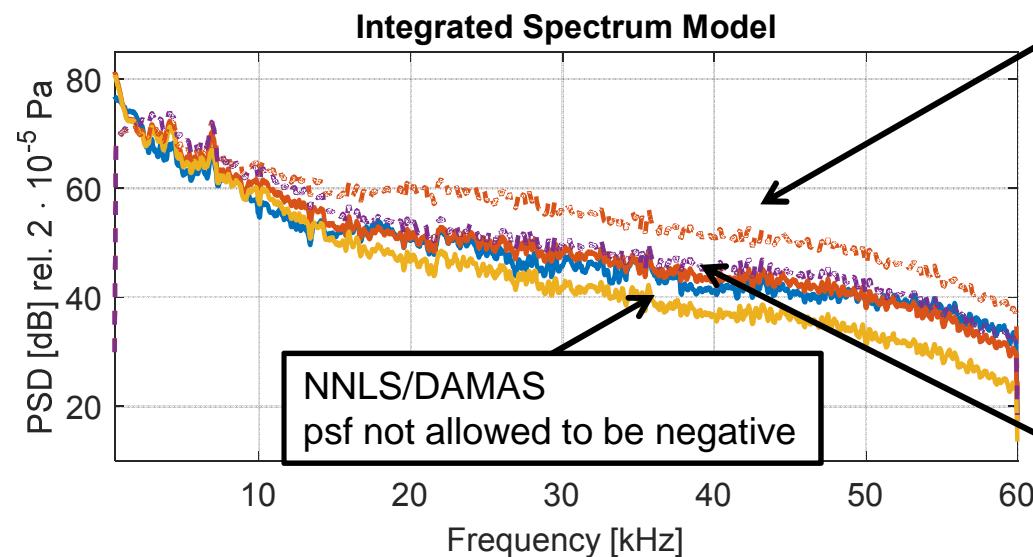
Rotation axis: z (right hand rule)

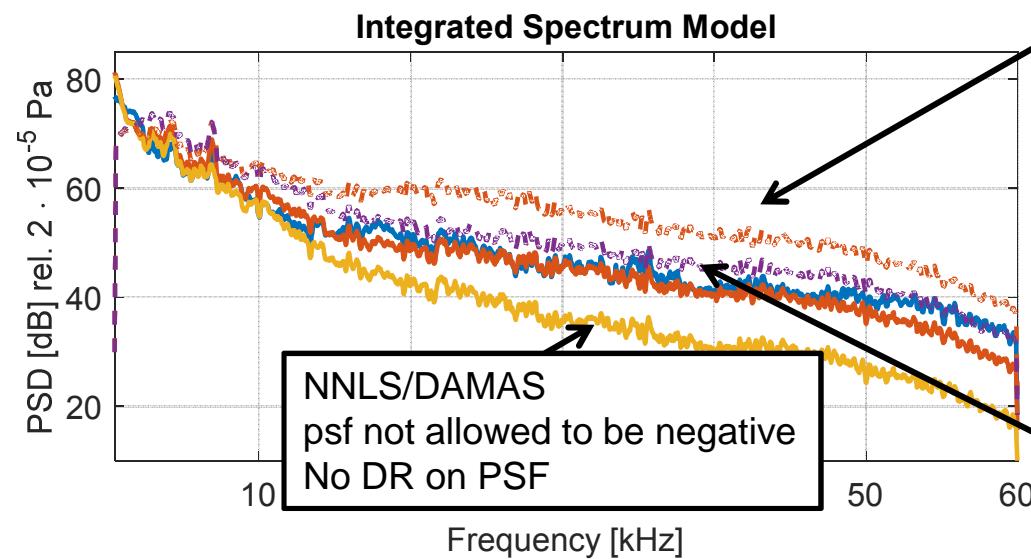
Point of rotation:

$$x = 0.13\text{m}, y = 1.175\text{m}, z = -0.675\text{m}$$

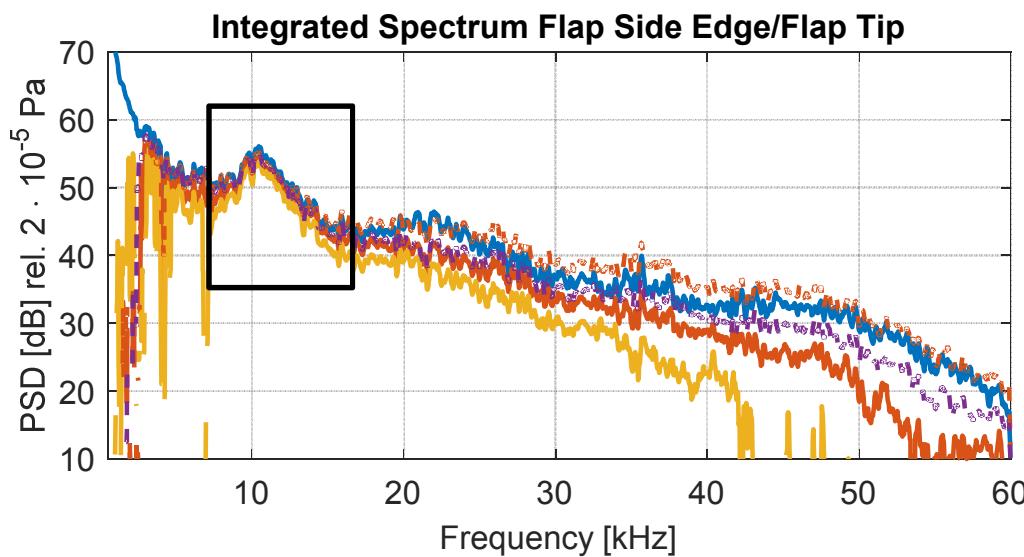
Rotated Focus Grids available @ TU-Cottbus Server





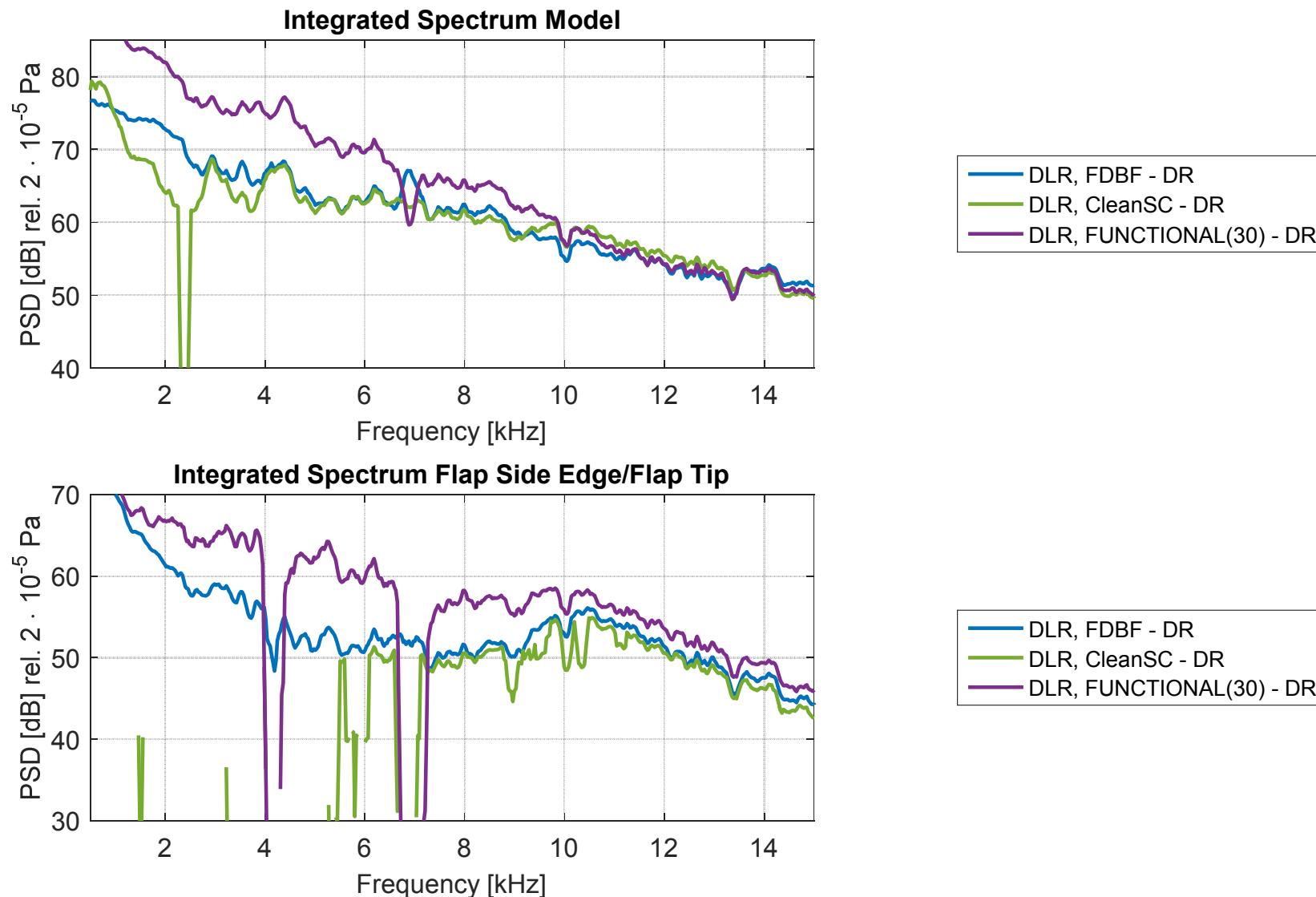


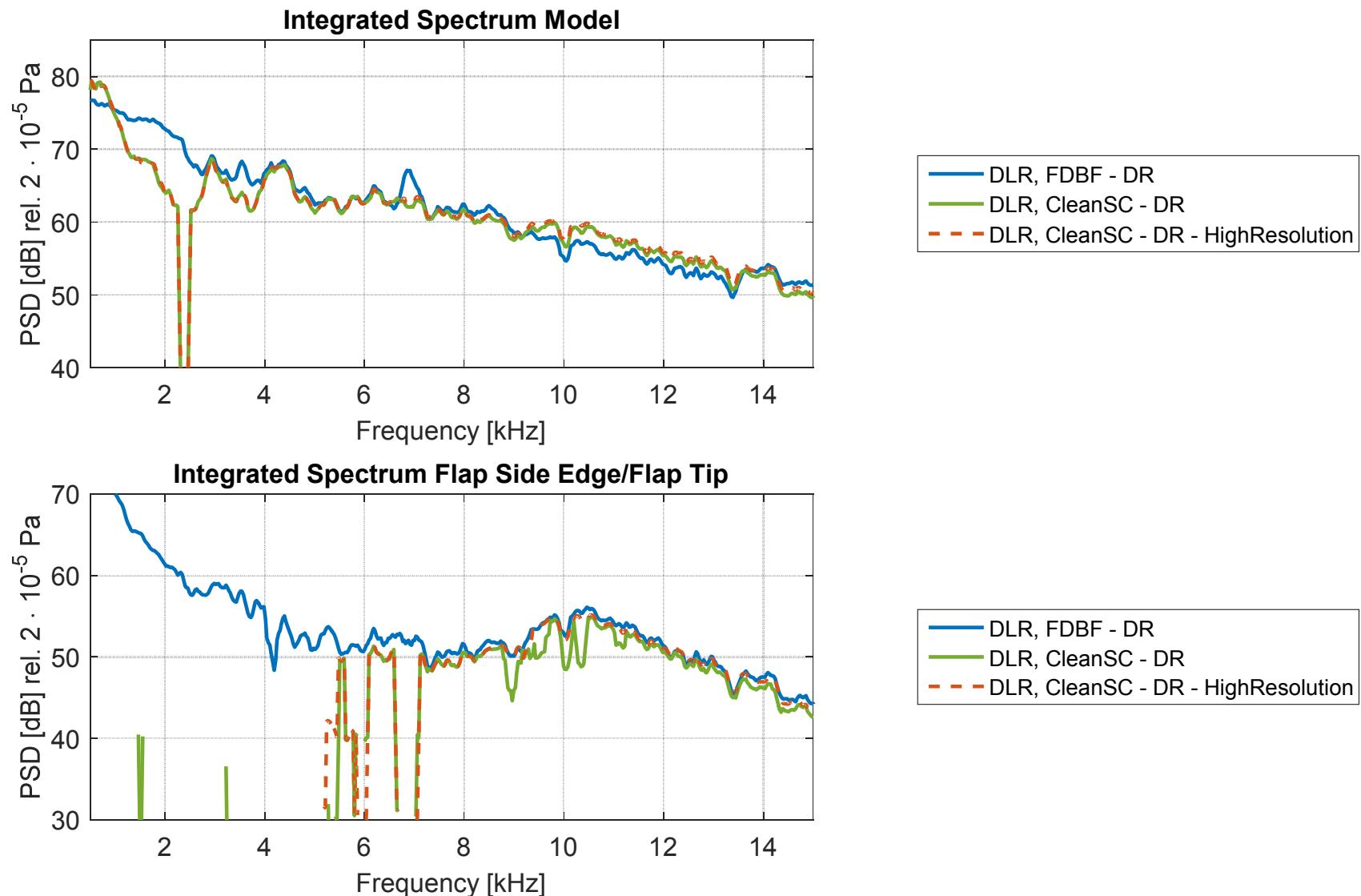
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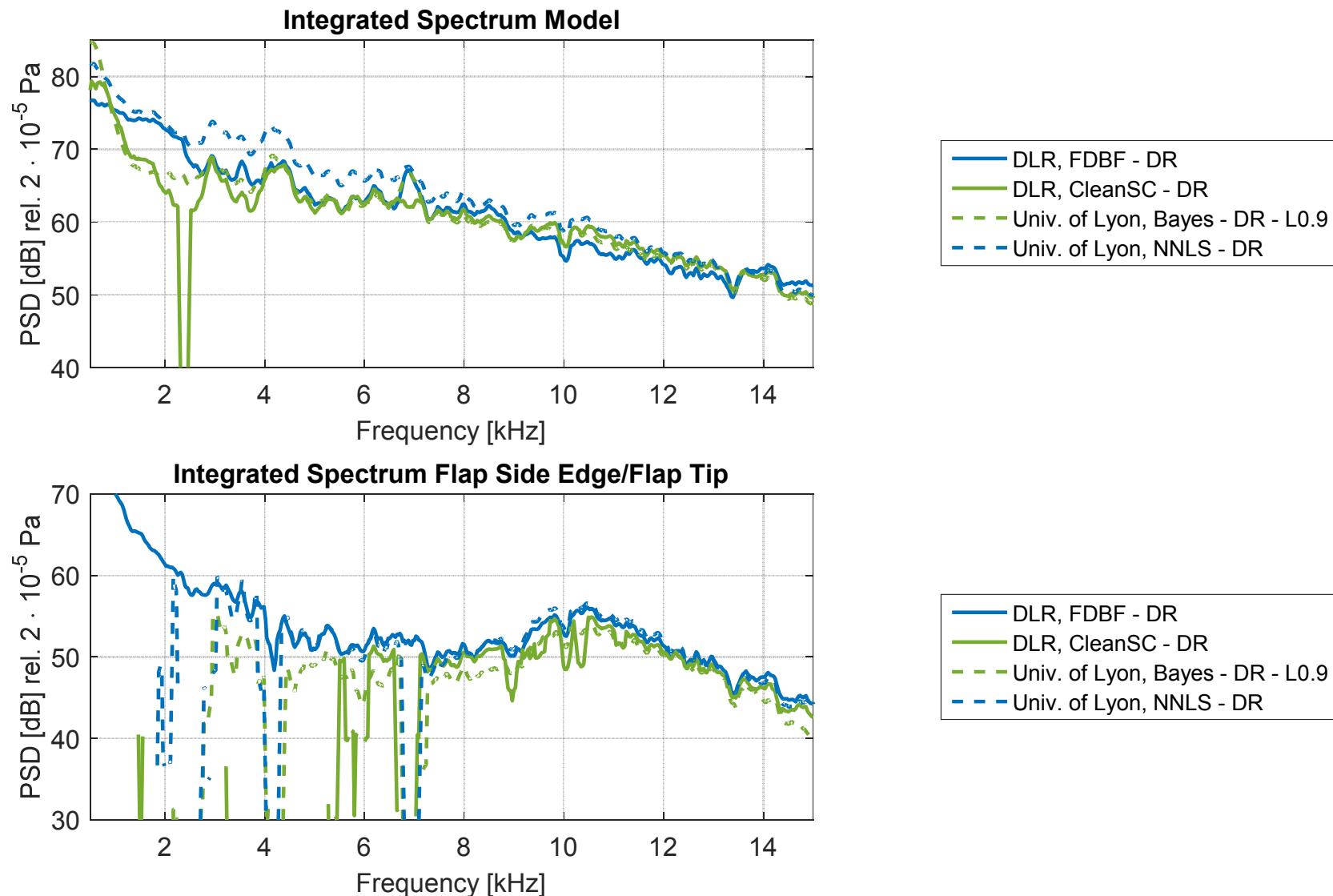


