



Conference Paper #09

Acoustic camera created using a distributed optical fibre –based sensor



Daniel Finfer (daniel.finfer@silixa.com), Georgios Efstathopoulos, Yousif Kamil, Sergey Shatalin, Tom Parker & Mahmoud Farhadiroushan

Berlin Beamforming Conference 2012

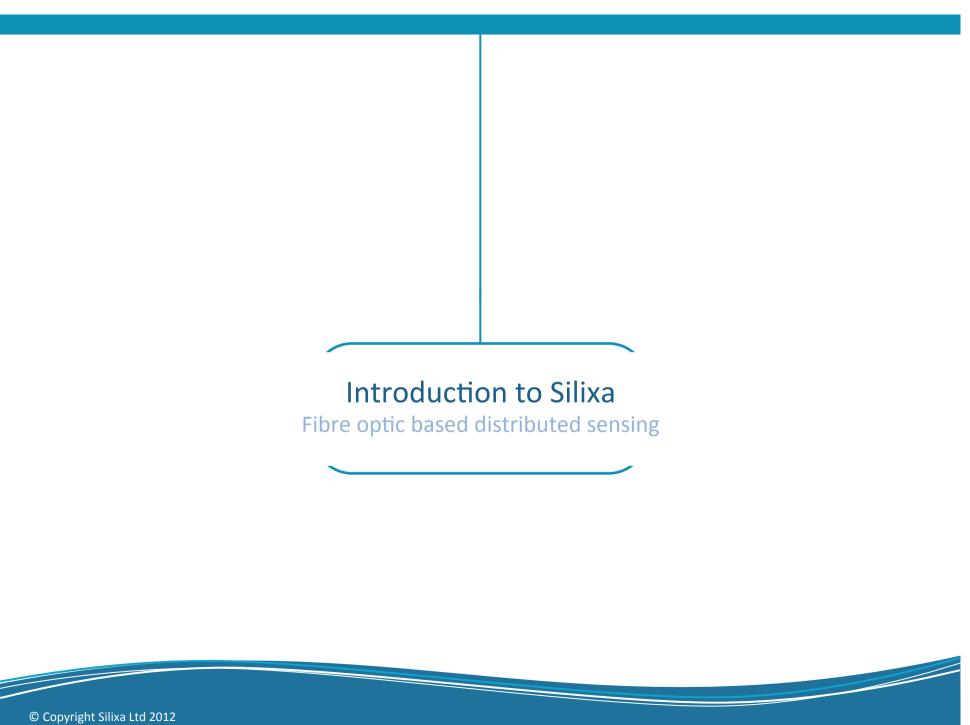
22-23 February 2012



Overview

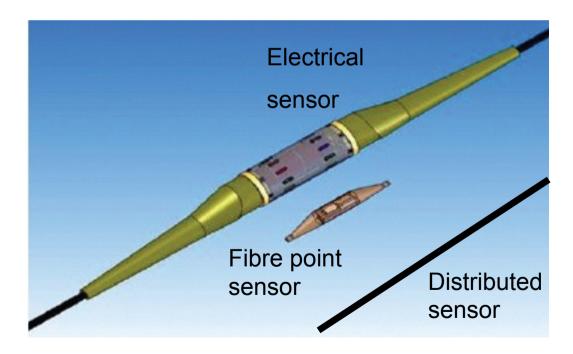


- -Introduction to Silixa and distributed sensing
- -Experiment overview
- -Signal processing approach
- -Results



Distributed & Point sensing comparison

- Continuous sensing cable
- Complete coverage
 - measurement every 1m along the entire cable length
- small size, flexible and cost-effective