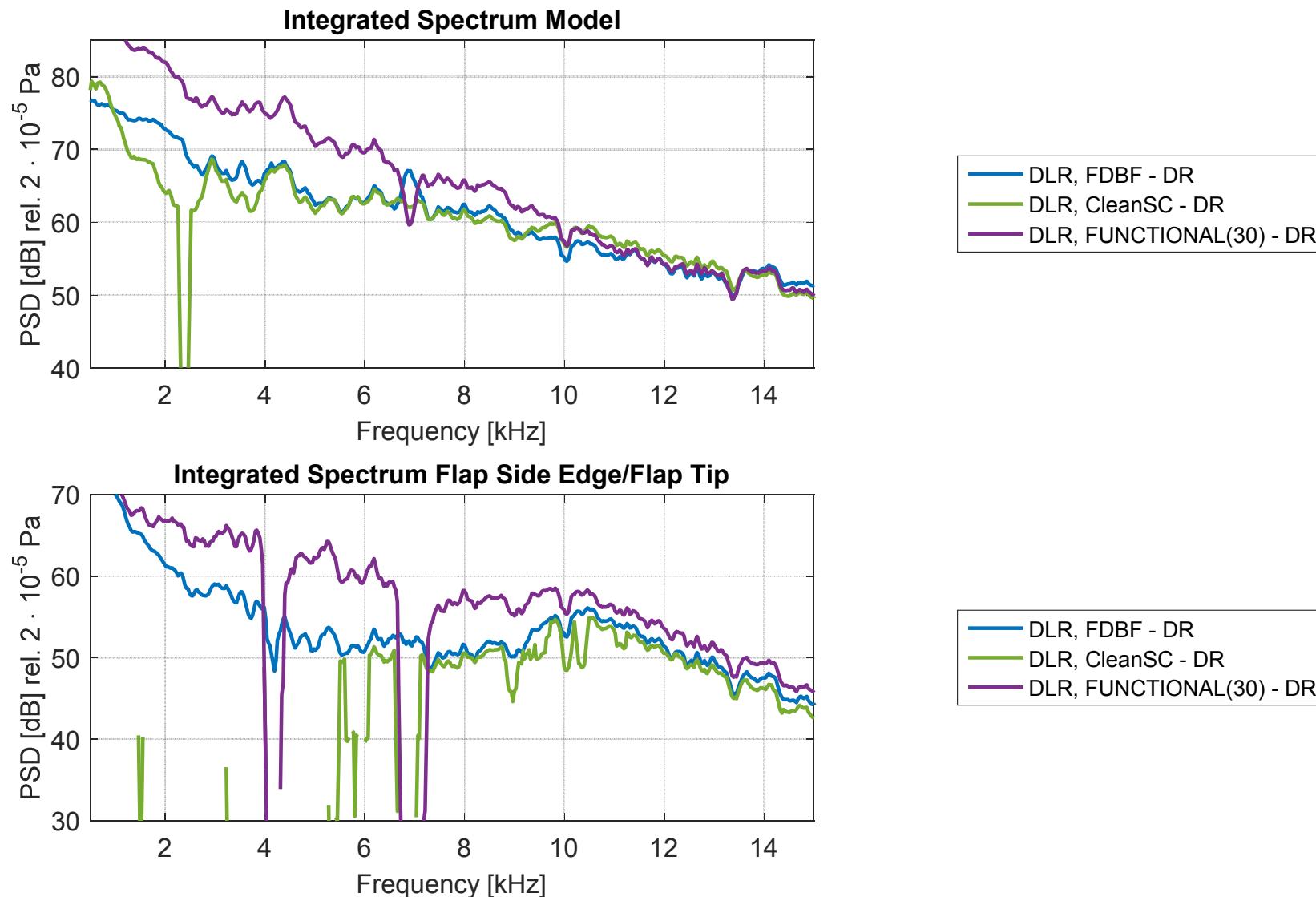
DAMAS
EigDR

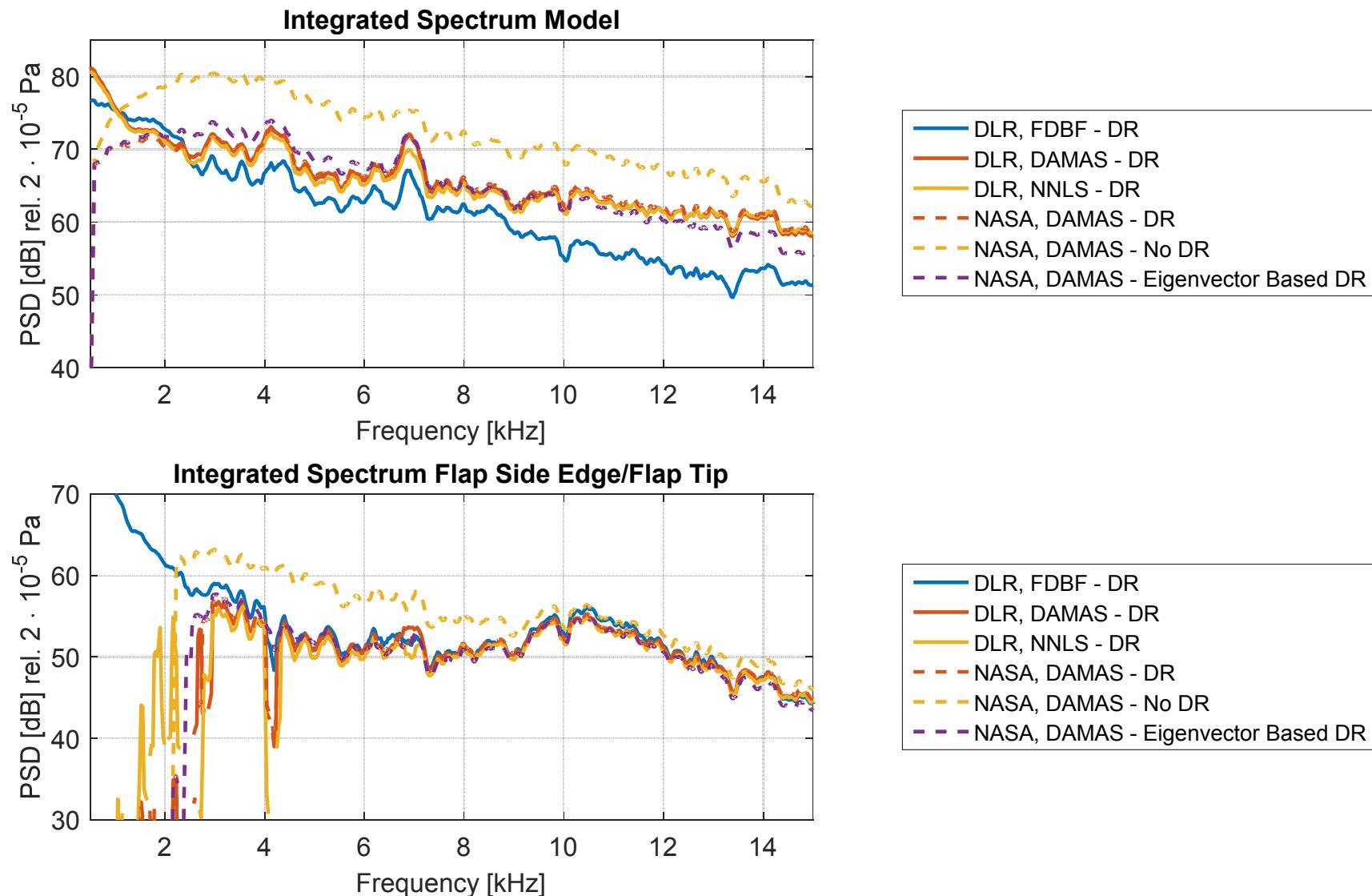


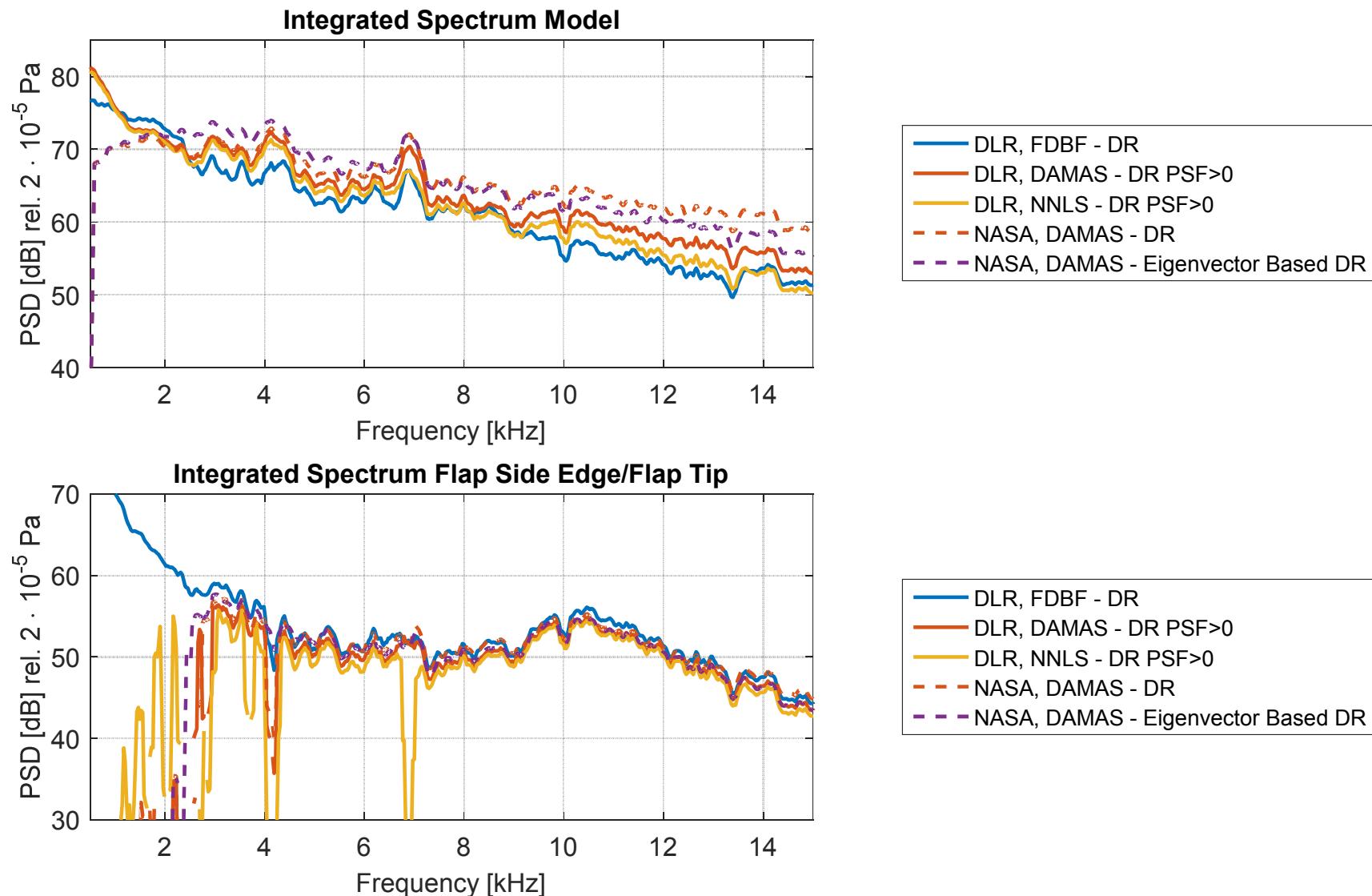


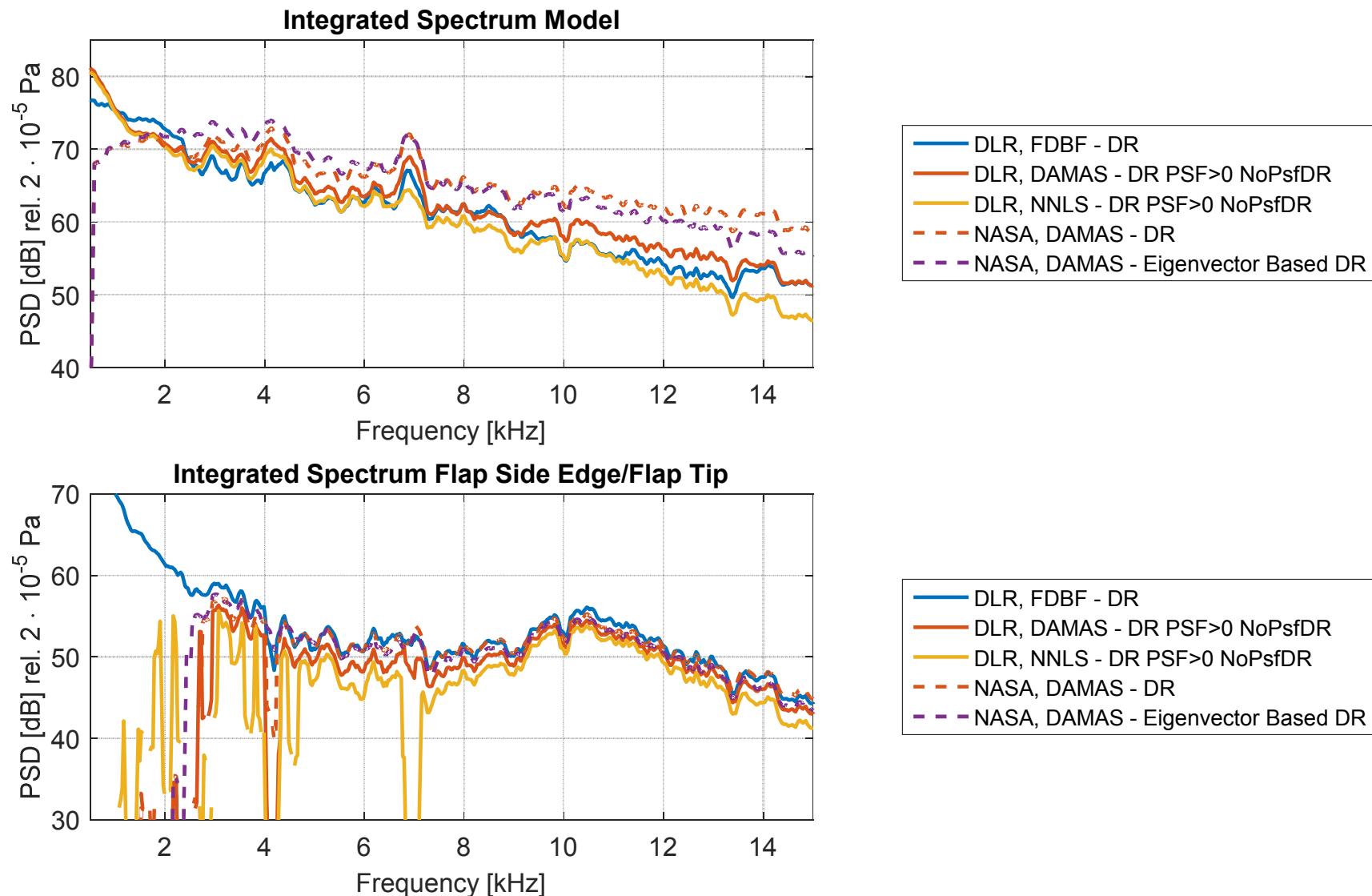


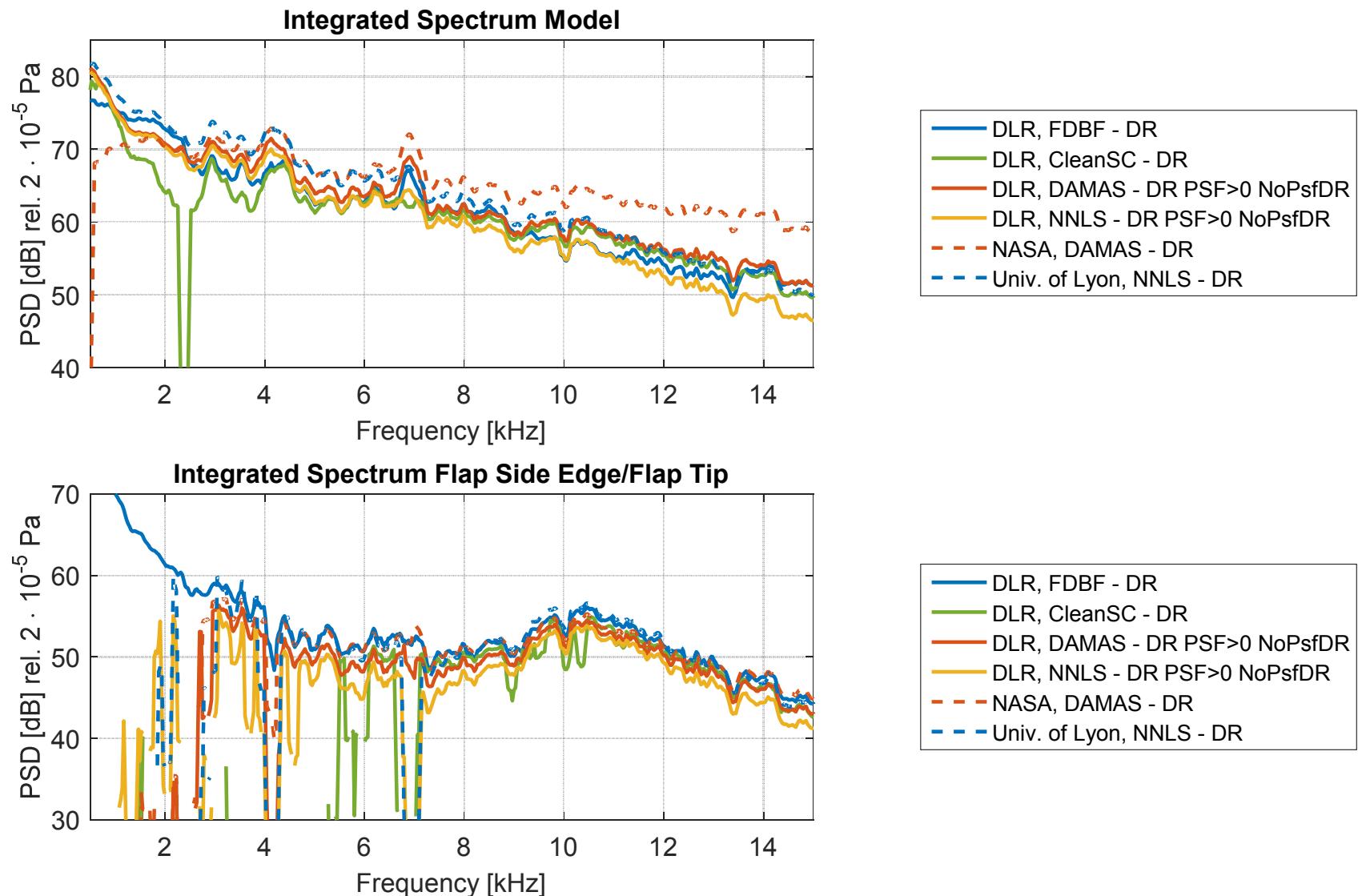


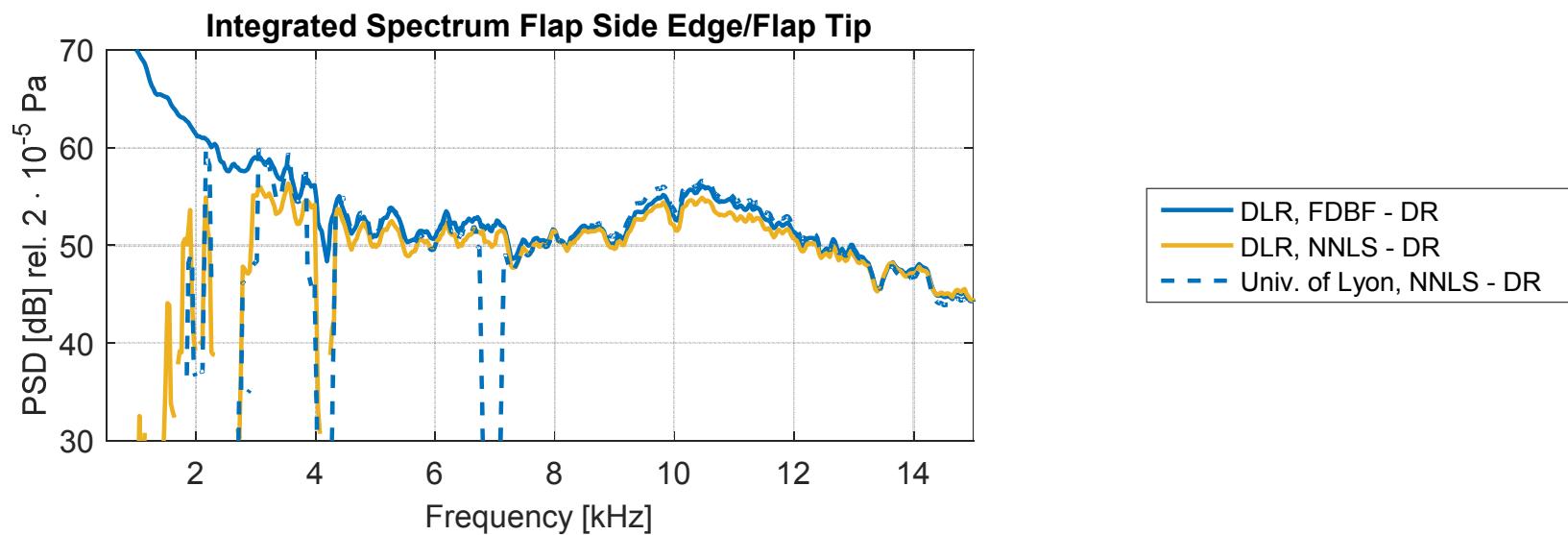
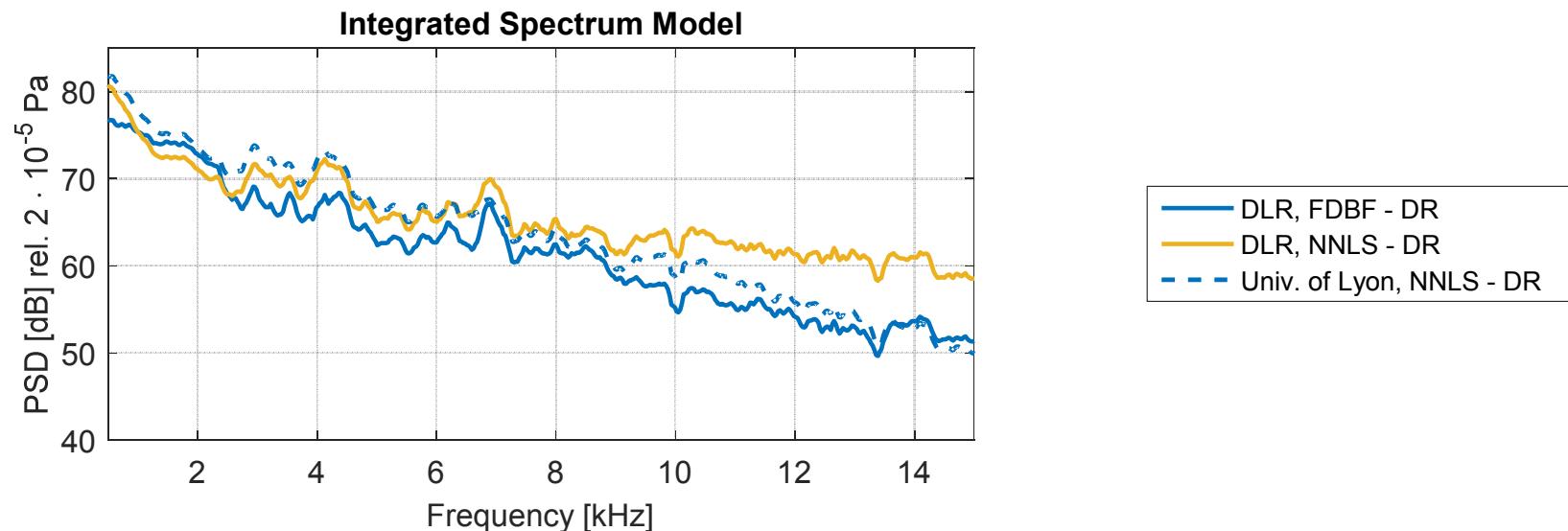


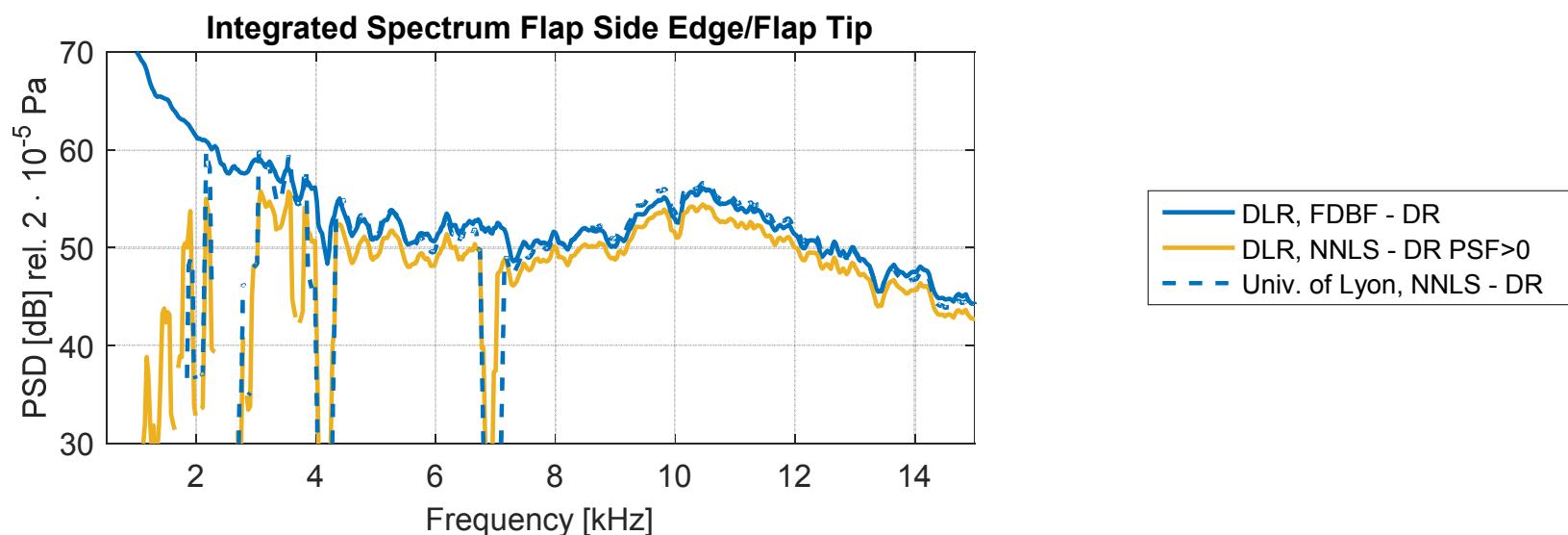
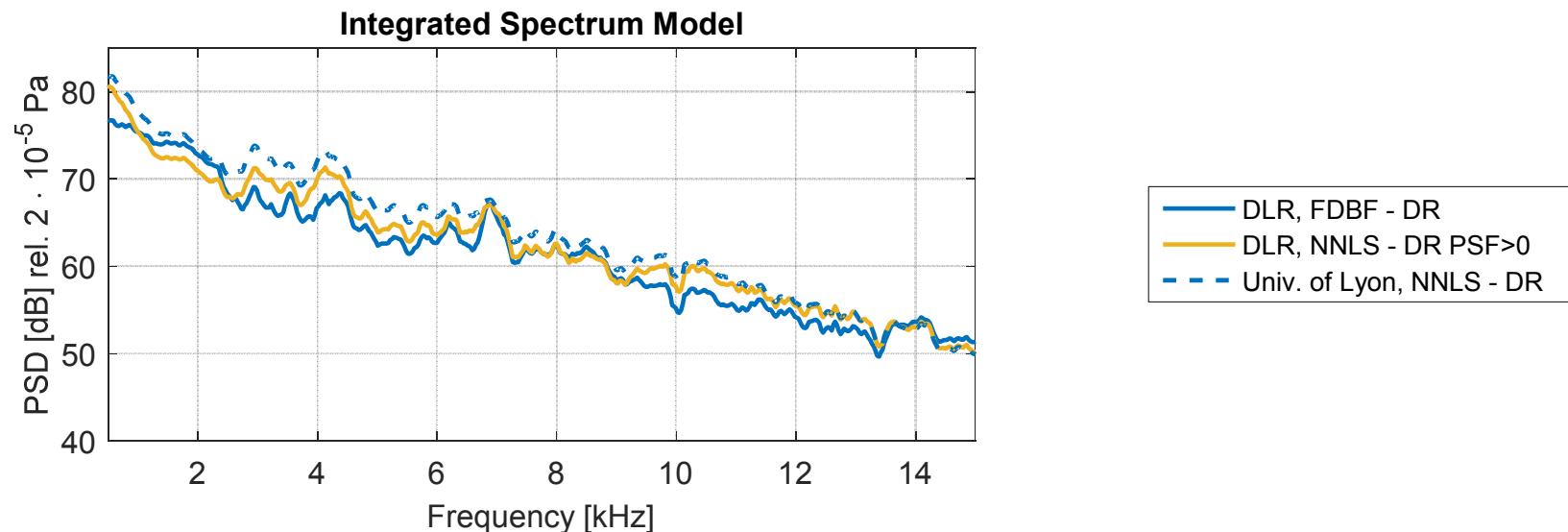