intelligent Distributed Acoustic Sensor



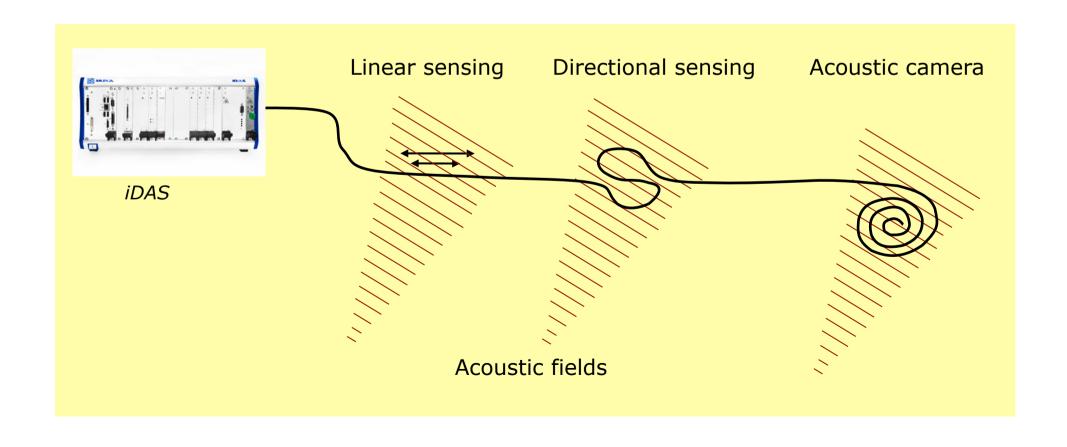


- » Simultaneous measurement of acoustic amplitude, phase and frequency at every metre along fibre
- » No cross talk
- » 90 dB dynamic range
- » Acoustic sensing array

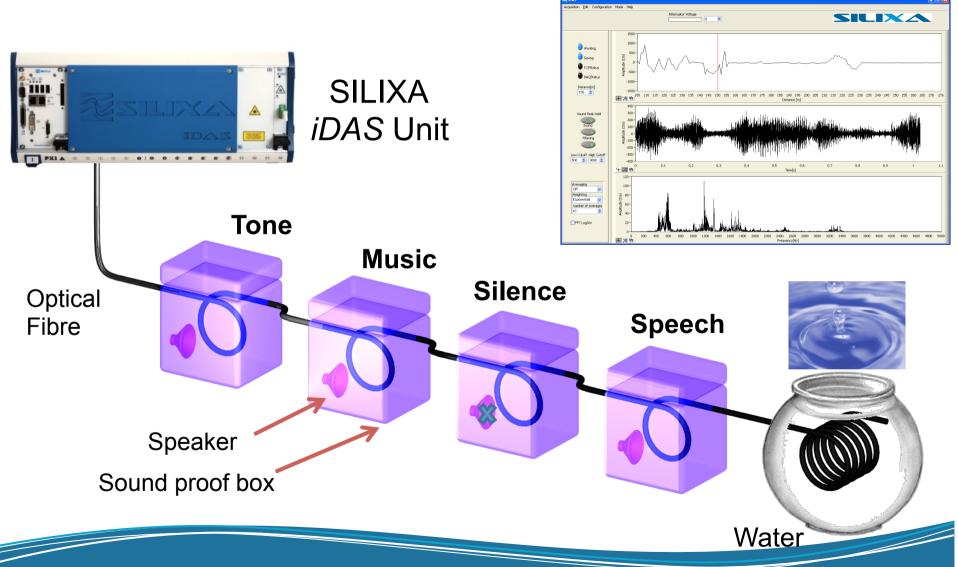
Copyright Silixa Ltd 2011

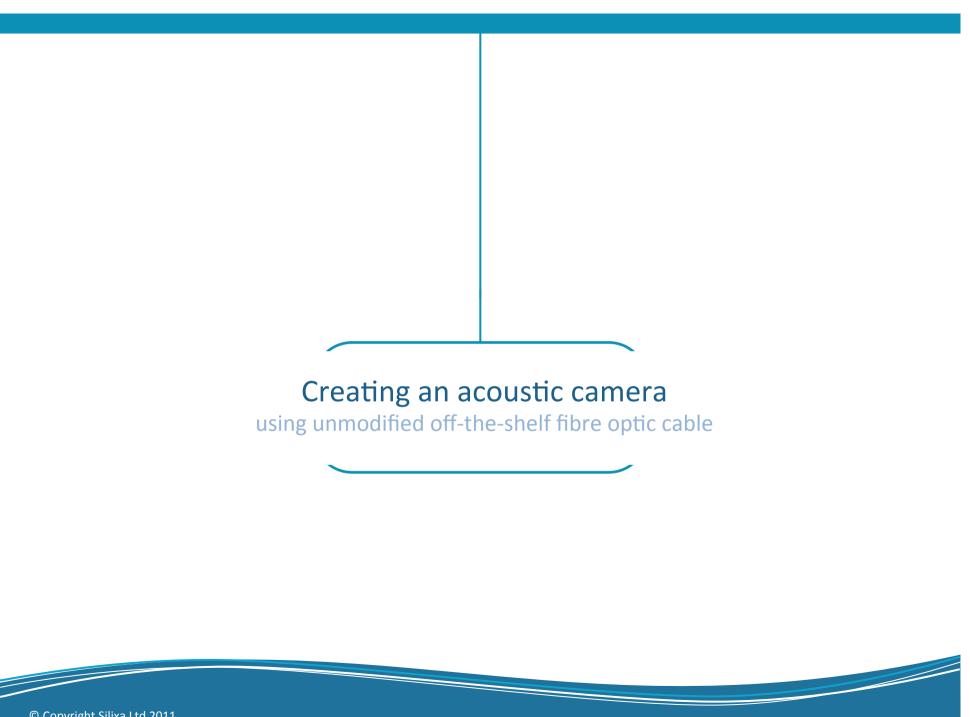
Measuring sound is just the beginning!