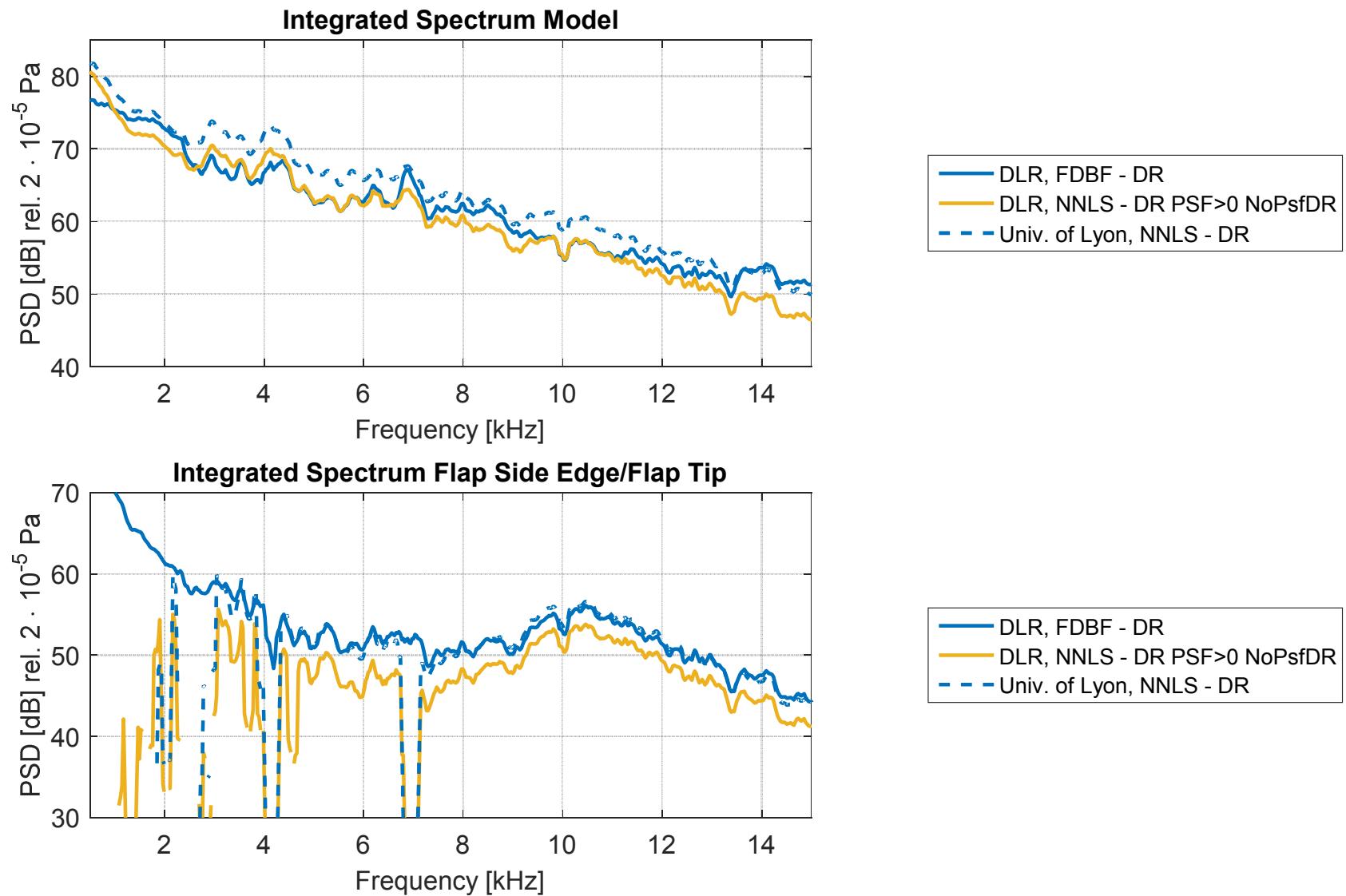


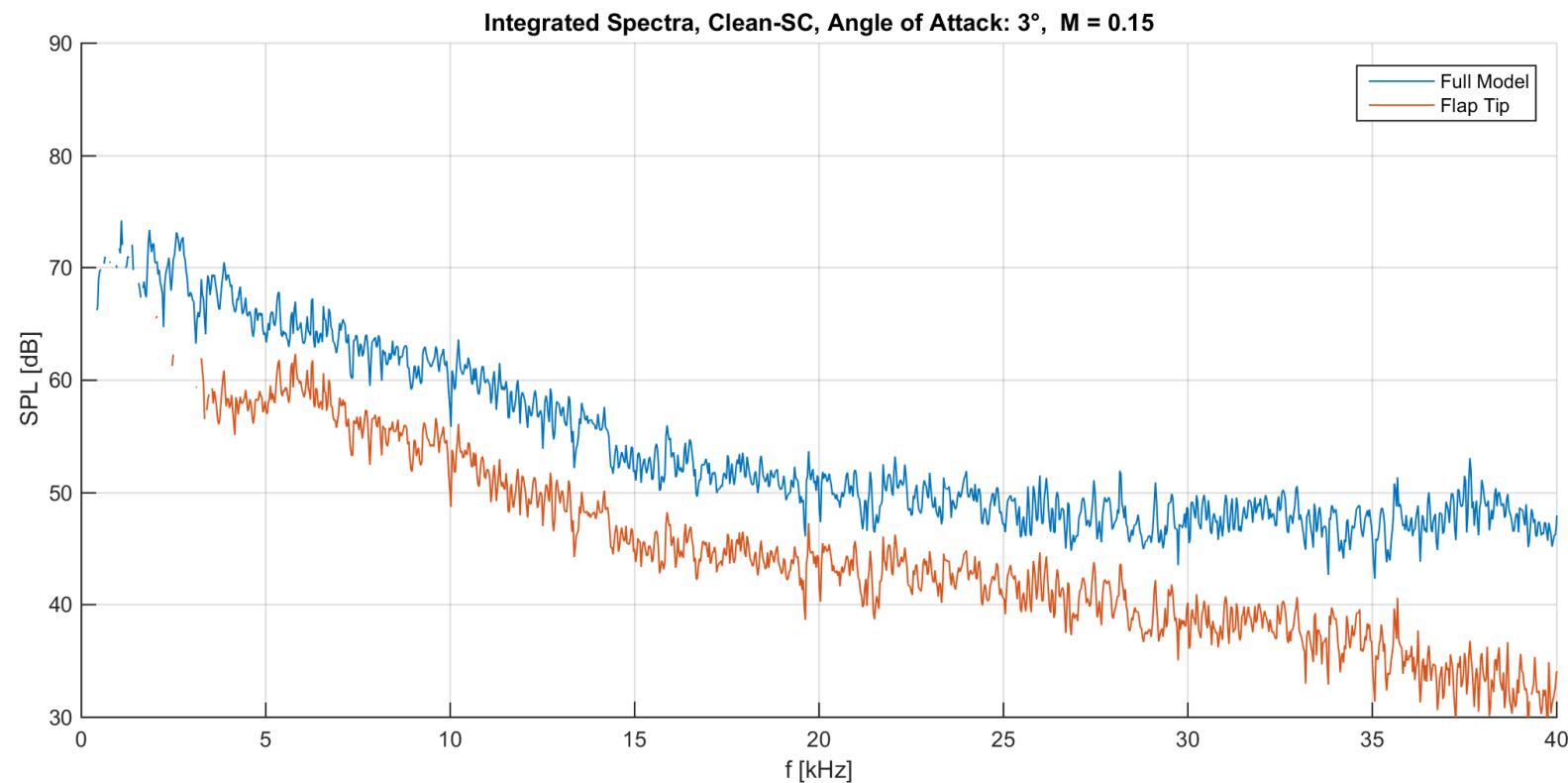


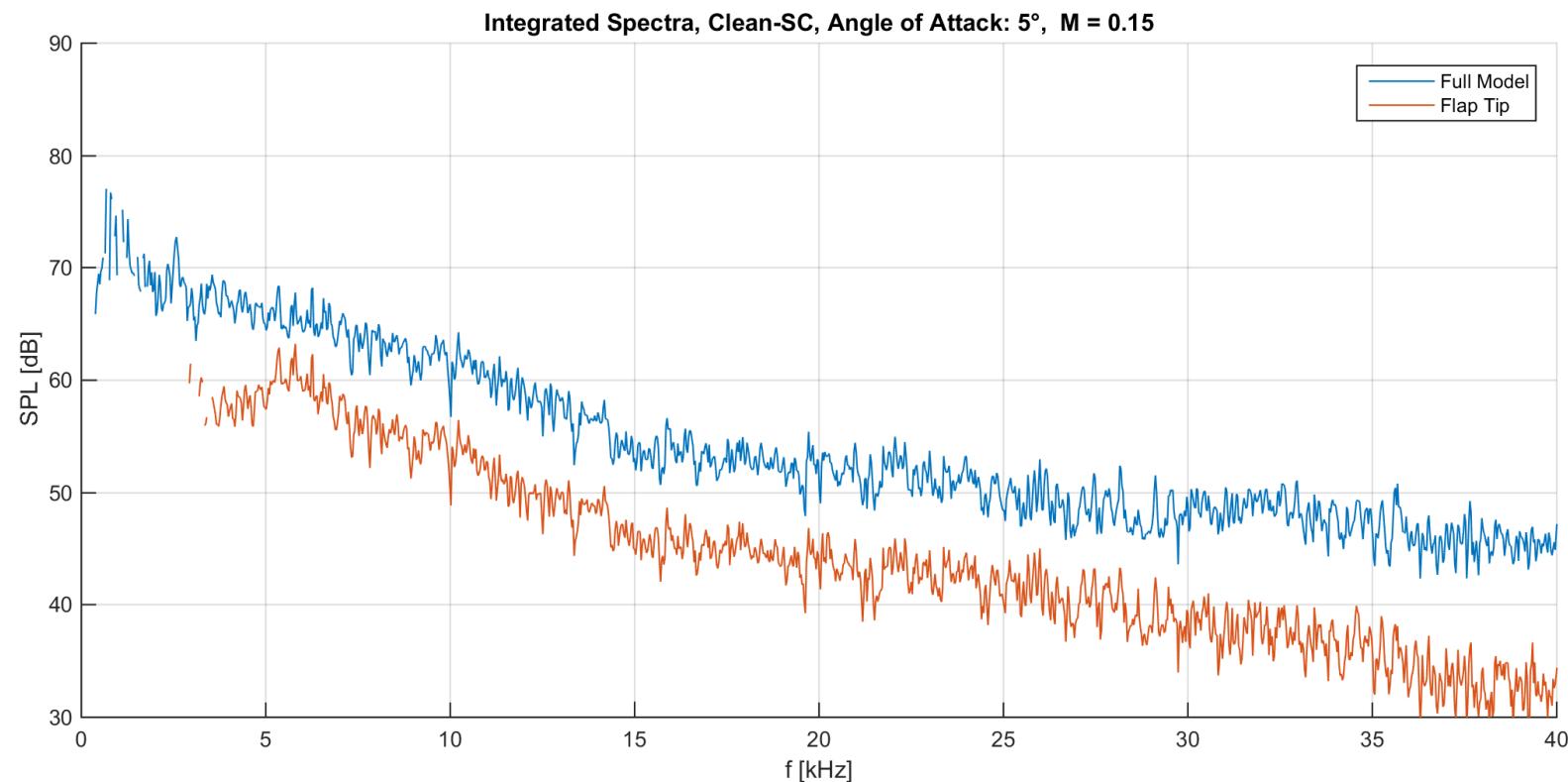


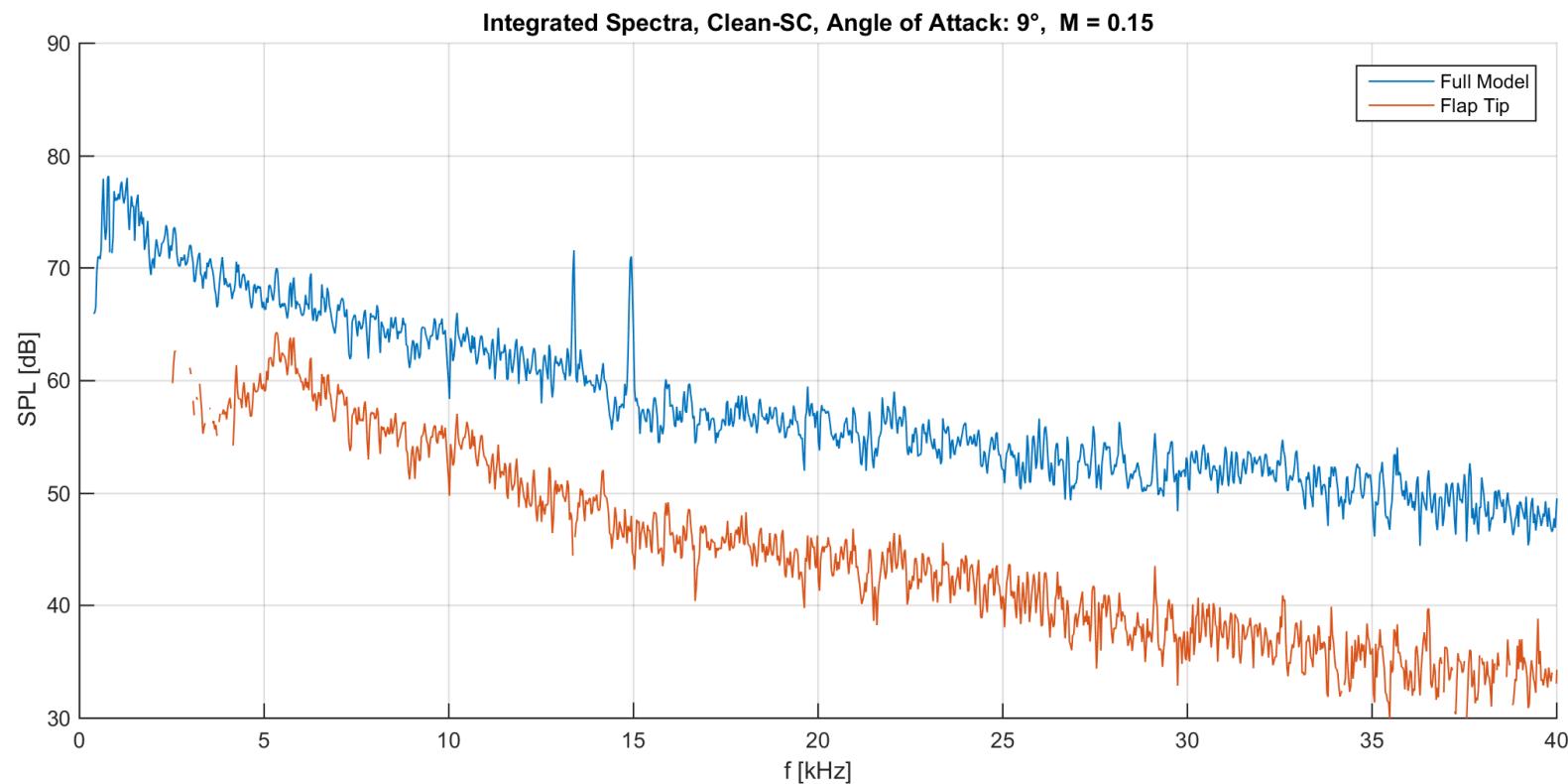


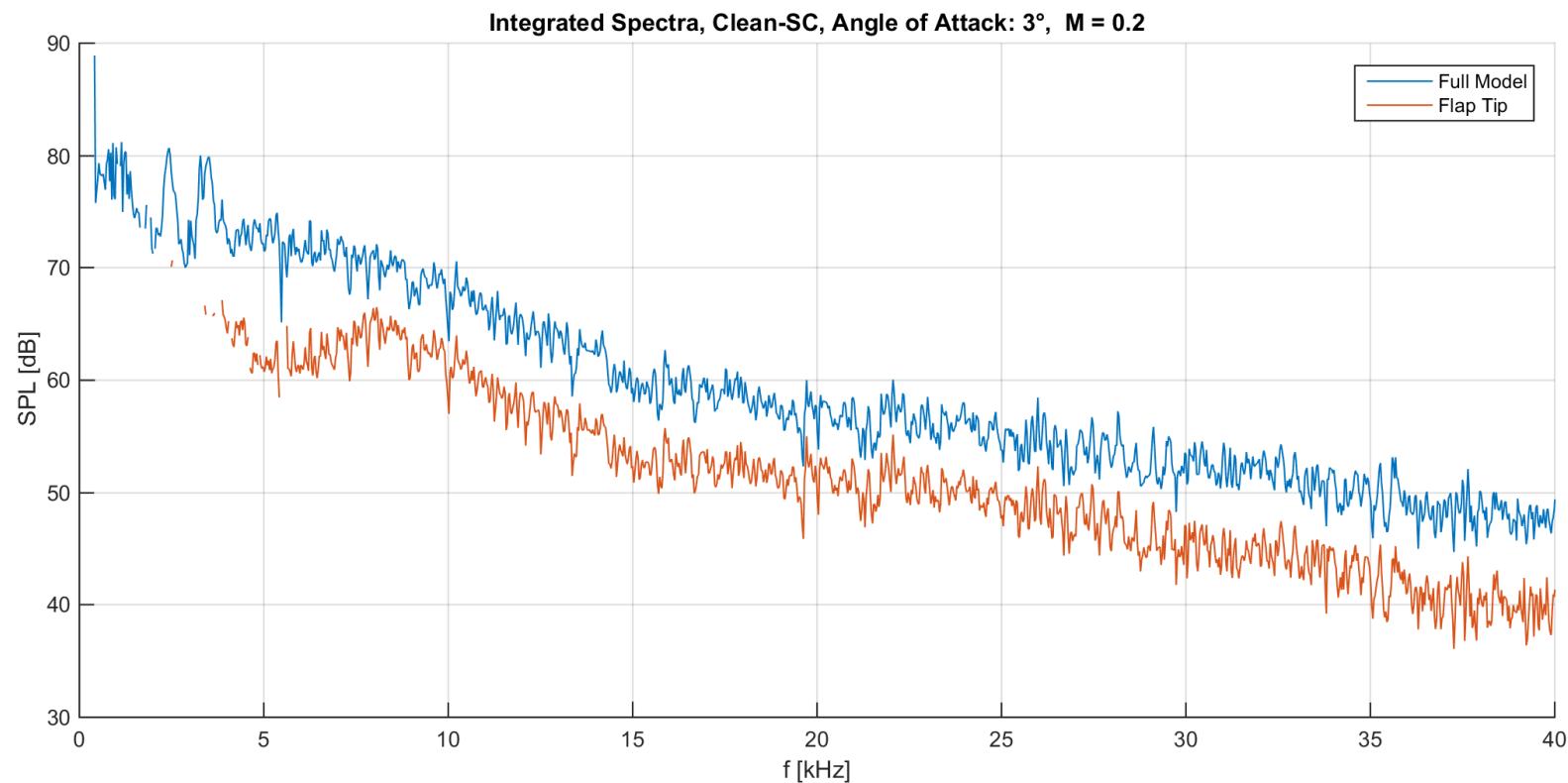


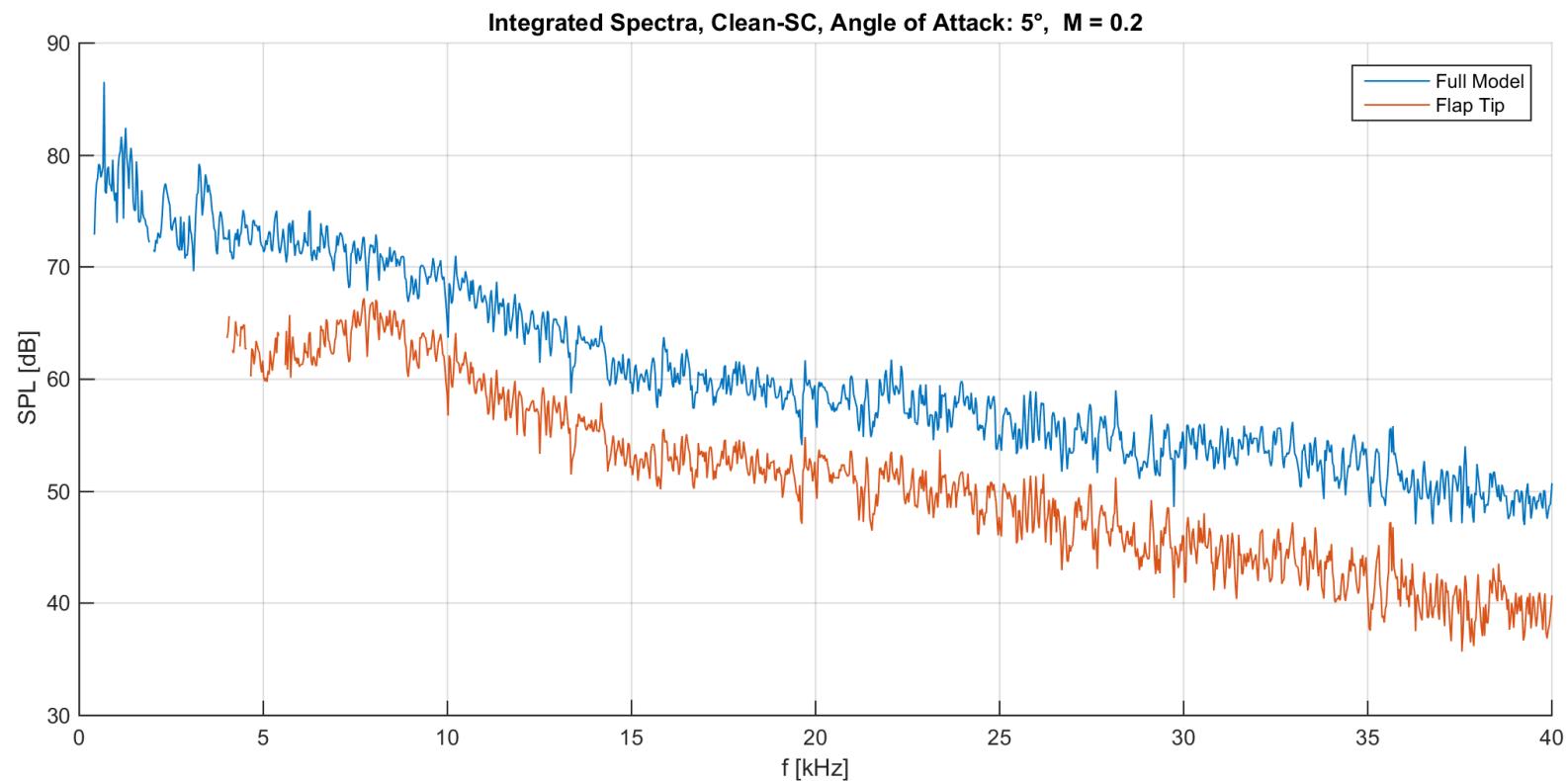


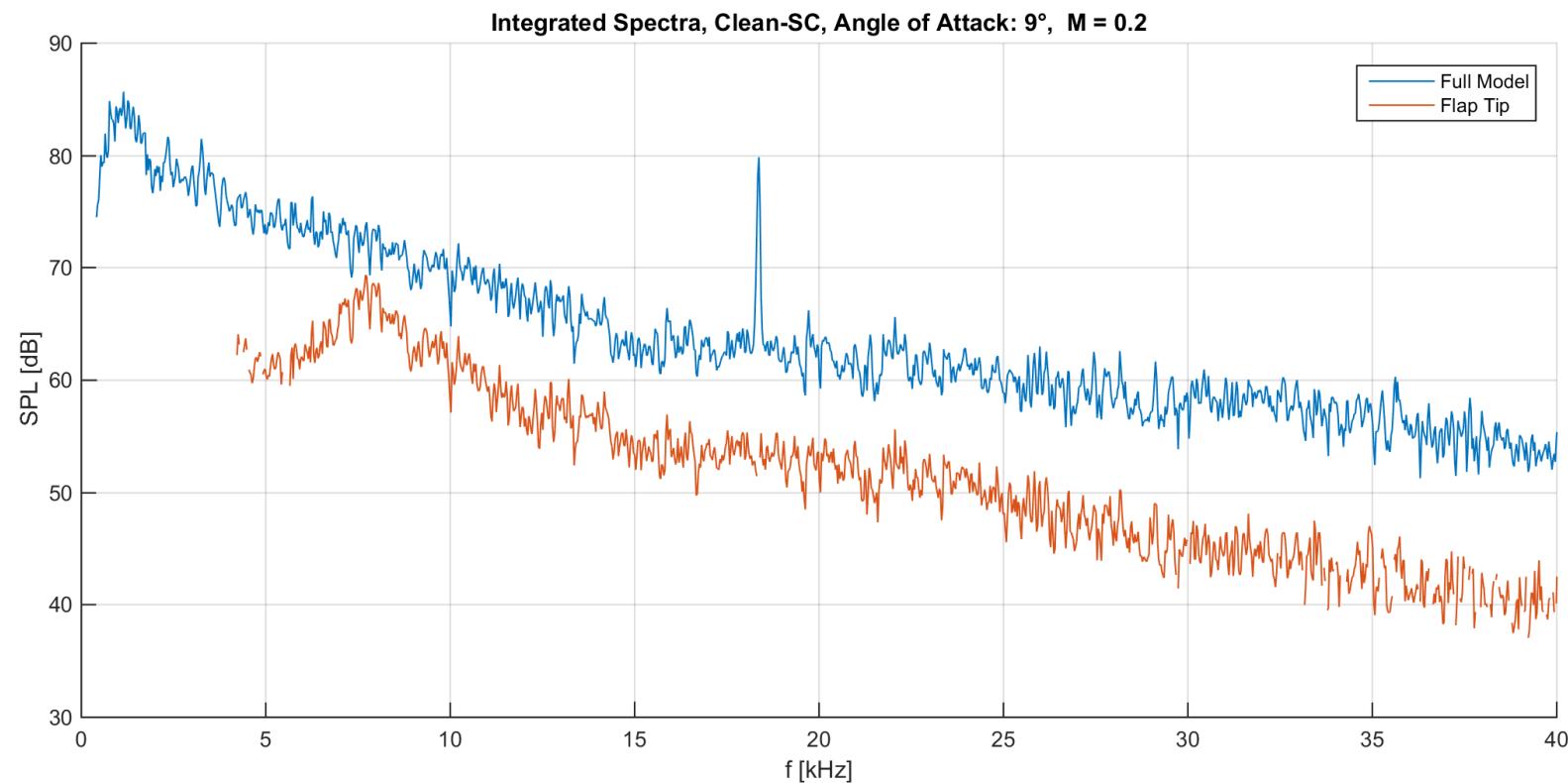


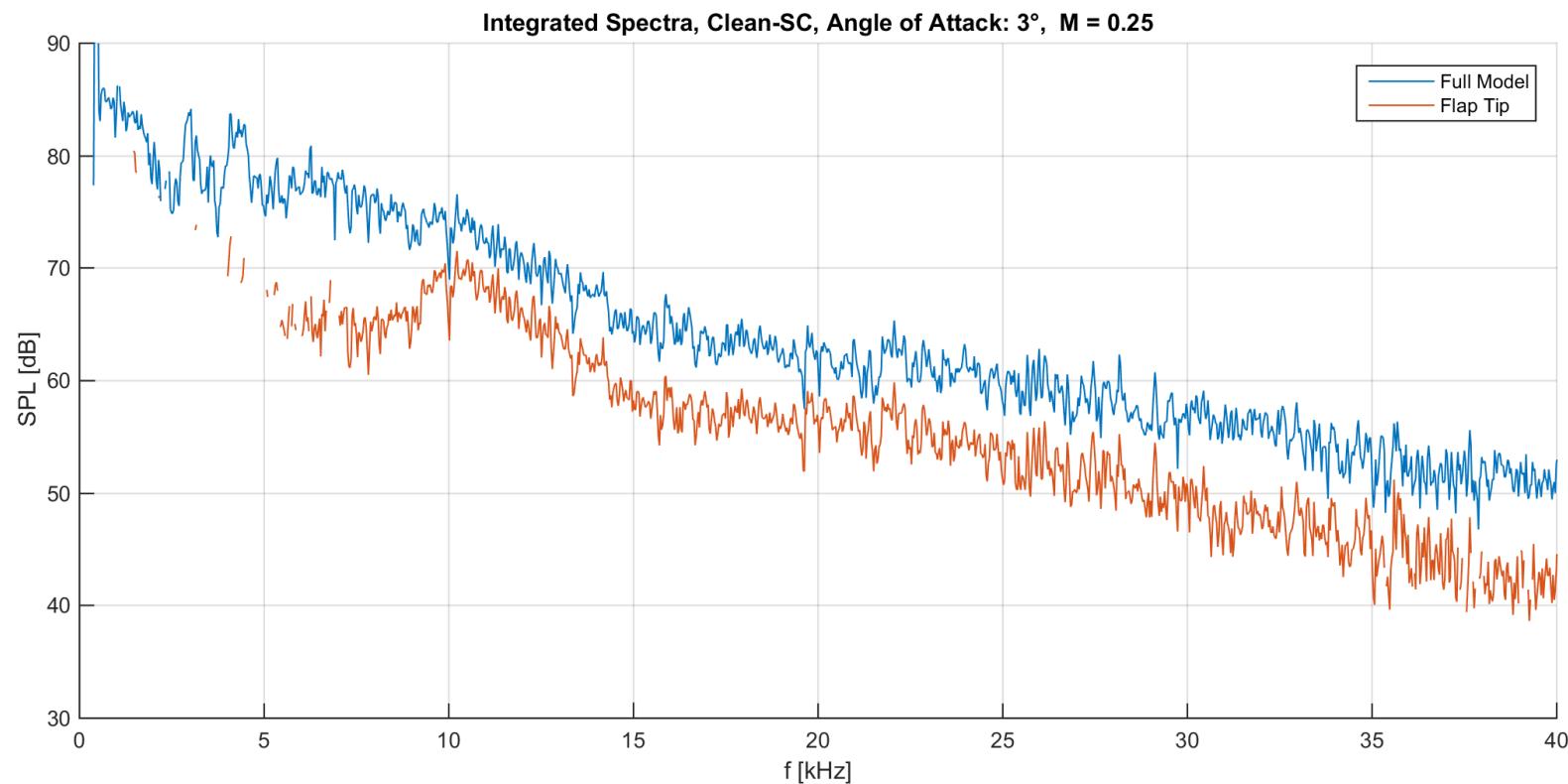


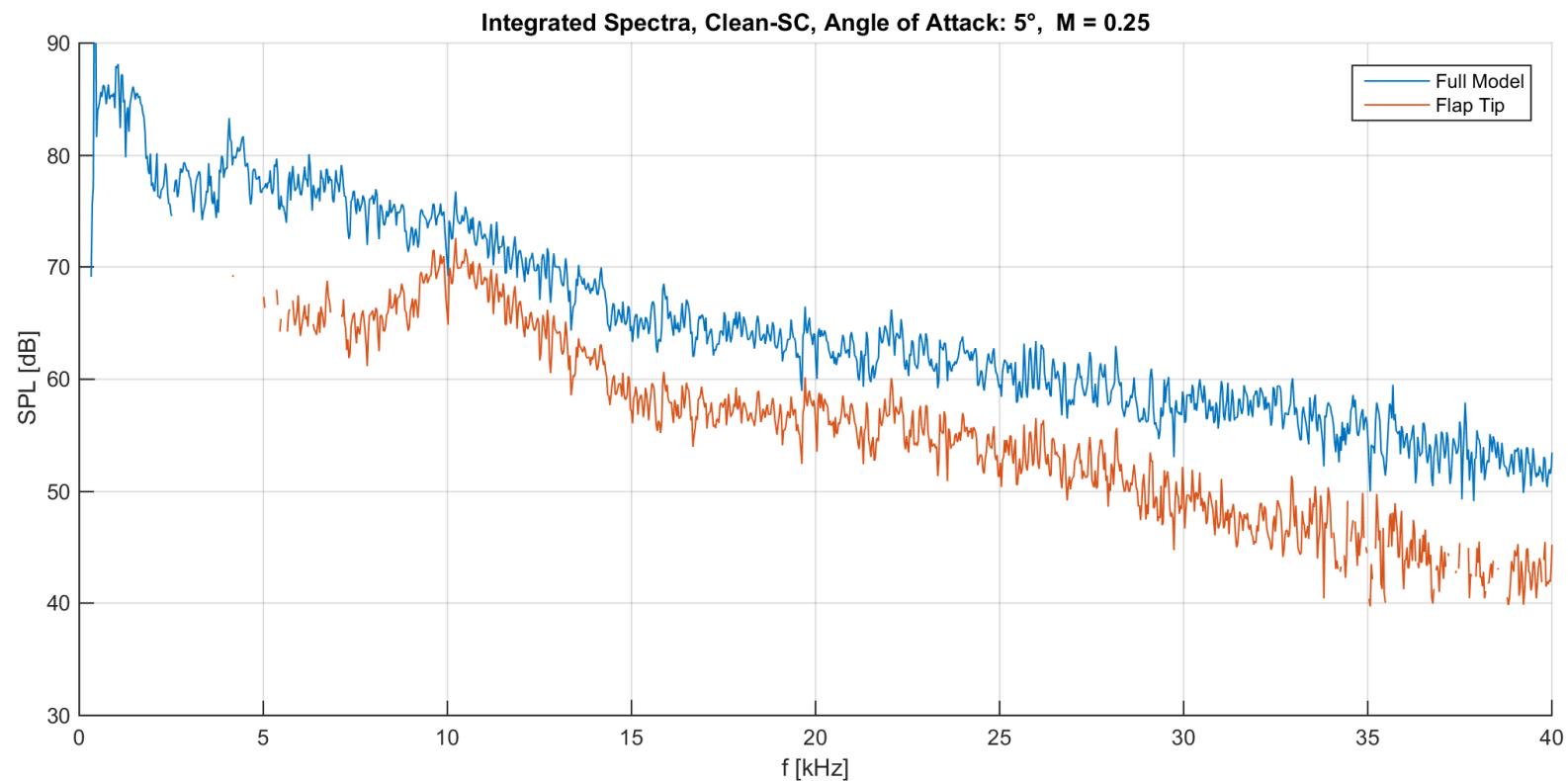


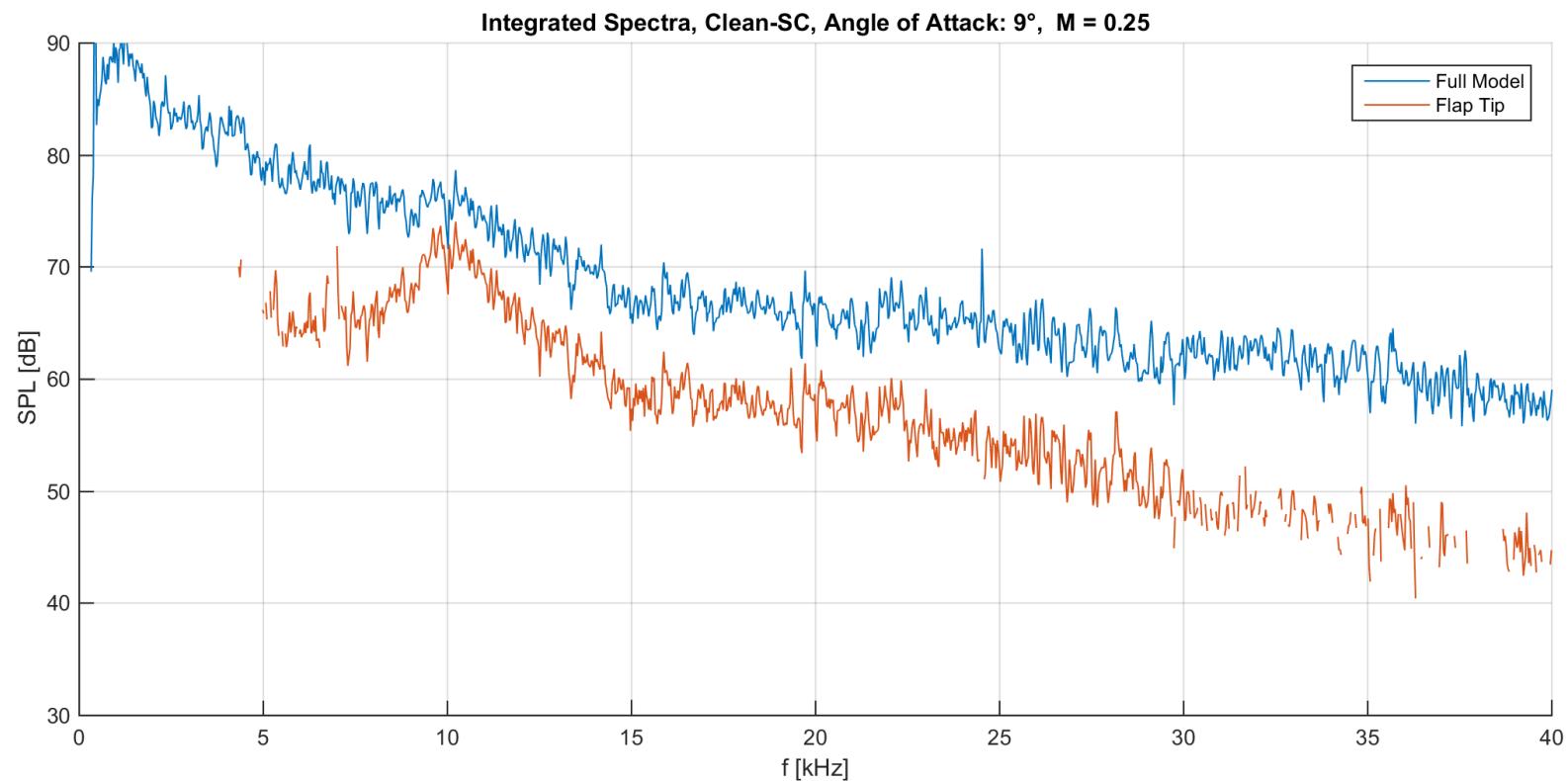