- Different fibre arrangements lead to a multitude of sensing opportunities
- For example, combining data from multiple points (which are all synchronised) allows beamforming to image the acoustic field far from the sensing fibre (e.g. acoustic camera/ telescope)



SILIXA *iDAS* Demonstration

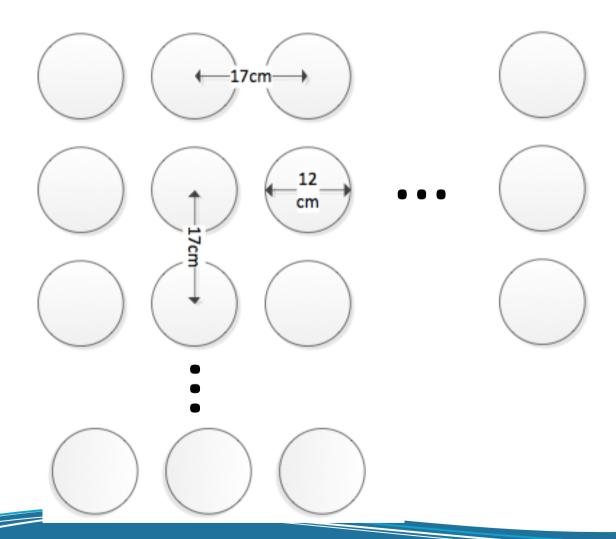






Experiment Overview



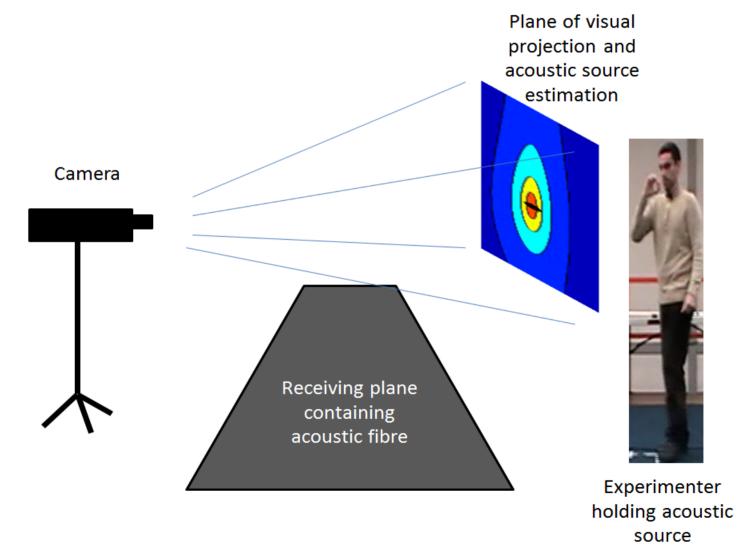




Experiment Overview

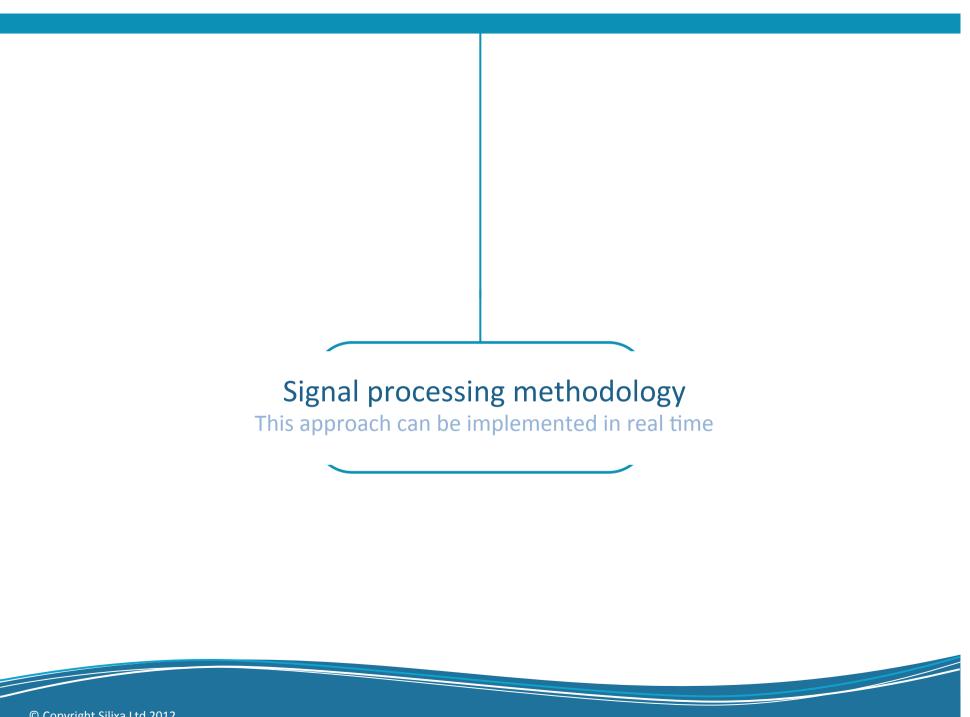














Signal Model



 Assuming that there are M narrowband sources with the same frequency for far field, the signal at the array can be modelled as:

$$\underline{x}(t) = \sum_{i=1}^{M} \underline{a}\left(\underline{\underline{r}}, \theta_{i}, \phi_{i}\right) m_{i}(t) + \underline{n}(t)$$

where $\underline{a}\left(\underline{\underline{r}},\theta_{i},\phi_{i}\right)$ is the manifold vector associated with the ith source and can be modelled as

$$\underline{a}\left(\underline{\underline{r}},\theta_{i},\phi_{i}\right) = e^{j\underline{\underline{r}}.\underline{k}(\theta_{i},\phi_{i})}$$

- The direction of arrival of the incoming signals can be estimated using different methods
- Subspace methods (e.g. MUSIC, ESPRIT) are considered as high resolution DOA estimation techniques
- If sources are moving, the direction of arrival of the signals will change with time
 and hence the manifold vector is considered to be a slow varying function of time.
 Estimates from different time intervals have to be associated together in a
 multitarget environment. These estimates have to be smoothed.



Direction of Arrival Estimation



 The MUltiple SIgnal Classification (MUSIC) algorithm is <u>chosen</u> for its high accuracy and simplicity. It is based on the formulation of the covariance matrix:

$$\mathbf{R}_{xx} = \mathbf{E}\{\underline{x}(t)\underline{x}(t)^{H}\}$$

$$\mathbf{R}_{xx} = \mathbf{R}_{sig} + \mathbf{R}_{noise}$$

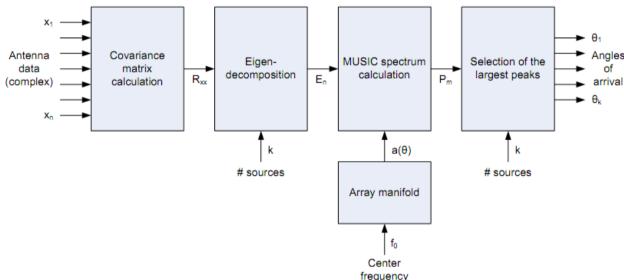
- The MUSIC algorithm utilizes the eigenstructure of the covariance matrix to estimate the direction of arrival.
- In particular, it uses the fact that the rank of the R_{sig} matrix increases with increasing number of signals. Thus, assuming that the number of sources M is known, the eigenvectors corresponding to the minimum N-M eigenvalues E_n are orthogonal to the signal manifold vectors
- The MUSIC spectrum is formed by utilizing this orthogonality and searching over the following cost function:

$$z(\theta,\phi) = \frac{1}{\underline{a}\left(\underline{\underline{r}},\theta,\phi\right)^{H} E_{n}(E_{n})^{H} \underline{a}\left(\underline{\underline{r}},\theta,\phi\right)}$$