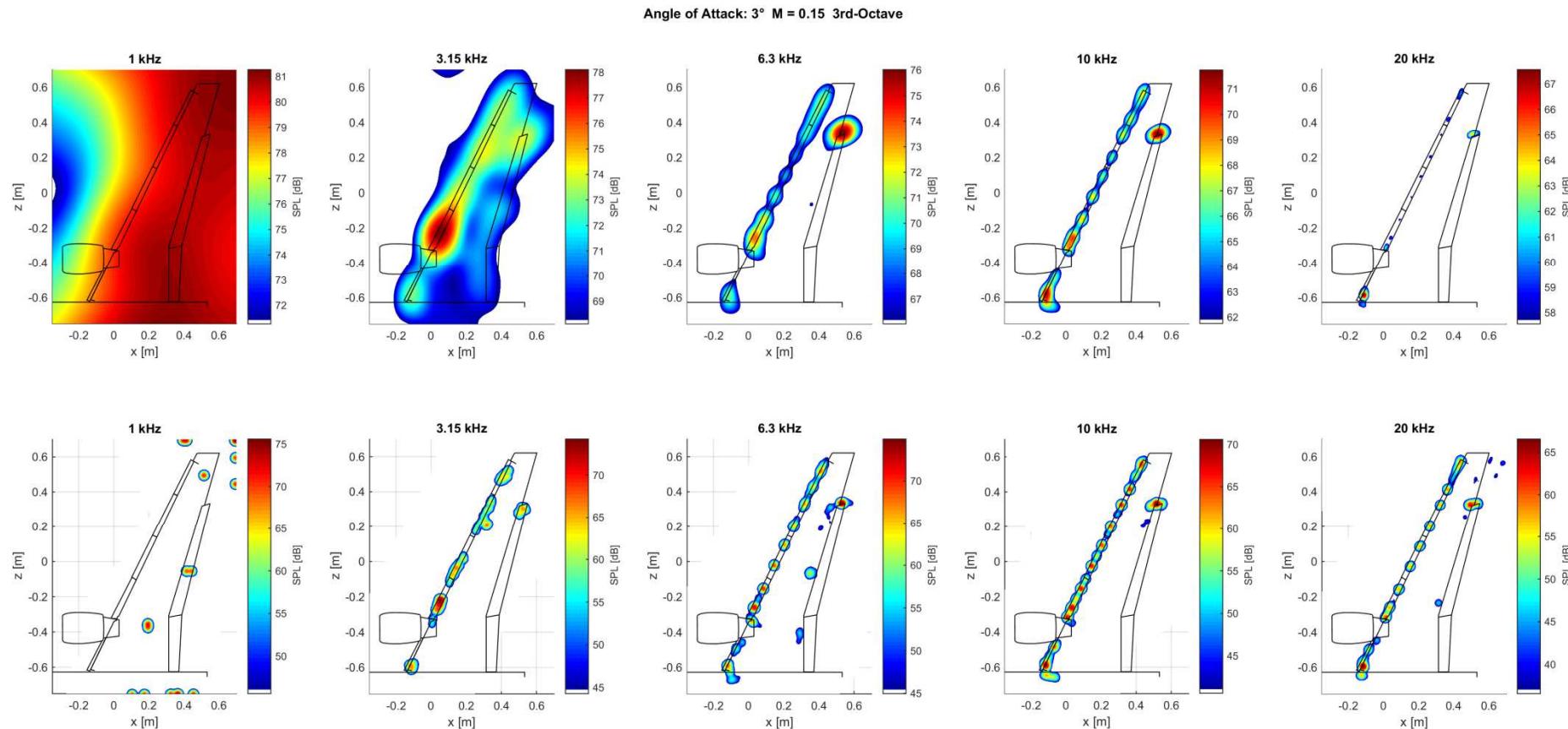










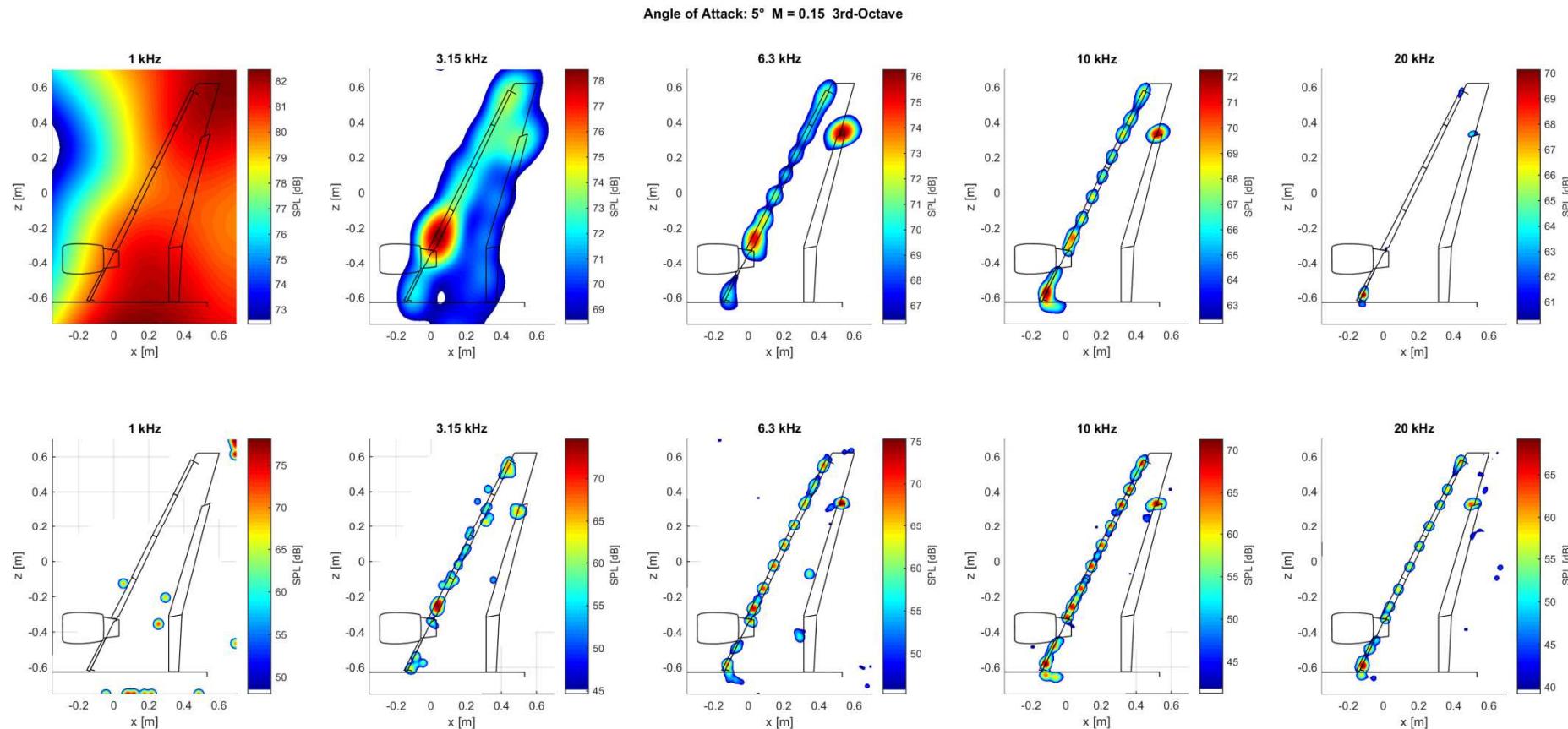


Upper Row: FDBF

Lower Row: CleanSC

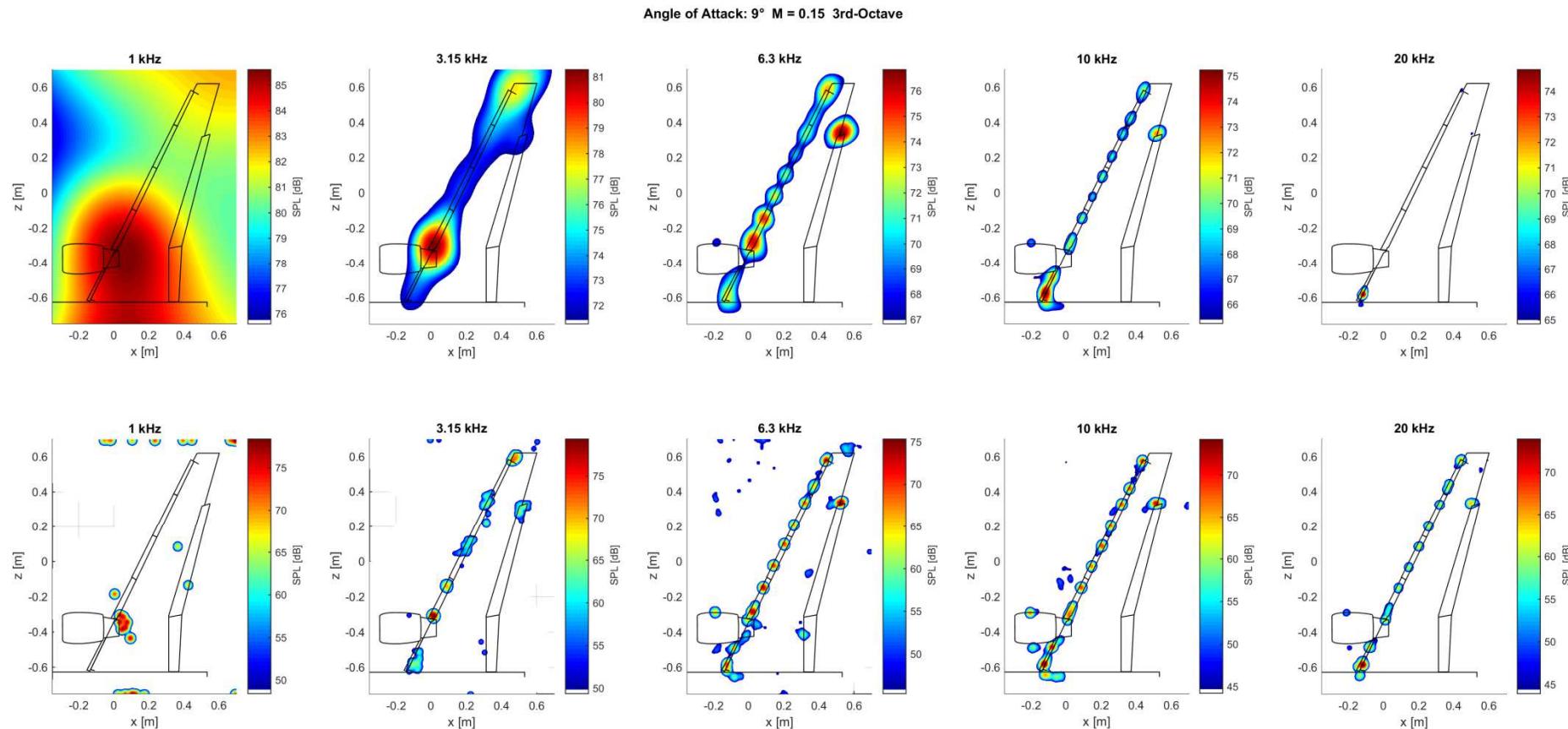
Resolution $\Delta x = 0.01m$





Resolution $\Delta x = 0.01m$