Direction of Arrival Estimation





- The MUSIC algorithm fails when signals are perfectly correlated (multipath) the rank of the signal covariance matrix \mathbf{R}_{sig} reduces. To restore the rank of the matrix, spatial smoothing techniques are used to modify the covariance matrix.
- Spatial smoothing technique is therefore <u>used</u>. Spatial smoothing is based on preprocessing the array response by partitioning the array of sensors into subsrrays and generate the average of the subarray covariance matrices



Tracking – Particle Filtering



• A state space model is defined for the ith target consisting of an azimuth θ and elevation ϕ

$$\underline{z}_i(k+1) = \underline{z}_i(k) + \underline{\dot{z}}_i(k)T + \underline{w}(k)$$
$$\underline{z}_i(k) = [\theta_i(k) \ \phi_i(k)]^T$$

- The noise in the state space model is assumed to follow Gaussian distribution with a standard deviation of 5 degs. The movements in the azimuth and elevation directions are assumed to be uncorrelated.
- The measurement model is the output of the MUSIC algorithm which is assumed to be corrupted with a Gaussian noise as well with a standard deviation of 5 degs. $y_i(k) = \underline{z}_i(k) + \underline{e}_i(k)$
- The idea behind particle filtering is to represent the posterior density function by a set of random samples with associated weights and to compute estimates based on these samples and weights

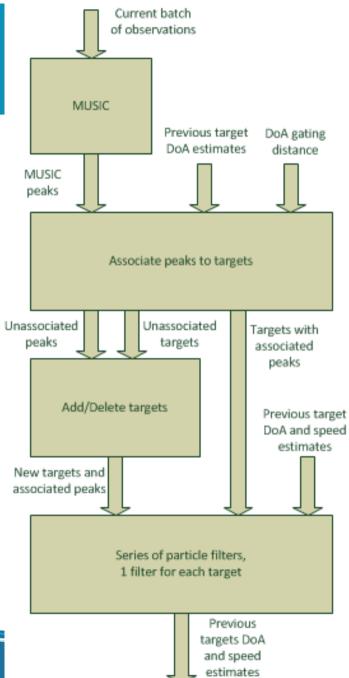
$$p\left(\underline{z}_{i}(k)\middle|\underline{y}_{i,l}(0:k)\right) = \sum_{n=1}^{N_{S}} v_{n,i}(k) \,\delta(\underline{z}_{i}(k) - \underline{z}_{i}(k)^{n})$$

for l = 1,...L MUSIC peaks for the *ith* target

· The weights are chosen to be updated using the following

$$v_{n,i}(k) \propto v_{n,i}(k-1)p\left(\{\underline{y}_{i,l}(k)\} | \underline{z}_i(k)^n\right)$$