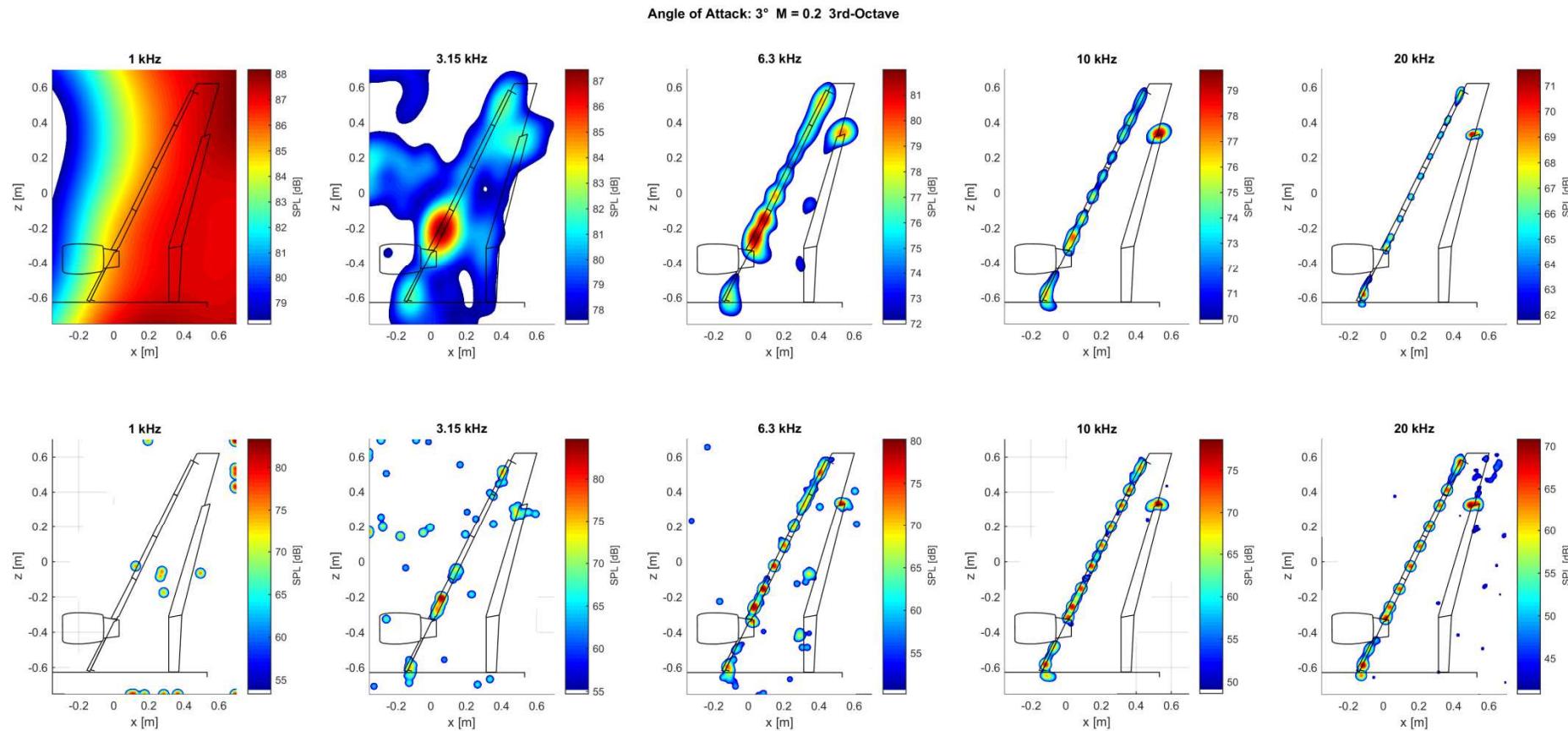


Upper Row: FDBF

Lower Row: CleanSC

Resolution $\Delta x = 0.01m$





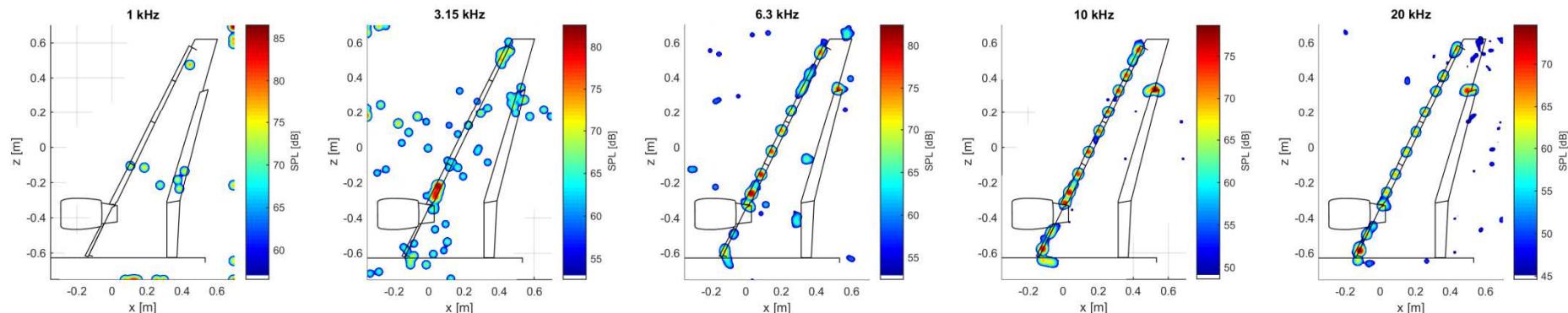
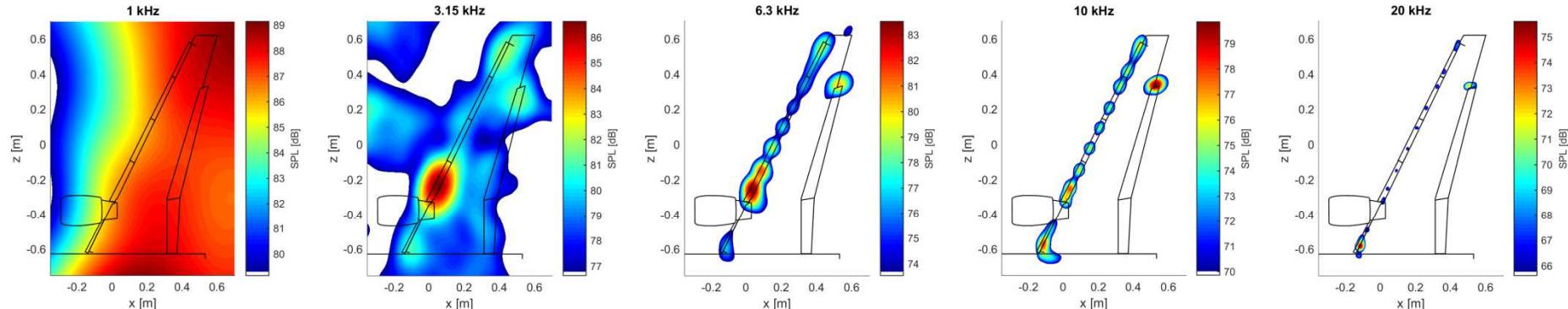
Upper Row: FDBF

Lower Row: CleanSC

Resolution $\Delta x = 0.01m$

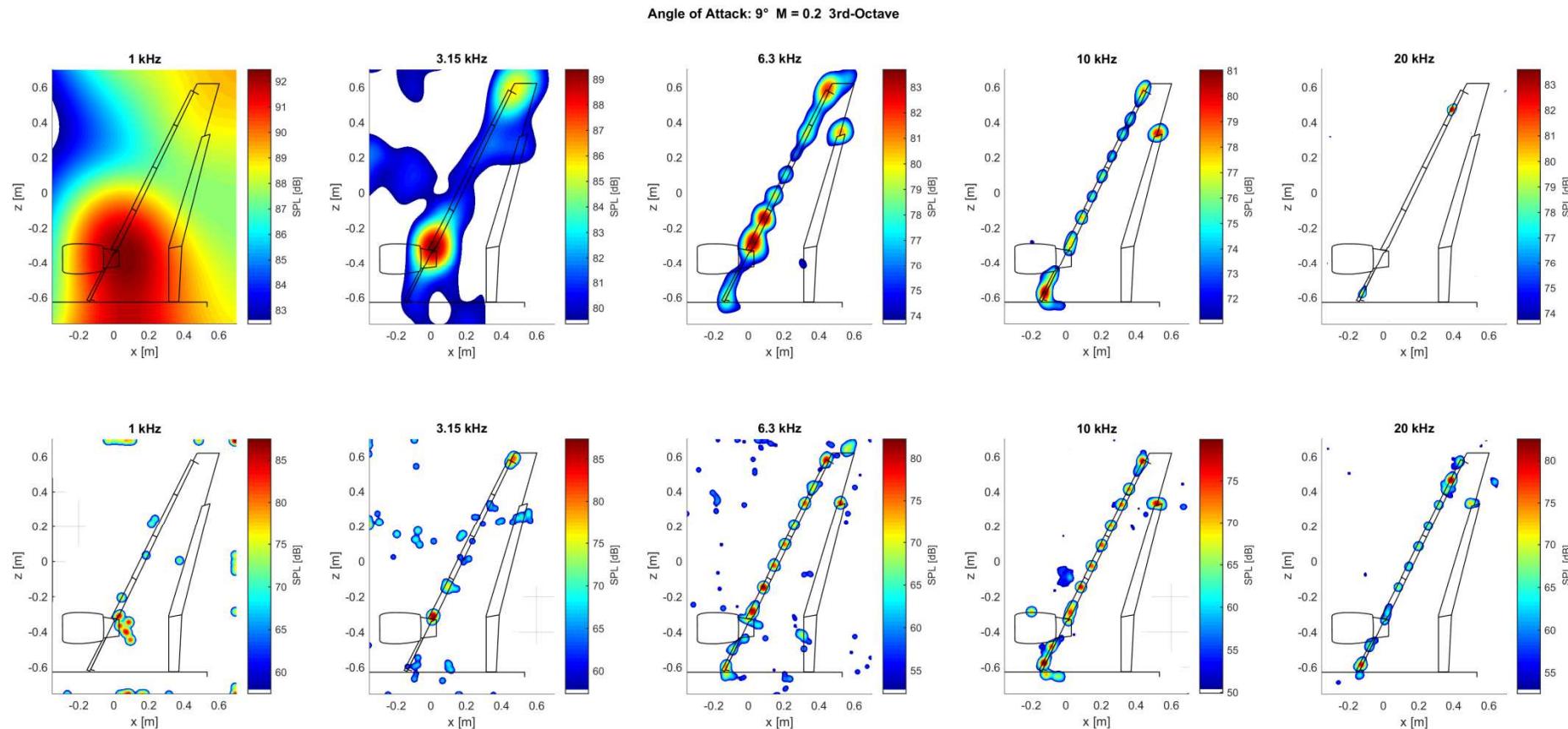


Angle of Attack: 5° M = 0.2 3rd-Octave



Resolution $\Delta x = 0.01m$



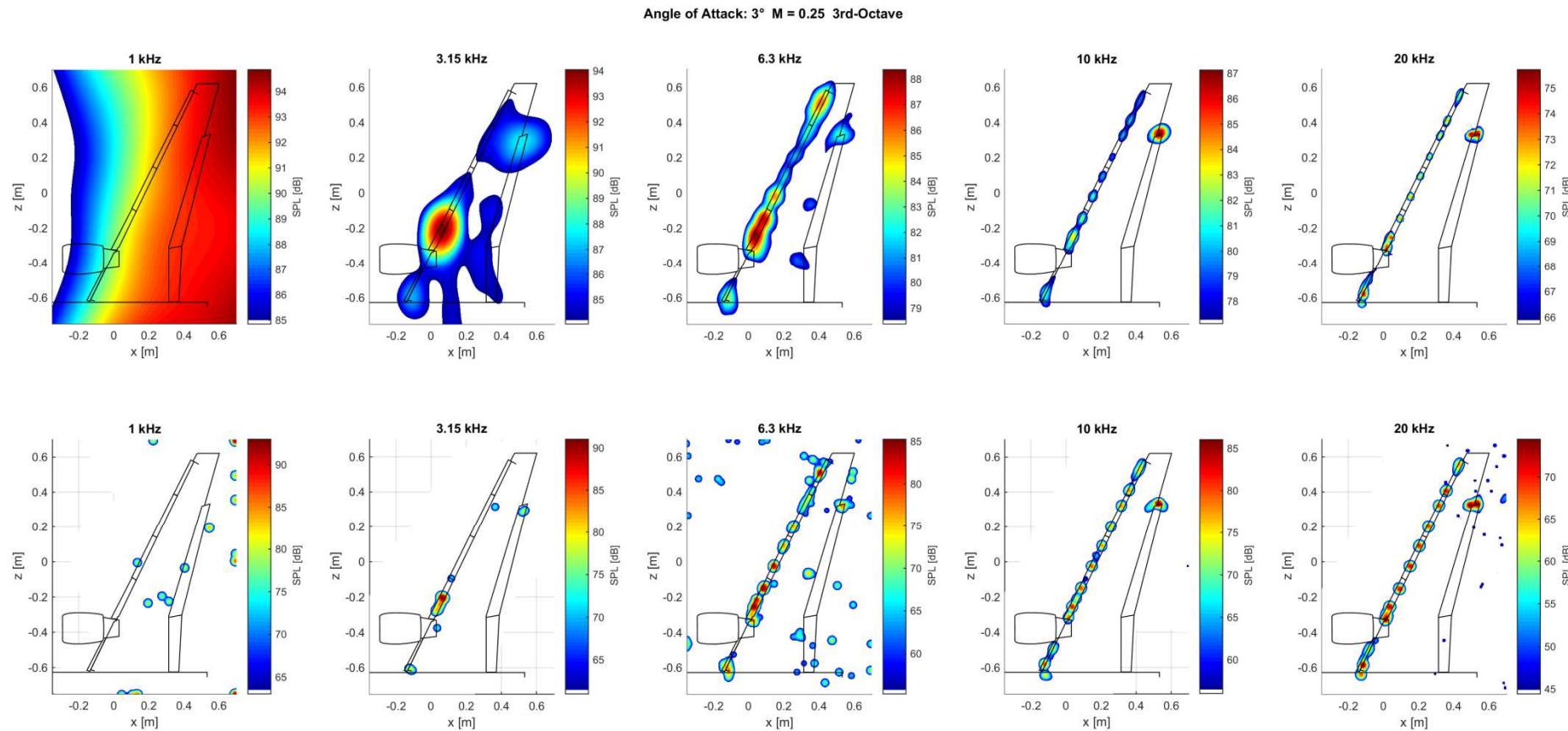


Upper Row: FDBF

Lower Row: CleanSC

Resolution $\Delta x = 0.01m$



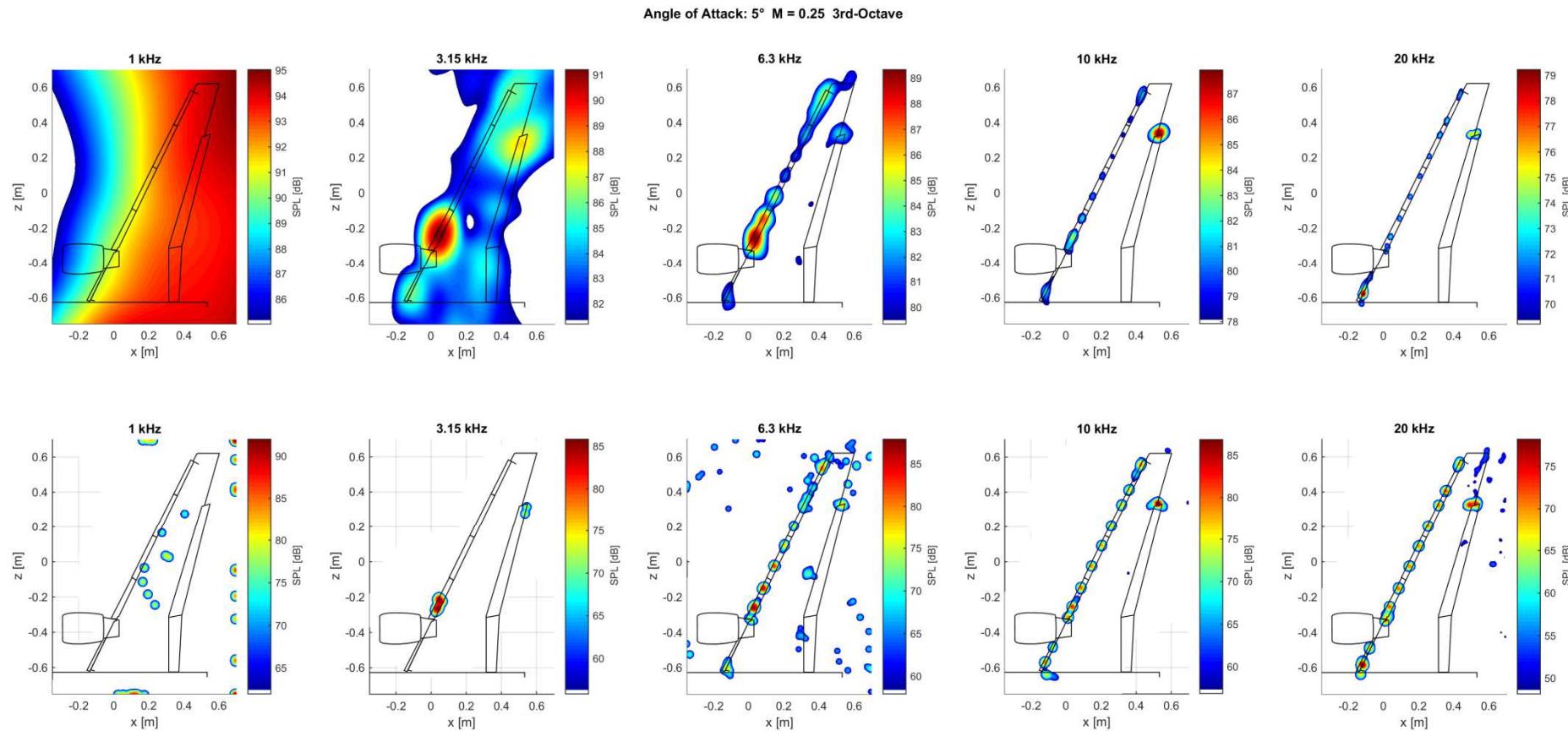


Upper Row: FDBF

Lower Row: CleanSC

Resolution $\Delta x = 0.01m$



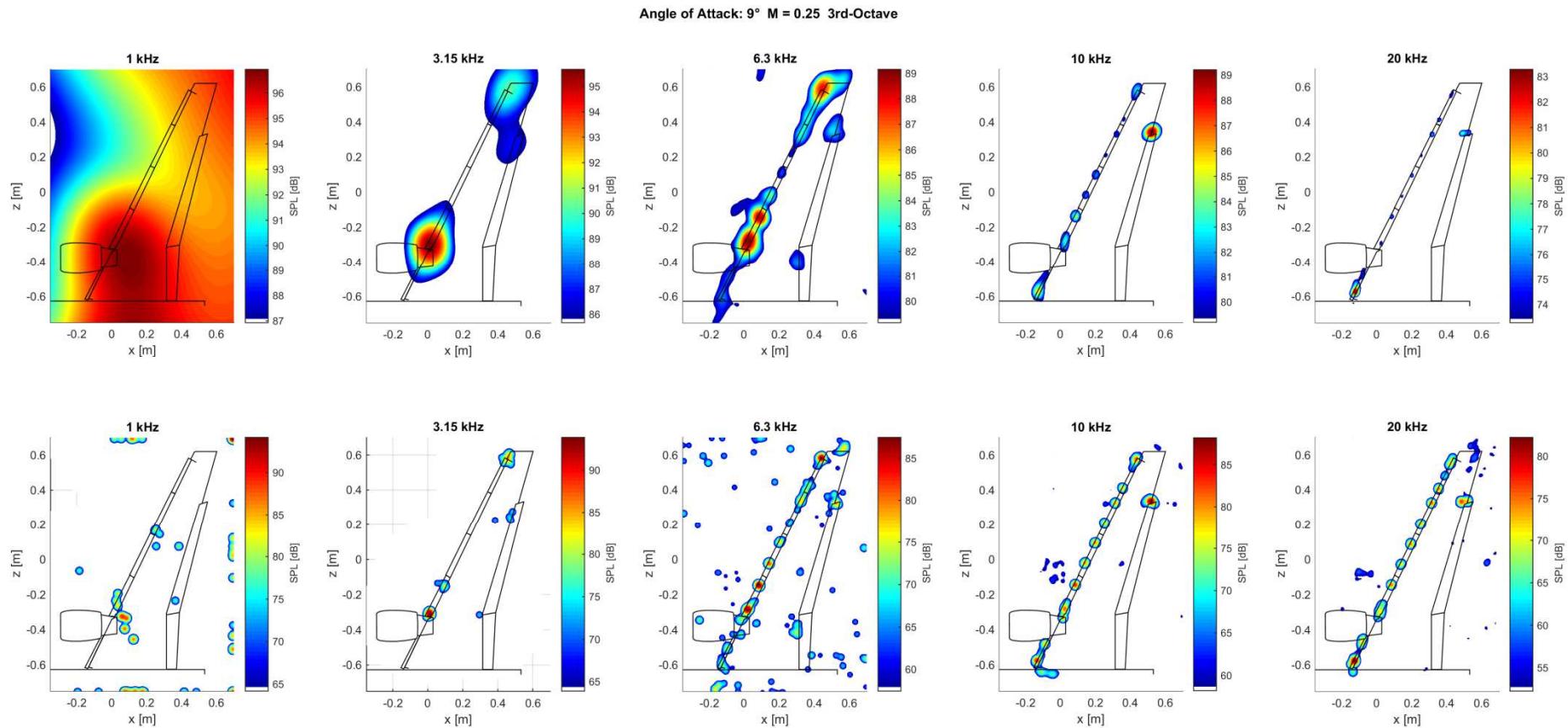


Upper Row: FDBF

Lower Row: CleanSC

Resolution $\Delta x = 0.01m$





Upper Row: FDBF

Lower Row: CleanSC

Resolution $\Delta x = 0.01m$