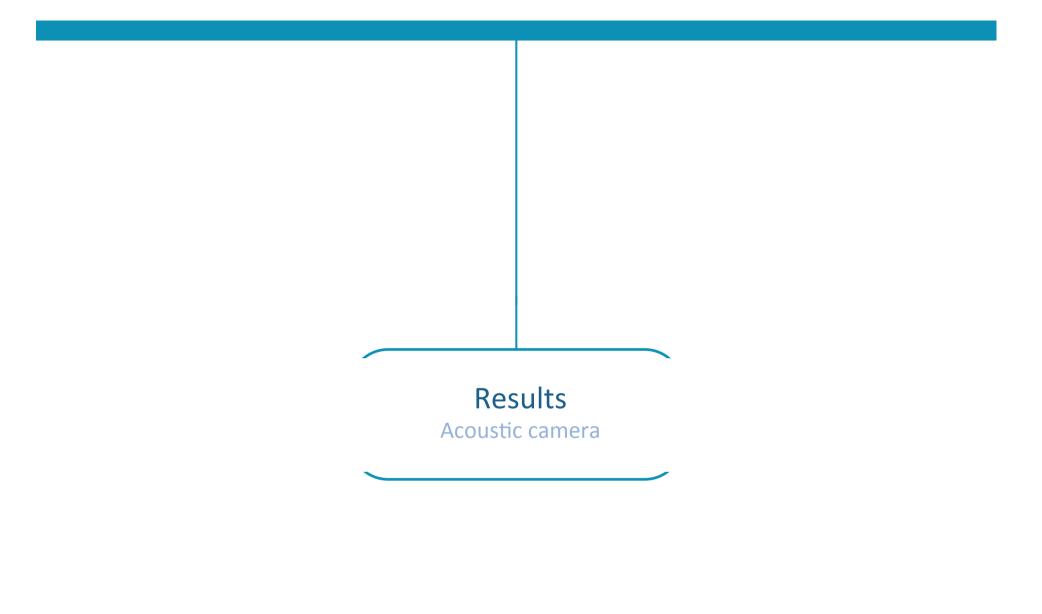








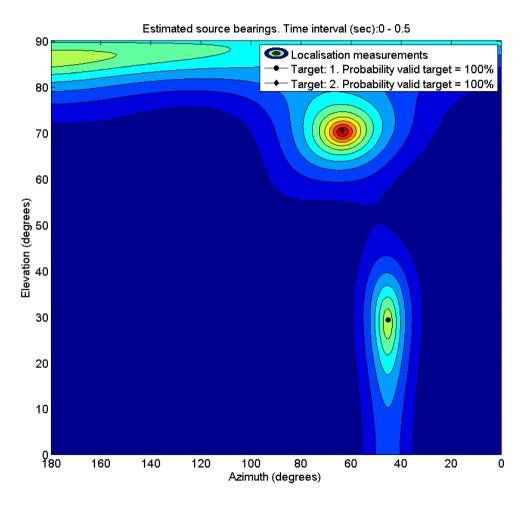
026



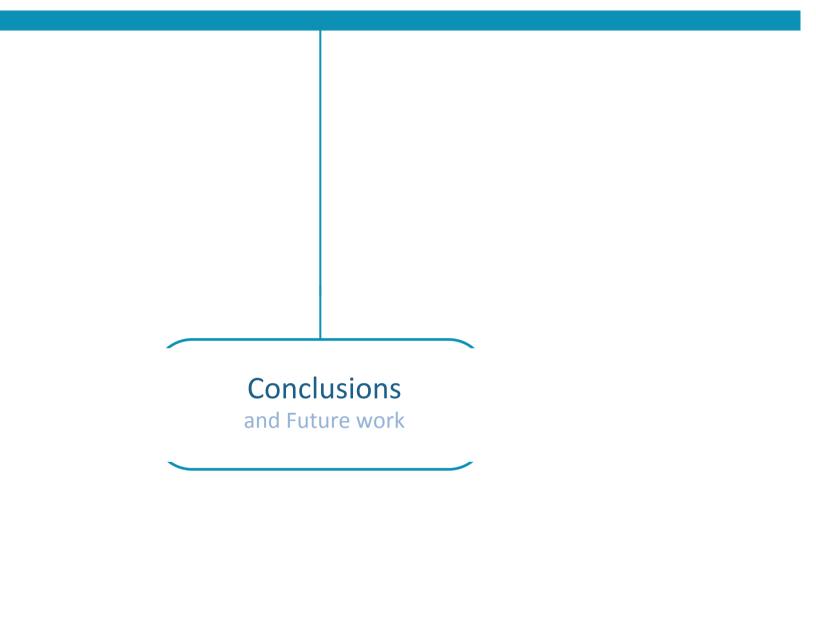


A snapshot from the video output











Conclusions



- -iDAS can be used to develop a far field acoustic camera
- -This result requires some predictive filtering to reduce multi-path effects
- -This result can be implemented in real-time

Next step

-Design and implementation of a massive acoustic lens



End



Thank You



T: +44 (0) 208 327 4210 F: +44 (0) 208 953 4362

W: www.silixa.com

E: enquiries@silixa.com